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Design of a Microprocessor-Based Control
System for the Monterey Institute for
Research in Astronomy 36" Telescope

by

David Paul Wood

June 1992

Thesis Advisor:
Co-Advisor

W. Bruce Weaver
I. M. Ross

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Design of a Microprocessor-Based Control System for the Monterey
Institute for Research in Astronomy 36" Telescope

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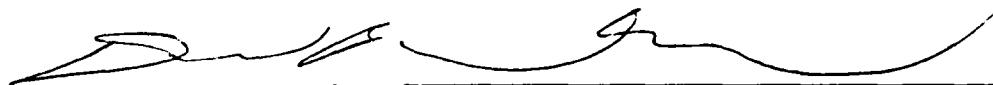
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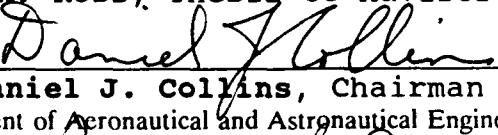


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ABSTRACT

The use of computer technology to control large pointing systems can significantly improve performance and reduce human work load. The goal of this thesis was to design software for an inexpensive, yet accurate and efficient control system for the 36-inch reflecting telescope owned and operated by the Monterey Institute for Research in Astronomy. Within this thesis, a computer program is developed to automatically move the telescope to a set of celestial coordinates and track with an accuracy of one-tenth of an arc second for five minutes within 75° of the zenith. Set times are anticipated to be between four and thirty seconds. Corrections are made to celestial coordinates to account for precession, nutation, aberration and atmospheric refraction effects. The user is provided an interface to the computer-based system that allows storage and editing of 100 star positions, editing of the system parameters and display of the telescope's status. Manual control of the telescope is also permitted at any time. Safety of the telescope structure is the primary concern of system software.

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TABLE OF CONTENTS

I.	INTRODUCTION	1
A.	STATEMENT OF PROBLEM	1
B.	THESIS SCOPE	2
C.	BACKGROUND	4
1.	THE MONTEREY INSTITUTE FOR RESEARCH IN ASTRONOMY	4
2.	CURRENT HARDWARE CONFIGURATION	5
3.	CURRENT PROCEDURES	7
II.	COORDINATE SYSTEMS	9
A.	CELESTIAL COORDINATES	9
B.	CONVERSION OF CELESTIAL TO TERRESTRIAL COORDINATES	11
III.	TIME	17
A.	UNIVERSAL COORDINATED TIME	17
B.	TIME SYSTEM CONVERSION	17
C.	TIME SYNCHRONIZATION	19
D.	CORRECTING COORDINATES FOR TIME	20
IV.	CORRECTIONS TO CELESTIAL COORDINATES	23
A.	PRECESSION	25

B.	NUTATION	29
C.	ABERRATION	30
D.	REFRACTION	32
V.	HARDWARE	37
A.	SYSTEM OVERVIEW	37
B.	COMPUTER	38
VI.	PROGRAM DEVELOPMENT	41
A.	MAIN PROGRAM DESCRIPTION	42
B.	SUBROUTINE DESCRIPTIONS	47
1.	Display Subroutines	47
a.	InitialDisplay	47
b.	StarDisplay	49
c.	Windows	49
d.	WindowsPop	49
e.	TeleStatusDisplay	49
f.	ShowTime	50
g.	TeleAngles	50
2.	Data Entry Subroutines	51
a.	TeleAlign	51
b.	Atmospheric	51
c.	StarDataEntry	52
d.	ChooseStar	52
3.	Data Correction Subroutines	54
a.	Precession	54

b.	Nutation	56
c.	Aberration	56
d.	Refraction	57
4.	Time Correction Subroutine	57
a.	SiderealTimeCalc	57
5.	Key and Video Handling Subroutines	58
a.	KeyCode	58
b.	KeyCodeNoWait	58
c.	VideoState	58
6.	Telescope Movement	59
a.	SlewCommands	59
b.	SetCommands	60
c.	TrackCommands	60
C.	ERROR AND WARNING MESSAGES	61
1.	The Quit Confirmation Window	61
2.	The File Save Error Window	61
3.	The File-Not-Found Error Window	62
4.	The Equinox Correction Error Window	62
5.	The Slew Confirmation Window	62
6.	The Change Telescope Status Window	63
7.	The RA/DEC Format Error Window	63
8.	The Slew Error Window	63
9.	The Zenith Angle Warning Window	64
10.	The Zenith Angle Halt Movement Window . .	64
11.	The Telescope Status Help Window	64
12.	The CTS-10 Failure Window	65

VII. PROCEDURES	67
VIII. PROGRAM VALIDATION	71
A. TRACKING VALIDATION	71
B. CELESTIAL COORDINATE CORRECTION VALIDATIONS . .	72
C. MOTOR COMMAND VALIDATION	73
IX. CONCLUSIONS AND RECOMMENDATIONS	75
A. CONCLUSIONS	75
B. RECOMMENDATIONS FOR ADDITIONAL WORK	75
APPENDIX A MIRACTRL.EXE	79
A. MIRACTRL.BAS LISTING	79
B. SAMPLE DATA FILE "STARFILE.DAT"	158
LIST OF REFERENCES	159
INITIAL DISTRIBUTION LIST	160

LIST OF FIGURES

Figure 1 - The MIRA 36" reflecting telescope with equatorial mount	6
Figure 2 - Existing hardware prior to this thesis . . .	7
Figure 3 - The celestial sphere	10
Figure 4 - Right Ascension and Declination	11
Figure 5 - Hour Angle and Declination	12
Figure 6 - Definition of the A and B Frames	13
Figure 7 - Azimuth (Φ), Elevation (Θ) and Zenith Angle ($z = 90^\circ - \Theta$) related to rectangular coordinates	14
Figure 8 - Right Ascension (α) and Declination (δ) related to celestial rectangular coordinates	15
Figure 9 - Effects of nutation on coordinate systems .	26
Figure 10 - Causes of nutation	27
Figure 11 - Refraction through a medium	33
Figure 12 - Proposed hardware configuration	37
Figure 13 - Program modular concept	43
Figure 14 - Simplified flow chart of the main program .	44
Figure 15 - Simplified flow chart of the main loop . .	48
Figure 16 - The initial display screen with user box .	50
Figure 17 - StarDataEntry screen	53
Figure 18 - ChooseStar screen	55
Figure 19 - The telescope status display screen with example data	71
Figure 20 - Eventual hardware configuration	76

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I. INTRODUCTION

Digital computing equipment, in the form of personal computers, can provide a powerful tool to control large physical systems. This thesis explored the feasibility of upgrading the control system of one such system, an astronomical telescope owned by the Monterey Institute for Research in Astronomy (MIRA). A computer program was written to control the telescope's movements for star tracking to several arc seconds degree of accuracy. A tracking rate error of less than a tenth of an arc second degree of accuracy for five minutes within 75° of the zenith was required. Additionally, requirements to interface a personal computer to the existing hardware were determined. Priority was given throughout the design process to the safety of the telescope structure.

A. STATEMENT OF PROBLEM

The MIRA telescope was to be given the capability to meet the following design criteria:

- Given a known star position, slew on command at the maximum acceptable slew rate to within one degree of the star's calculated position.
- Once within one degree of the calculated position, set via the right ascension and declination stepping motors to within one step of the calculated position.

- The star's calculated position is required to be within ten arc seconds of the star's actual apparent position.
- Manual adjustments to eliminate error between actual apparent and calculated positions are to be made via existing hand paddles or the computer keyboard.
- Calculated corrections to a star's tabulated position are to include precession, nutation, aberration and refraction of light through the Earth's atmosphere.
- Once error has been sufficiently reduced, track the star's apparent position by compensating for rotation of the Earth and estimated refraction of the atmosphere.
- A user interface is required, using standard astronomical vocabulary and symbols. This interface shall include the capability to read, write and edit a listing of one hundred star positions from a formatted data file.
- The telescope's status is to be displayed at all times during telescope movement. The status display is to provide the following information: The telescope's position in celestial and terrestrial coordinates, local, sidereal and universal times and the state of each motor.
- The user interface shall also display adequate error and warning messages to inform the user of special circumstances.
- Allow guiding of the telescope manually by hand paddles or keyboard at any time to eliminate or reduce subsequent errors.

Use of a digital computer was chosen as the simplest design solution. Motor controller hardware, telescope position data and user choices were to be interfaced to this computer via the computer program developed by this thesis.

B. THESIS SCOPE

Requirements to upgrade existing MIRA hardware to encompass an automatic control system for the telescope fell into two categories; hardware and software. This thesis

determined the hardware required to interface the existing motor sets and position determination devices with a personal computer and developed the software which that computer would use to control motor movement. The tracking accuracy requirement of one-tenth of an arc second required that the relation of the Earth's movement to any target star be precisely determined and accounted for. Refraction of light through a variable atmosphere and aberration effects due to the Earth's motion were also considered.

The concept of a virtual and a physical telescope is used throughout this thesis. The virtual telescope is an abstraction in software upon which operations are performed without effecting the movement of physical objects. The virtual telescope represents ideal viewing conditions, a stable Earth with no atmosphere and no Sun to perturb the path of incoming light. Coordinates for a star, corrected only for precession, define the desired position of the virtual telescope for a given time. This position is easily calculated by an astronomer and is used to interface with the computer system. In reality, however, many sources of error other than precession must be considered, such as nutation, aberration and refraction of light through the Earth's atmosphere. The correction of a star's coordinates for all of these errors defines the required position of the physical telescope. An astronomer, in effect, positions the virtual telescope based on incomplete, yet readily obtainable data.

The control system, guided by the computer, corrects the pointing instructions and positions the physical telescope. In effect, the virtual telescope is defined as pointing to a star's correct geometrical coordinates. The physical telescope must point elsewhere in order to see an object at those coordinates.

Corrections to the virtual telescope include precession, nutation, aberration, and atmospheric refraction effects. Structural considerations such as telescope flexure, fork twist, droop, and encoder eccentricity also cause errors which were not considered within this thesis. These corrections require careful measurements of the physical telescope and must be made prior to system completion.

To prevent distortion, the telescope's main mirror is only rear and side defined; it is not front constrained in its housing. Were the telescope to become inverted, this mirror would lose alignment and even possibly be at risk. Safety of this mirror was the overriding design concern.

C. BACKGROUND

1. THE MONTEREY INSTITUTE FOR RESEARCH IN ASTRONOMY

The Monterey Institute for Research in Astronomy is the only privately-owned professional astronomical observatory in the United States to be founded in this century. MIRA is a non-profit corporation dedicated to research and education in astronomy. It maintains a close association with two

institutions of higher learning in Monterey, the Monterey Peninsula College and the U.S. Naval Postgraduate School (NPS). A memorandum of understanding exists between MIRA and NPS allowing the use of government-owned equipment on MIRA property and the teaching of NPS students by MIRA astronomers.

MIRA owns and operates a 36-inch reflecting telescope located on Chews Ridge in the Los Padres National Forest. The Oliver Observing Station is located at approximately 5000 feet altitude above sea level on land owned by the U.S. Forest Service. Additionally, facilities in Monterey include administrative offices and an engineering and receiving building with computer, electronics and machining facilities.

2. CURRENT HARDWARE CONFIGURATION

The telescope at Chews Ridge is a 36-inch reflecting telescope. A 36-inch f/4 paraboloid mirror reflects incident light to a second, smaller cassegrain hyperboloid mirror mounted axially in front of it (see Figure 1). This second mirror directs light to the Guide and Acquisition Package (GAP). The telescope has a back focus of fourteen inches. The GAP is mounted axially, with a zero to six-inch back focus and may contain many instruments, including a spectrometer, a CCD camera, photographic equipment, and an optical eyepiece. The entire telescope is mounted equatorially, which means that it rotates about an axis parallel to the Earth's axis.

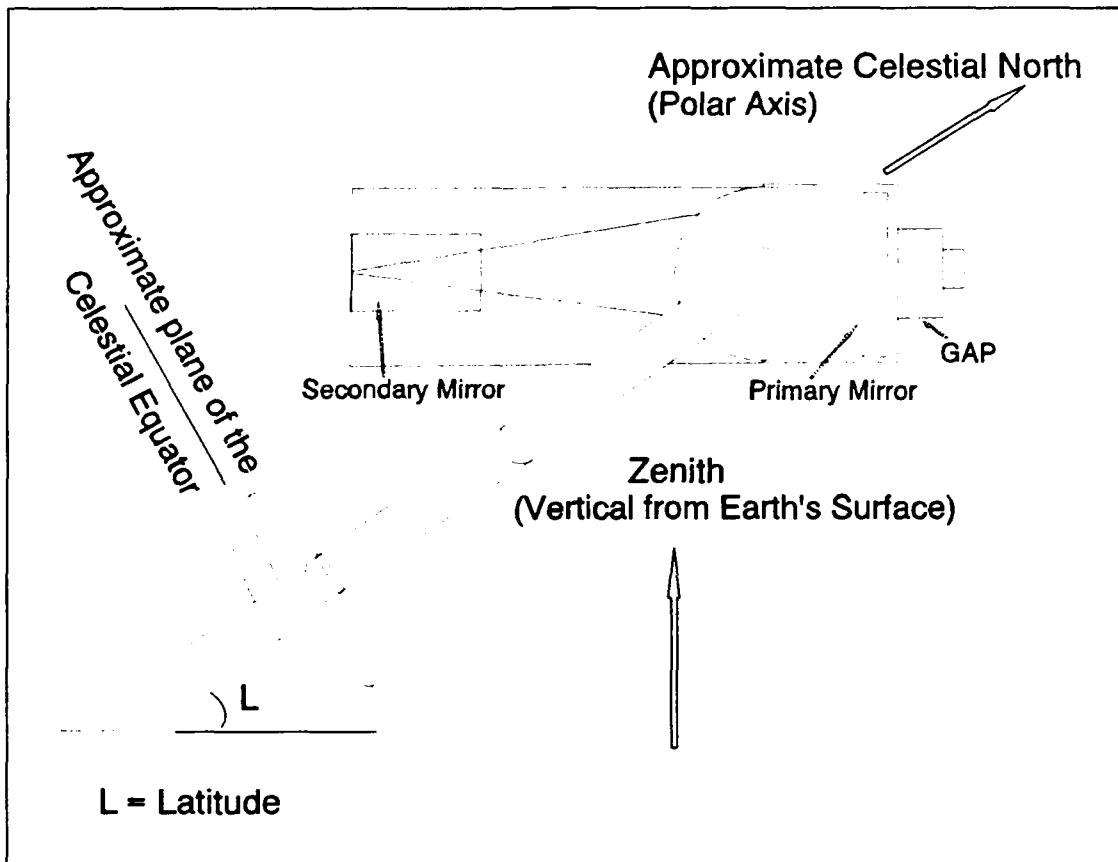


Figure 1 - The MIRA 36" reflecting telescope with equatorial mount

Rotation about the polar axis is then equivalent to changing the right ascension of the telescope (see Chapter II). Rotation about an axis orthogonal to the polar axis is also possible, which changes the telescope's declination, or celestial latitude. Two sets of motors are used to control telescope movement. Two one-quarter horse power slew motors, one each for right ascension and declination, are used for large telescope movements. For finer movements, such as tracking a star, two additional stepping motors are used.

An optical encoder on each of the two axes detect movements of the telescope. This information, once referenced to a known location, is decoded by existing electronics and sent via an RS-232 link to provide the telescope's current (uncorrected) celestial coordinates in the form of declination and local hour angle. Right ascension must be calculated from the local hour angle by knowing the current sidereal time.

3. CURRENT PROCEDURES

Prior to this project, the four motors were controlled electro-mechanically by means of switches mounted on hand paddles. Tracking the star, that is, adjusting the right

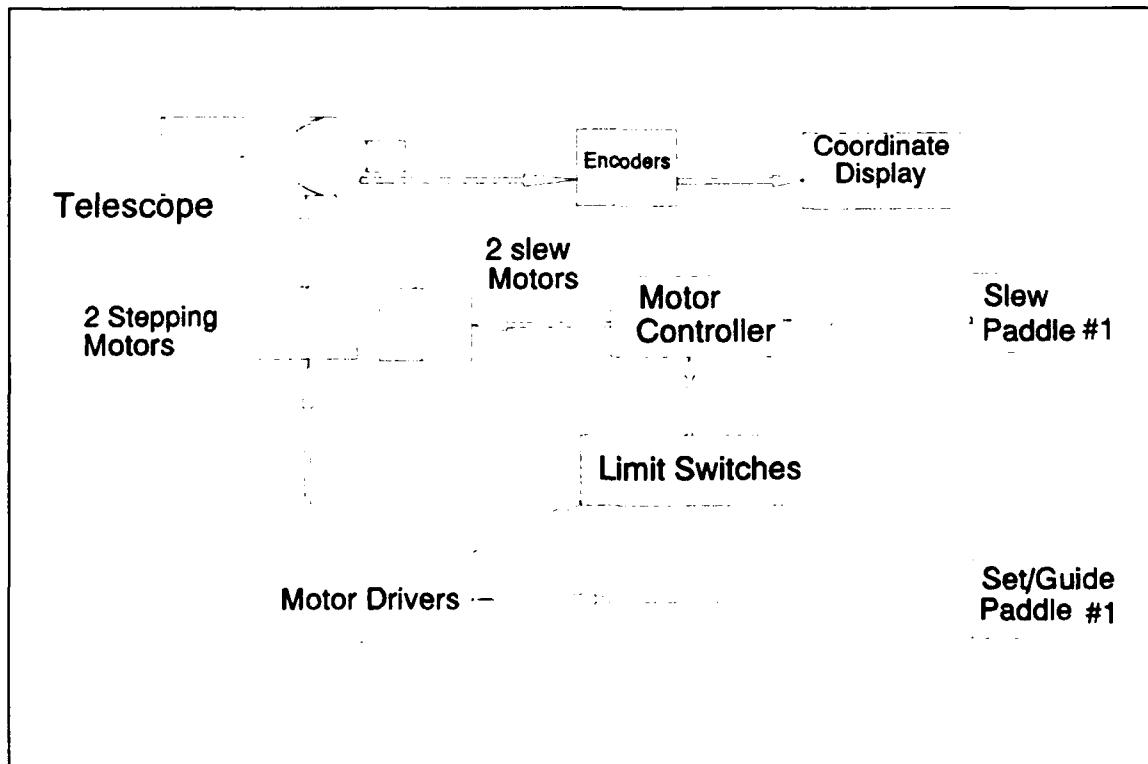


Figure 2 - Existing hardware prior to this thesis

ascension of the telescope at a rate approximately equal to the rotation rate of the Earth, is automatic. Guiding the telescope entails making adjustments for effects not associated solely with the Earth's rotation, such as atmospheric refraction. In this case, the virtual and physical telescopes are identical. An astronomer wishing to observe a star would keep the telescope properly pointed by continual minute adjustments to the stepping motors. This process was long, tedious and tiring which lead directly to the need for an automatic system. Upon completion of the required hardware, astronomers of the future shall be able to sit in a warm-room environment while concentrating on an experiment's data instead of telescope operation.

II. COORDINATE SYSTEMS

A. CELESTIAL COORDINATES

The celestial coordinate system used is centered on the Earth and extends outward to the celestial sphere. The celestial sphere is a hypothetical spherical surface upon which all stars are said to reside. Celestial North is aligned with the Earth's North Pole. Declination corresponds to latitude on the celestial sphere and right ascension is analogous to terrestrial longitude. Where terrestrial longitude is referenced to the prime meridian at Greenwich, England, right ascension is referenced from the prime celestial meridian. The prime celestial meridian passes through the vernal equinox, sometimes called the First Point of Aries. Figures 3, 4 and 5 show these relations.

Right ascension is defined as the angle, measured eastward along the celestial equator, between the vernal equinox and the great circle passing through the poles and the body. Declination is the angle, measured north- or southward along the meridian, from the celestial equator to the body. North declination is taken as positive, south as negative.

An hour circle is a great circle passing through the celestial poles. An hour circle passing through an observer's zenith is also called a celestial meridian. Hour angles are

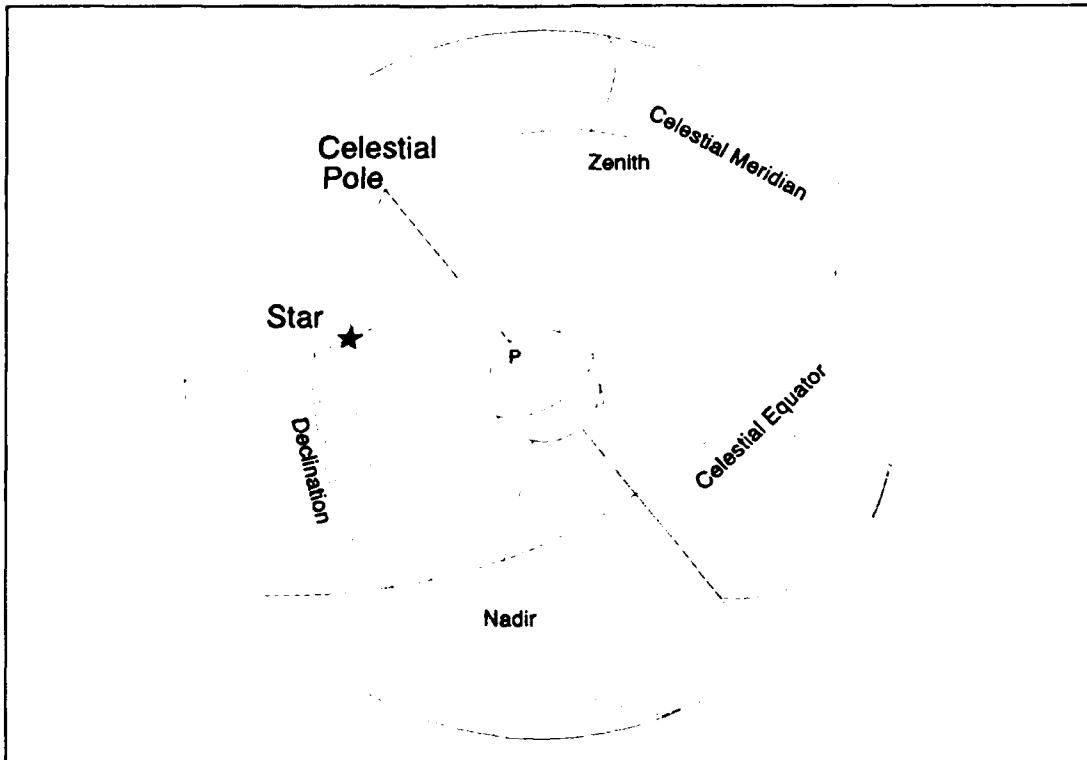


Figure 3 - The celestial sphere

angles measured east- or westward from a celestial meridian to a body. Local hour angles are referenced from an observer's meridian to a body, with westward angles being positive and eastward angles being negative.

From any point on Earth, the point directly overhead is known as the zenith. Its opposite point, directly through the Earth is the nadir. An angle from the horizon to a body is known as the elevation. The zenith angle of a body is simply the angle from the zenith to the position of the body on the celestial sphere. It may be recognized as the co-elevation, or 90 degrees minus the elevation angle. The azimuth of a body is the angle from local North, measured eastward in a

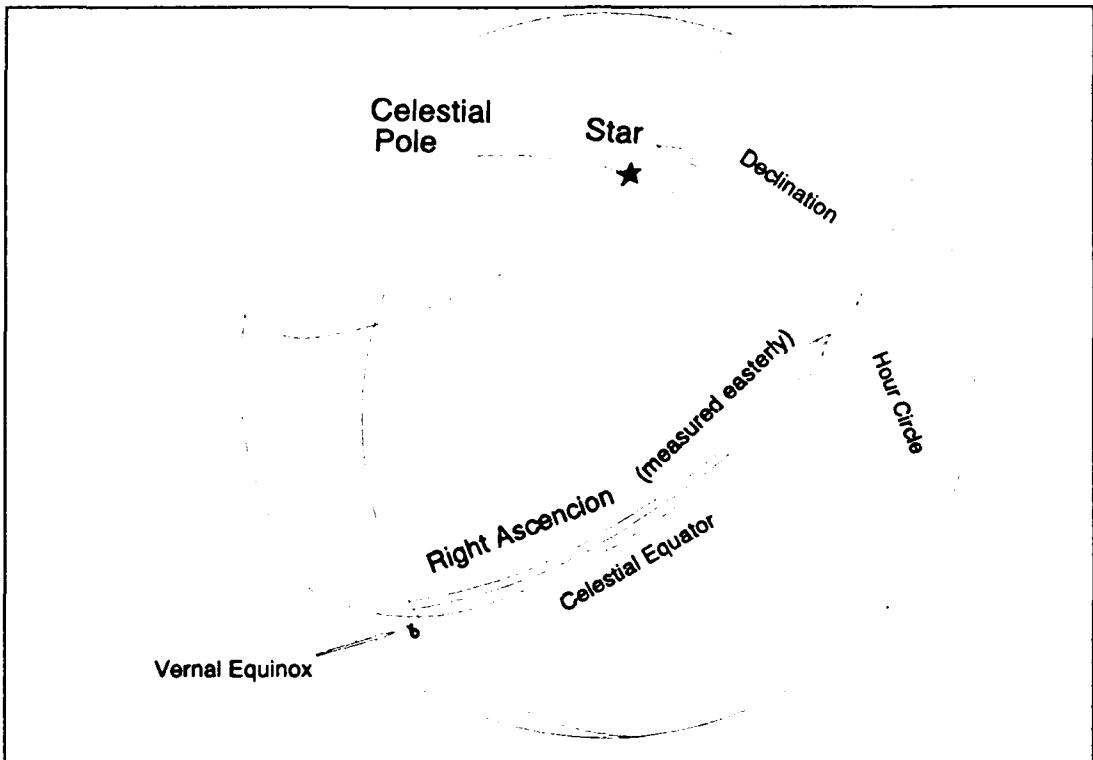


Figure 4 - Right Ascension and Declination

circle, to the body. A body lying due west of an observer would therefore have an azimuth angle of 270 degrees.

B. CONVERSION OF CELESTIAL TO TERRESTRIAL COORDINATES

Safety of the telescope was the primary design criterion. For alignment and safety reasons the telescope could not be pointed below an elevation of 15° , even during testing of the new system. Therefore, care is taken to predict the position in which the telescope would be during a movement. To accomplish this, the celestial coordinates of the active star are converted to their respective terrestrial coordinates, in the form of azimuth and zenith angles. The two coordinate

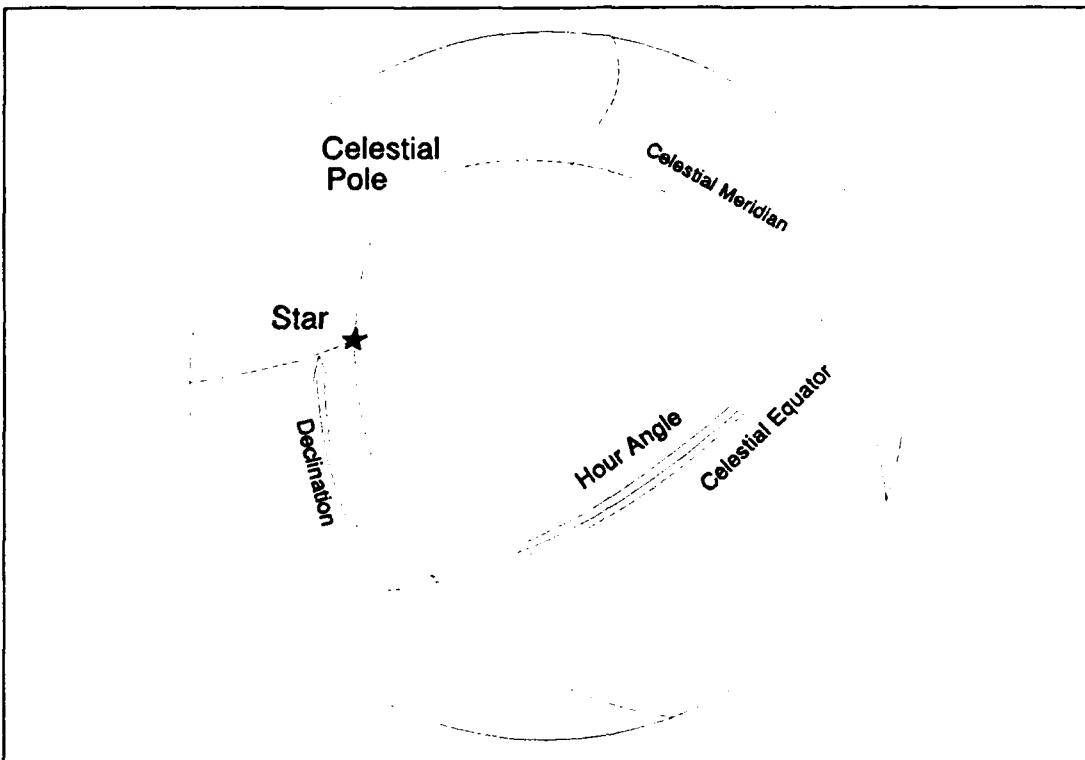


Figure 5 - Hour Angle and Declination

systems are related through a direction cosine approach as follows:

A rectangular coordinate system is determined on the Earth's surface, at the latitude (L) and longitude (γ) of the observing site. This coordinate system is designated by the letter A. Local north is aligned with the number one axis, the number three axis points to the zenith and the number two axis completes a right-handed orthogonal set. Figure 6 shows this arrangement.

The A-frame is related to an intermediate frame, the B-frame, as shown in Figure 6. The number three axis of the B-

frame points to celestial North. This relationship is shown in direction cosine notation in Equation 2.

The B-frame is related to another intermediate frame, denoted by the letter C, by Equation 3. The angle between the number one axes of frames B and C is equal to the longitude of the observing site, measured westerly from Greenwich.

The final, celestial, frame of reference is designated the D-frame. It is related to the C-frame by the Greenwich Hour Angle of Aries, which is the angle between number one axes. The Greenwich Hour Angle of Aries is denoted as $GH\Delta\gamma$. This rotation is shown in direction cosine notation in Equation 4.

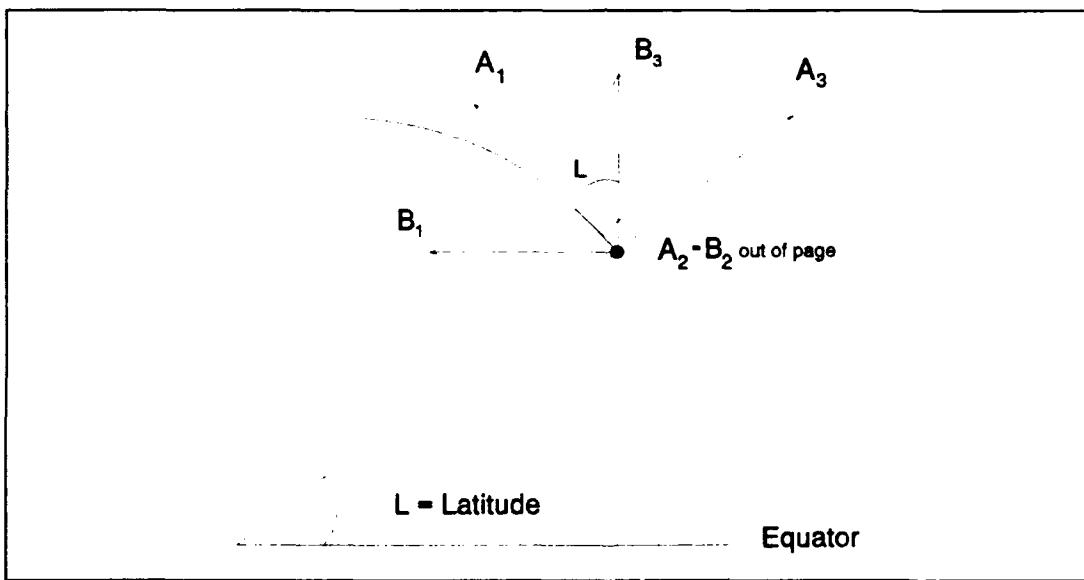


Figure 6 - Definition of the A and B Frames

$$D^B C^A = \begin{bmatrix} \sin(L) & 0 & -\cos(L) \\ 0 & 1 & 0 \\ \cos(L) & 0 & \sin(L) \end{bmatrix} = [\overline{B}_i \cdot \overline{A}_i] \quad (2)$$

$$D^C C^B = \begin{bmatrix} -\cos(\lambda) & -\sin(\lambda) & 0 \\ \sin(\lambda) & -\cos(\lambda) & 0 \\ 0 & 0 & 1 \end{bmatrix} = [\overline{C}_i \cdot \overline{B}_i] \quad (3)$$

$$D^D C^C = \begin{bmatrix} \cos(GHA\gamma) & -\sin(GHA\gamma) & 0 \\ \sin(GHA\gamma) & \cos(GHA\gamma) & 0 \\ 0 & 0 & 1 \end{bmatrix} = [\overline{D}_i \cdot \overline{C}_i] \quad (4)$$

$$\overline{d} = {}^D C^C {}^C C^B {}^B C^A \overline{a} \\ = {}^D C^A \overline{a} \quad (5)$$

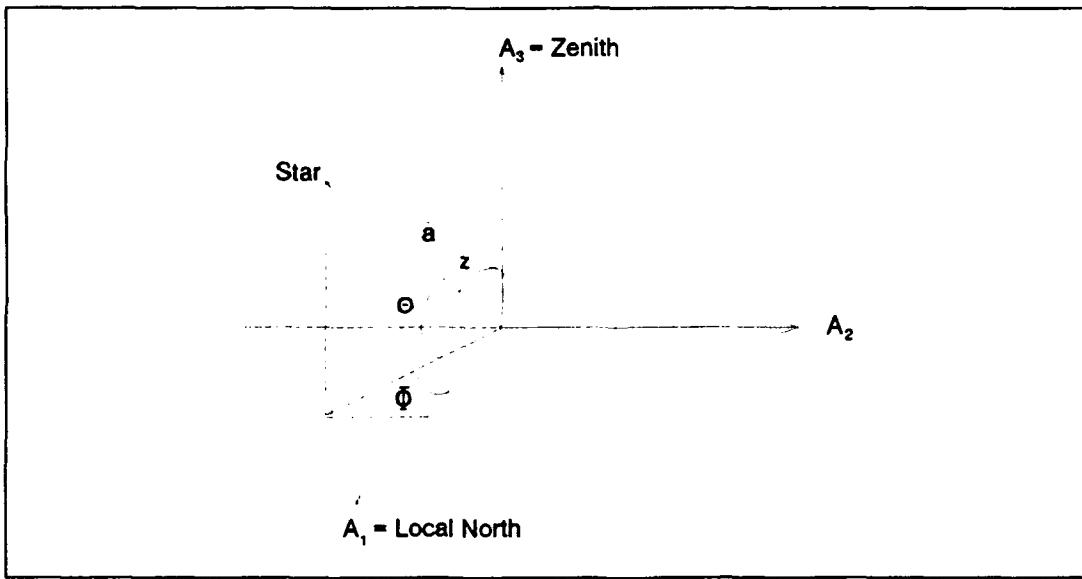


Figure 7 - Azimuth (Φ), Elevation (θ) and Zenith Angle ($z = 90^\circ - \theta$) related to rectangular coordinates

The rectangular coordinate vectors must be represented in some more useful form. The unit terrestrial coordinate vector

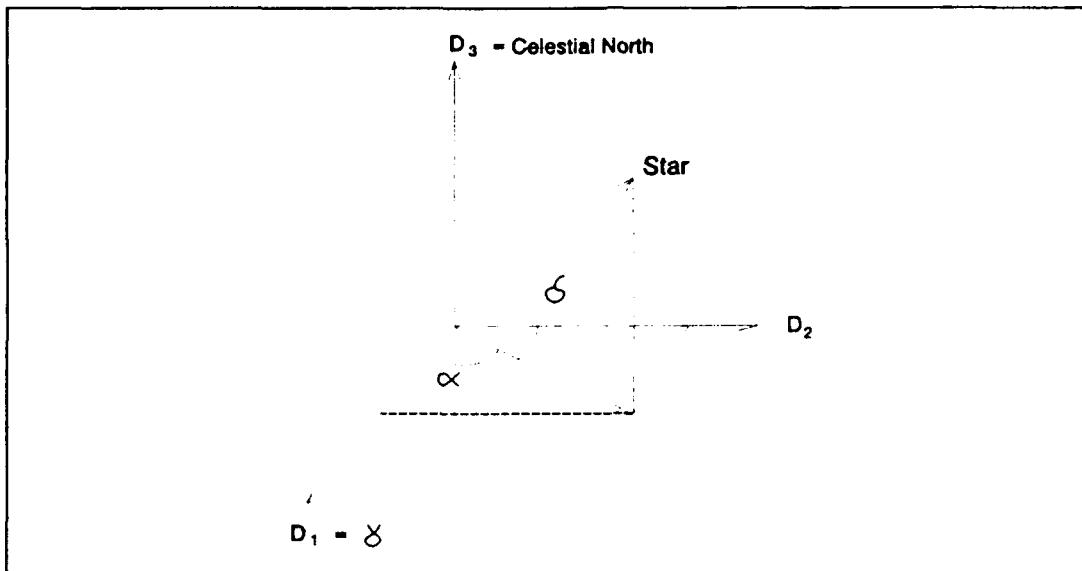


Figure 8 - Right Ascension (α) and Declination (δ) related to celestial rectangular coordinates

that points to a star, a , may be related to azimuth and elevation (or zenith angle) as shown in Figure 7. Note that azimuth is measured clockwise from the A_1 axis and the zenith angle is equal to the co-elevation. Specifically,

$$\begin{aligned}\bar{a} &= \begin{bmatrix} a_1 \\ a_2 \\ a_3 \end{bmatrix} \\ &= \begin{bmatrix} \cos(\theta)\cos(\phi) \\ -\cos(\theta)\sin(\phi) \\ \sin(\theta) \end{bmatrix}\end{aligned}\tag{6}$$

Similarly, the celestial coordinate vector may be broken into right ascension and declination as shown in Figure 8 and below.

$$\begin{aligned}\vec{d} &= \begin{bmatrix} d_1 \\ d_2 \\ d_3 \end{bmatrix} \\ &= \begin{bmatrix} \cos(\delta)\cos(\alpha) \\ \cos(\delta)\sin(\alpha) \\ \sin(\delta) \end{bmatrix}\end{aligned}\tag{7}$$

Solving for the terrestrial coordinates, assuming that α and δ are known, and noting that ${}^D C {}^A = [{}^A C {}^D]^{-1}$ for an orthogonal set,

$$\vec{a} = {}^A C {}^D \vec{d}\tag{8}$$

When solving for azimuth, quadrant dependence of the answer must be taken into account. This solution assumes that the distance between the center of the Earth and the observing site is negligible when compared to the distance from the observing site to an observed body. This is true for all observations of stars, but would not be acceptable for viewing closer objects, such as Earth-orbiting satellites.

III. TIME

A. UNIVERSAL COORDINATED TIME

Universal coordinated time (UTC) is the time standard for the National Bureau of Standards and is based on the solar day. This standard of time is based on the atomic second, which is defined as "the duration of 9,192,631,770 periods of the radiation corresponding to the transition between two hyperfine levels of the ground state of the cesium atom 133" (Reference 2, Volume 1), and is measured by a cesium beam atomic clock. In the United States, radio stations WWV (Fort Collins, CO) and WWVH (Kauai, HI) broadcast time ticks based on UTC at one-second intervals on a continuous basis. The fine timing requirements inherent in astronomical tracking make it desirable to know UTC directly. This was accomplished by reception of the WWV radio signal and decoding into time information.

B. TIME SYSTEM CONVERSION

Universal coordinated time (UTC) may be easily obtained by radio signal. Earth's rotation makes any local coordinate system a moving frame of reference. Since Earth makes one revolution in one sidereal day, this leads naturally to the use of sidereal time for astronomical calculations. Conversion between the two systems is therefore necessary.

The Greenwich mean sidereal time (GMST) is the sidereal time at Greenwich, England as referenced from the mean vernal equinox. Therefore, GMST is also the Greenwich hour angle of the mean equinox. At midnight (0 hour) on any given day GMST may be determined by the following relation:

$$GMST \text{ at } 0 \text{ hour} = 24110.54841^s + 8640184.812866^s T_U \\ + 0.093104^s T_U^2 - 6.2 \times 10^{-6} T_U^3 \\ T_U = \frac{(Julian Date - 2451545.0)}{36525} \quad (9)$$

T_U is defined as the interval of time since noon Universal Time on 1 January, 2000, expressed in Julian centuries. See Reference 1, page B6.

Greenwich apparent sidereal time (GAST) is found by applying a moderate correction to GMST to account for the difference between the true and mean equinoxes.

In the everyday world, we deal with local time as a function of the time zone in which we are. The time at our location is, therefore, the same as our neighbors. Astronomers and others concerned with precise time measurements, however, view local time much differently. Each longitude on Earth has a different hour angle from Greenwich. Local time is defined as Greenwich time plus east longitude. Therefore, you and your neighbor will only share the same local time if you are directly north or south of one another. This is known as local mean solar time.

Local mean sidereal time is then Greenwich mean sidereal time plus east longitude and local apparent sidereal time is Greenwich apparent sidereal time plus east longitude.

From the above it can be easily seen that the local hour angle, important to many calculations regarding stellar positions, may be determined by subtracting the apparent right ascension from the local apparent sidereal time.

C. TIME SYNCHRONIZATION

The Earth revolves on its axis 360 degrees per sidereal day. At this rotation rate, one tenth of a arc-second would only take six-thousandths of a second. The accuracy required for the control system therefore mandates an extremely accurate method for determining and maintaining time information.

In the United States, two radio stations transmit time standard information on a 10 MHz carrier wave. The stations are operated and maintained by the National Institute of Standards and Technology, which was previously known as the National Bureau of Standards. Radio stations WWV, in Fort Collins, Colorado, and WWVH, on Kauai, Hawaii, broadcast these signals which are synchronized to Universal Coordinated Time by way of the U.S. cesium atomic clock.

The method chosen to allow the computer system time clock to be synchronized to this time standard was an electronic radio receiver/decoder made to fit an IBM-compatible personal

computer. This board is manufactured by Odetics Corporation's Precision Time Division, formerly known as Coordinated Time Link, of Anaheim, California. It integrates an AM radio receiver with digital signal processing electronics which decode the WWV/WWVH signal into time and date information. The board is serviced by an external dipole antenna. An onboard battery and clock increment the computer's system clock between synchronizations.

It should be noted that failure of the CTS-10 onboard battery will cause the computer's system clock to be reset to the CTS-10 factory default (afternoon of 19 December, 1991) until the system has re-synchronized.

D. CORRECTING COORDINATES FOR TIME

Many stars are listed in catalogs which give their celestial coordinates for a certain date and time. These reference dates and times are known as epochs. Epoch is Latin for 'time'. Epochs differ primarily due to precession and nutation in the interval separating them. The plane of the Earth's equator, and hence the celestial equator, varies with precession and nutation. Referencing stars' positions to a standard epoch in effect references them to an inertial frame of reference. The two most common epochs for celestial reference are B1950.0 and J2000.0. As one may expect, they refer to the years 1950 and 2000 A.D. The time of the coordinates is taken as midnight on 0 January of that year.

Failure to account for differing epochs can lead to substantial errors in the celestial coordinates of a body. For example, precession alone will cause an error of approximately seven minutes of time in right ascension and five minutes of arc in declination between 1992 and 2000.

The Astronomical Almanac (Reference 1), updated every year, also lists bright stars, with the coordinates given at one-half of the way through the current year. A bright star listed in the 1992 Astronomical Almanac would then be referenced to epoch 1992.5.

Of the corrections made to correct a set of coordinates for time variance, precession is the one commonly referenced to an epoch. The others, nutation and aberration, may be determined separately, through formulae which are known for a given interval of years. Refraction, obviously, does not change with the Earth's position, but is a function of the atmosphere and local stellar coordinates.

Precession corrections in this thesis to the current epoch, i.e. the current date and time, are made by referencing a star to J2000.0, then back to the current epoch. This assures a commonality of treatment for all stars entered.

IV. CORRECTIONS TO CELESTIAL COORDINATES

The Earth is not a perfect sphere of uniform density (i.e. a centrobaric body). This causes close celestial bodies to exert disturbing gravitational moments on it. Disturbing moments with long periods of oscillation are called precession, those with short periods are called nutation. These moments cause the Earth's pole to "wobble" in relation to the plane of the ecliptic. This motion must be taken into account for precise angular determination of stars.

Movement of the Earth about the Sun causes apparent displacement of a celestial body called annual aberration. At the degree of accuracy required, aberration effects, too, must be used to correct a star's apparent position.

The Earth's atmosphere changes in pressure and temperature regularly, causing changes in the refraction of light incident upon it. Refraction must be modeled and its effects taken into account. For these and other reasons, a star's tabulated position must be modified in order to determine where it will actually appear in the night sky. The sight must be corrected for the phenomena of refraction, precession, aberration and nutation.

Table (1)

Correction	Typical Size	Cause	Residual Error after correction
Precession: - lunisolar - planetary	< 50 arc"/yr <.09 arc"/yr	Gravitational attraction of celestial bodies	<.05 arc"/yr N/A
Nutation	< 9.2 arc"	Lunar gravitational attraction	<.09 arc"
Aberration: - annual - diurnal	< 20.5 arc" < 0.03 arc"	motion of Earth	<.03 arc" N/A
Refraction	< 21 arc"	Earth's atmosphere	<.1 arc"
Proper motion of stars	<<.1 arc"	Motion of stars	N/A

Each of these phenomena will be described in detail and the necessary corrections explained in this chapter. Of these corrections, only precession is applied to the virtual

telescope. This is because precession is readily calculable by an astronomer and is commonly calculated for each star position taken from a tabulated reference. Nutation, aberration and refraction are calculation-intensive and are solely the responsibility of the computer. These corrections are applied to the physical telescope, but are not shown to the astronomer on the status screen. The user is thereby insulated from the minutia of the control system and can concentrate on an intuitive understanding of the telescope's position.

Table (1) summarizes corrections to celestial coordinates.

A. PRECESSION

The Astronomical Almanac, Reference 1, defines precession as:

"the uniformly progressing motion of the pole of a freely rotating body undergoing torque from external gravitational forces. In the case of the Earth, the component of precession caused by the Sun and Moon acting on the Earth's equatorial bulge is called lunisolar precession; the component caused by the action of the planets is called planetary precession. The sum of lunisolar and planetary precession is called general precession."

Precession of the Earth's axis is caused primarily by the effects of lunisolar moments due to the much larger gravitational attractions of those bodies. One precessional period is approximately 26,000 years and accounts for the apparent changing of the "North" star in different centuries.

Gravitational attraction would be symmetric about a perfect sphere of uniform mass distribution and no precession would therefore occur. Since the Earth is not a uniform sphere, attraction to distant bodies varies with density and distribution of mass. On the first order, Earth may be pictured as having a belt, or bulge, about the equator.

Figure 9 shows the "misalignment" of the Earth's North pole with the pole of the ecliptic. The angle ϵ is called the obliquity of the ecliptic and is the angle between the Earth's equatorial bulge and the plane of the ecliptic. Figure 10 shows the effects of such an arrangement.

In the figure, the lower part of the ellipsoid would be more strongly attracted to the point O than the upper part due to a smaller distance OP. This causes a moment counterclockwise acting on the body G. This moment would cause a precession of the pole on body G.

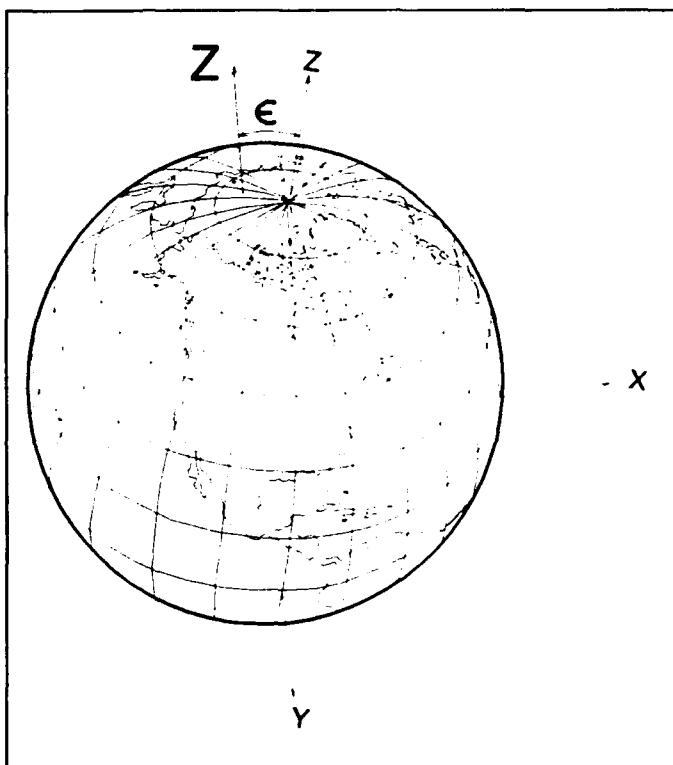


Figure 9 - Effects of nutation on coordinate systems

Precessional effects, although of very long period, are noticeable in astronomical terms. Errors on the order of arc minutes may be observed for every few years of difference from the referenced epoch if precessional effects are not taken into account. Since the MIRA requirement of a few arc seconds accuracy is far below that, precession must be taken into account.

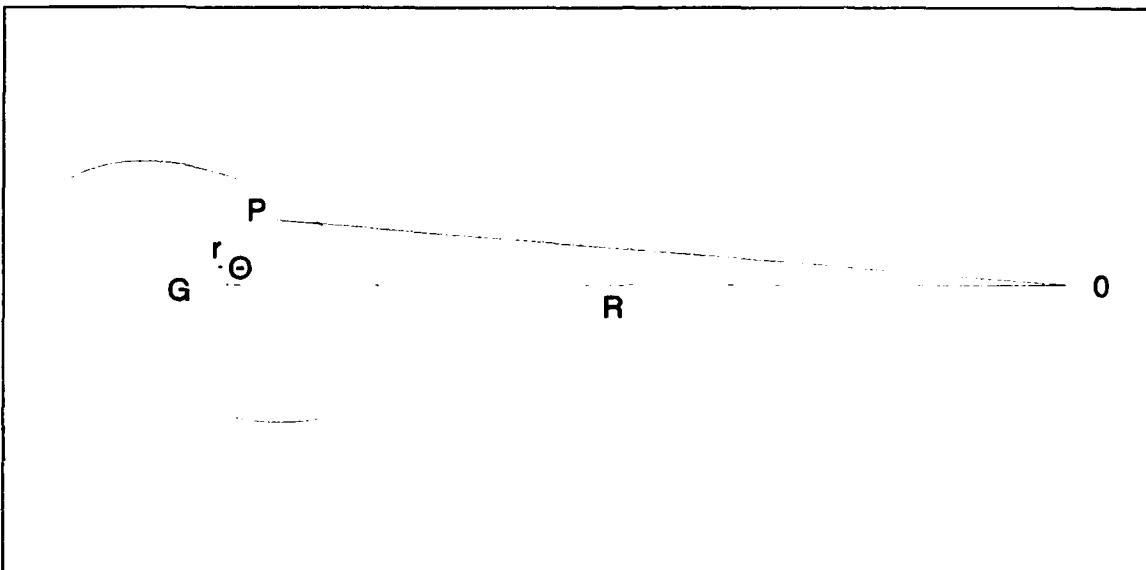


Figure 10 - Causes of nutation

Star data in the Astronomical Almanac is generally referenced to an epoch corresponding to the middle of the current year. Other star catalogs typically represent star coordinates from the standard epochs of B1950.0 or J2000.0 (See Chapter III for an explanation of epochs). Coordinates need, therefore, to be referenced to a known standard, then to the current date for use. The approach used referenced all stars first to the J2000.0 epoch and then to the current date.

Two precessional corrections were then made for each set of star coordinates.

From the Astronomical Almanac, page B19:

For reduction to J2000.0:

$$\begin{aligned}\alpha_o &= \alpha - M - N \sin(\alpha_m) \tan(\delta_m) \\ \delta_o &= \delta - N \cos(\alpha_m)\end{aligned}\quad (10)$$

$$\begin{aligned}\alpha_m &= \alpha - \frac{1}{2} (M + N \sin(\alpha) \tan(\delta)) \\ \delta_m &= \delta - \frac{1}{2} (N \cos(\alpha_M))\end{aligned}\quad (11)$$

For reduction from J2000.0:

$$\begin{aligned}\alpha_o &= \alpha + M + N \sin(\alpha_m) \tan(\delta_m) \\ \delta_o &= \delta + N \cos(\alpha_m)\end{aligned}\quad (12)$$

$$\begin{aligned}\alpha_m &= \alpha + \frac{1}{2} (M + N \sin(\alpha) \tan(\delta)) \\ \delta_m &= \delta + \frac{1}{2} (N \cos(\alpha_M))\end{aligned}\quad (13)$$

where the subscript zero refers to epoch J2000.0 and the subscript m refers to the mean epoch.

The precessional constants M and N are given by:

$$\begin{aligned}M &= 1.2812323^\circ T + 0.0003879^\circ T^2 + 0.0000101^\circ T^3 \\ N &= 0.5567530^\circ T - 0.0001185^\circ T^2 - 0.0000116^\circ T^3\end{aligned}\quad (14)$$

where $T = (\text{Julian Date} - 245\ 1545.0)/36\ 525$

The precessional coefficients of M and N change less than 0.00000005 degrees per year around the J2000.0 epoch. Since it would take approximately 50 years to accumulate errors on the one-tenth arc second order of magnitude, no attempt was made to change the precession coefficients annually.

B. NUTATION

Nutation, or "nodding", is another name for short-period oscillations in the precession moments. For the Earth, approximately 18.5 years describe one nutation period. When compared to the Moon, the Sun has almost no short-period effects and they are normally neglected. The effect of the Moon's gravitational attraction to the Earth's ellipsoidal shape is therefore responsible for the "nodding" of the Earth's pole.

$$\begin{aligned}\Delta\Psi &= -0.0048^\circ \sin(318.5^\circ - 0.053d) \\ &\quad - 0.0004^\circ \sin(198.8^\circ + 1.971d) \\ \Delta\epsilon &= +0.0026^\circ \cos(318.5^\circ - 0.053d) \\ &\quad + 0.0002^\circ \cos(198.8^\circ + 1.971d)\end{aligned}\tag{15}$$

where d = Julian Date - 244 7891.5
 ϵ = 23.44° = Obliquity of the ecliptic.

The corresponding rotation matrix is

$$N = \begin{bmatrix} 1 & -\Delta\Psi\cos(\epsilon) & -\Delta\Psi\sin(\epsilon) \\ \Delta\Psi\cos(\epsilon) & 1 & -\Delta\epsilon \\ \Delta\Psi\sin(\epsilon) & \Delta\epsilon & 1 \end{bmatrix} \quad (16)$$

The celestial coordinates may be corrected for nutation by multiplying the rotation matrix and the coordinate vector in rectangular form:

$$\overline{s}_2 = N\overline{s}_1 \quad (17)$$

The new coordinates must then be decoded from rectangular form to right ascension and declination.

C. ABERRATION

Stellar objects may be apparently displaced from their true, geometric positions by the motions of the observer and the object. The finite velocity of light may also cause the apparent location of a body to differ from its actual position. The combination of these effects is called aberration. Of the motions involved, the only one which is significant to the accuracy required is annual aberration. Annual aberration is that component of aberration caused by the motion of the Earth about the Sun.

First order annual aberration corrections to right ascension and declination are given by:

$$\begin{aligned}\Delta\alpha &= \frac{-k\sin(\lambda)\cos(\epsilon)\cos(\alpha)}{\cos(\delta)} \\ \Delta\delta &= -k\sin(\lambda)\cos(\alpha)\sin(\delta) \\ &\quad + k\cos(\lambda)[\cos(\epsilon)\sin(\alpha)\sin(\delta) - \sin(\epsilon)\cos(\delta)]\end{aligned}\tag{19}$$

where λ = true longitude of the Sun, the angle from the First Point of Aries to the Sun, measured eastward in the plane of the ecliptic.

ϵ = obliquity of the ecliptic

and the constant of aberration is given by:

$$\begin{aligned}k &= \frac{2\pi a}{Pc} (1-e^2)^{-\frac{1}{2}} \\ &= 20.496 \text{ arc seconds}\end{aligned}\tag{20}$$

where $a = 1.49598 \times 10^{13}$ cm = 1 AU

$e = 0.01675$ = mean eccentricity of the Earth's orbit

$P = 3.1558 \times 10^7$ sec = length of sidereal year

$c = 2.997925 \times 10^{10}$ cm/sec = velocity of light

The true longitude of the Sun is given by:

$$\lambda = 278.926^\circ + 0.98567d\tag{21}$$

where d = day of year + fraction of day since 0^h UT

The true longitude of the Sun is an approximation good to about one arc-second. However, since it affects aberration only through a product of trigonometric expressions, the above approximation of aberration is adequate to the one-tenth arc-second level.

D. REFRACTION

When a photon transits from one medium to another in its travels, it refracts. In other words, the direction of its travel is altered. For the case of a photon traveling through space from a distant star and impacting the Earth's atmosphere (a denser medium), the photon is refracted toward a vector normal to the atmosphere. This phenomenon is shown in Figure 11 below and described by Equation 21.

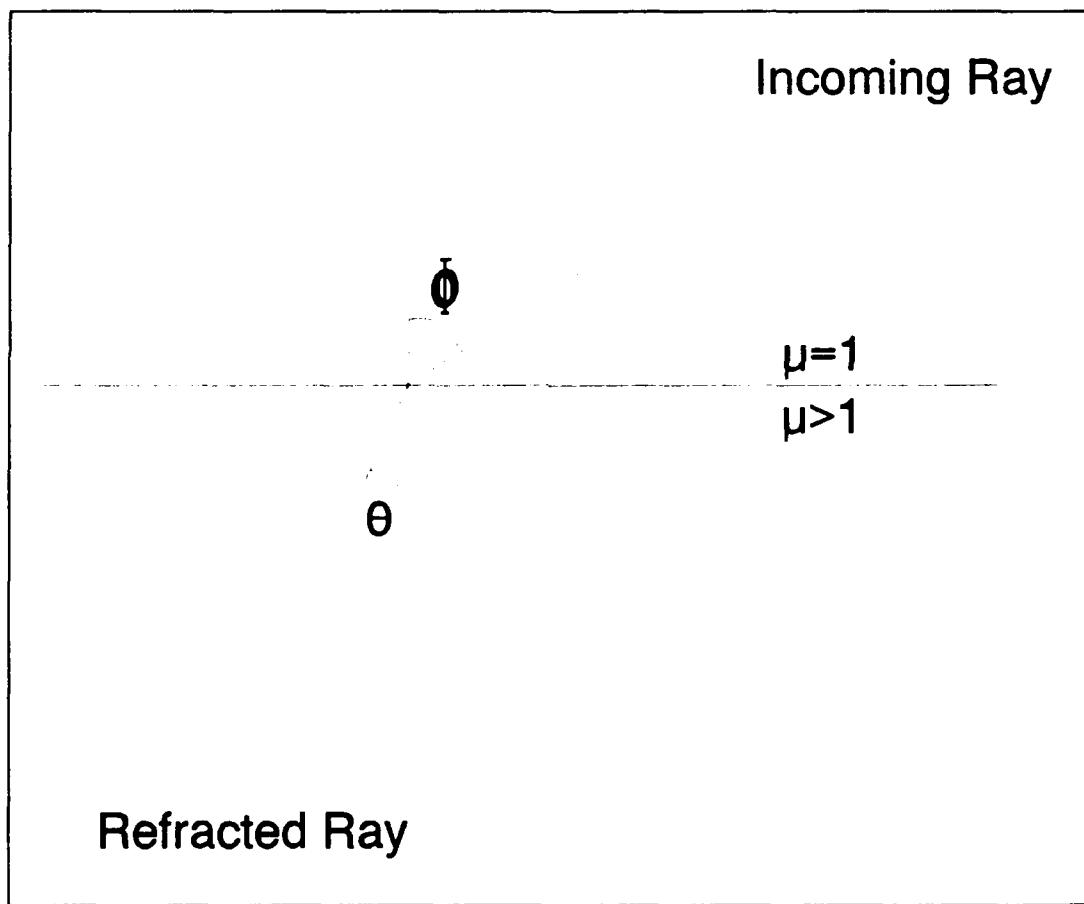


Figure 11 - Refraction through a medium

$$\mu_1 \sin(\Phi) = \mu_2 \sin(\Theta) \quad (22)$$

where μ_1 = The index of refraction of space (= 1),
 μ_2 = The index of refraction of the atmosphere,
 Φ = the angle of incidence between the incident ray and
the atmospheric normal,
 Θ = the angle of refraction between the refracted ray
and the atmospheric normal.

Since the Earth's atmosphere is constantly changing, refraction is dependent upon many factors and a single index of refraction is not possible to determine. Of these factors, the most prominent are atmospheric pressure and temperature and the wavelength of the observed light.

The MIRA telescope operates with zenith angles of less than 75 degrees for safety of the primary mirror. The following formulae allow a reasonably close approximation of atmospheric refraction for zenith angles less than 75° (more than 15° off the horizon) (See Reference 5).

$$K = \frac{R}{\tan(z_o)} \quad (23)$$

$$R = \frac{21.3''P \left(1 + 0. \frac{0057}{\lambda^2}\right) \tan(z_o)}{273 + t} \quad (24)$$

$$d\alpha = \frac{K \sec^2(\delta) \sin(H)}{\tan(\delta) \tan(\Phi) + \cos(H)} \quad (25)$$

where α = star's uncorrected right ascension
 α_0 = observer's right ascension
= GHA γ - observer's longitude
P = atmospheric pressure (mmHg)
t = atmospheric temperature ($^{\circ}$ C)
 λ = center wavelength of star's light (μ)
 δ = star's uncorrected declination
H = star's hour angle = $\alpha_0 - \alpha$
 Φ = observer's latitude

$$d\delta = \frac{K(\tan(\delta) \cos(H) - \tan(\Phi))}{\tan(\delta) \tan(\Phi) + \cos(H)} \quad (26)$$

This treatment of refraction can be extremely sensitive to changes in the atmospheric parameters. The total refraction correction can change on the order of one-tenth of an arc second in less than 1.7 seconds under nominal atmospheric temperature (10° C) and pressure (633 mmHg) at a nearly zero zenith angle. Refraction was therefore updated every one second, with the clock displays. Temperature changes of a single degree Celsius or pressure changes of less than a single millimeter of mercury can alter the refraction corrections by approximately one-tenth of an arc second at nearly nominal atmospheric conditions at a nearly zero zenith angle. While manual entry of these parameters is acceptable

for short period observations, eventual upgrade to automatic temperature and pressure sensing is recommended.

V. HARDWARE

A. SYSTEM OVERVIEW

The hardware to be added to the telescope system is shown in Figure 12 and may be compared to Figure 1, the current hardware configuration.

An IBM-compatible personal computer, based on an Intel 80486 microprocessor, is to be used as the controlling

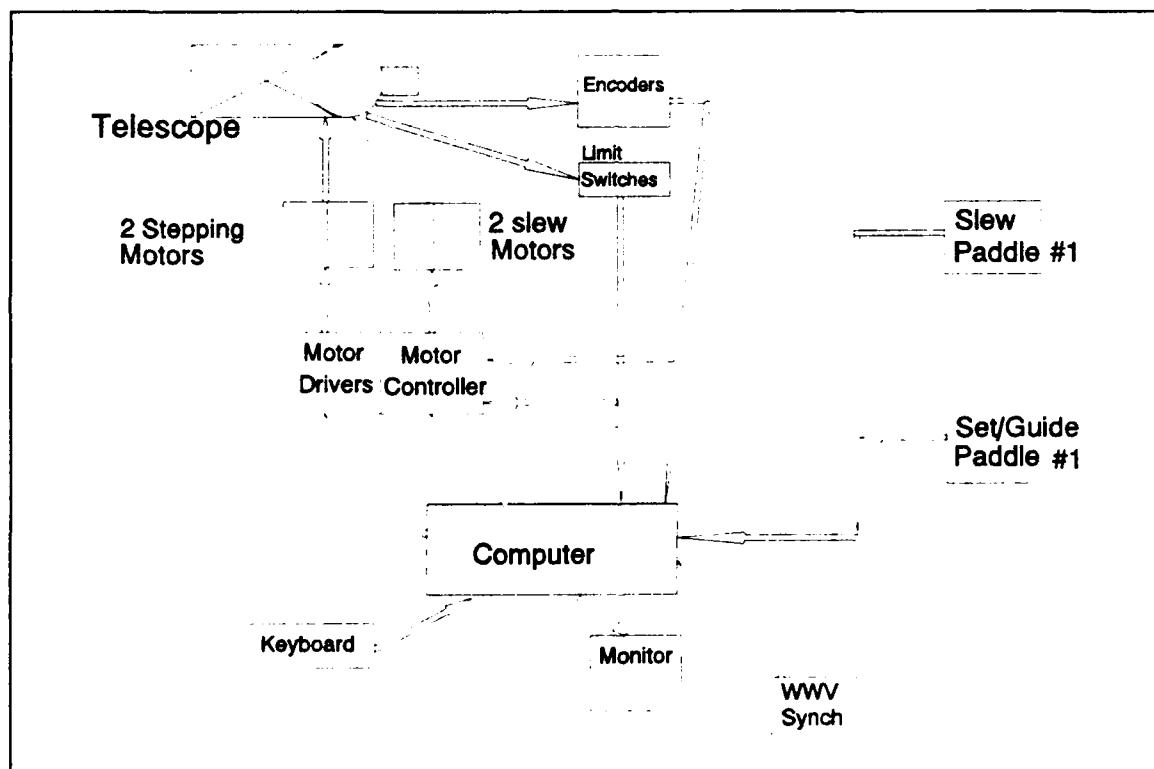


Figure 12 - Proposed hardware configuration

authority. System clock time is to be updated automatically via the WWV time synchronization board discussed in Chapter

III. Slew and set/guide paddles are to be interfaced to the computer via a TTL interface board. Motor drivers are either in an "on" or "off" state. The state is to be controlled by relays operated through an additional TTL interface board.

Physical limit switches are installed to prevent the telescope from being moved beyond its limits in the case of software failure. Two sets of physical limit switches are installed. The first set of switches cuts power to the slew motor controller in the direction required when the telescope's zenith angle reaches 76°. Operation in the opposite direction is allowed to facilitate recovery. This set is engaged first and will stop telescope slew. The second set of limit switches cuts power to the stepping motor drivers, thus preventing the stepping motors from driving the telescope to an unsafe position. In this case, also, only movement in the affected direction is stopped.

B. COMPUTER

An IBM-compatible personal computer was chosen for the system due to availability, cost, speed, and operating considerations. Of these, familiarity of the MIRA staff with the IBM-compatible operating system and the Microsoft QuickBASIC programming language were the key factors. A machine based on an Intel 80486 microprocessor was chosen to allow the system to be operated in a multi-tasking

environment. This eliminated the need for a dedicated computer, cost again being a major consideration.

To allow multi-tasking, the following hardware requirements are recommended as a minimum:

- 80486 microprocessor
- 4 MBytes RAM
- 1 hard disk drive
- 1 floppy disk drive
- CGA monitor
- a mathematics co-processor if a lower CPU is used

If multitasking is not desired, the computer requirements may be downgraded to 1 MByte RAM and an Intel 80286 microprocessor with an 80287 mathematics co-processor.

This type of computer comes standard with eight bus interface slots. Of these, at least five would be needed. One slot would be required for each of the TTL boards for interface to the two control paddles and the two types of motor controllers. One additional slot would be required for the WWV board. Additionally, one RS-232 standard serial interface port would be required for the encoder inputs. This type of port also comes standard with almost any brand of IBM-compatible computer.

VI. PROGRAM DEVELOPMENT

A computer program was developed in the Microsoft QuickBasic Language, version 4.00, to point the telescope to predetermined celestial coordinates and correct for any errors larger than one-tenth of an arc-second. The program uses rotation matrices, developed in Chapter II, to translate between celestial and terrestrial coordinate systems.

The program structure is modular. The main program controls 24 subroutines which deal with data display, data entry, data correction, time correction, key and video handling, and telescope movement. Each subroutine is described below, within its appropriate category.

The concept of a virtual telescope was fundamental to the software development. A 'virtual' model of the system, encompassed solely in software, was developed and the effects of any operation on this model were evaluated prior to any action being taken. This 'virtual' model is seen by the user and represents the target star's position corrected solely for precession. Corrections to raw star position data for precession, nutation, aberration, and refraction were completed and new motor states determined. These corrections and motor states determine the status of the physical telescope. The new position of the physical telescope was determined in both terrestrial and celestial coordinates. If

no errors were encountered, such as hard- or software limits to movement, the telescope was then issued motor commands.

Fig 13 shows the highest-level modular flow of the program.

A. MAIN PROGRAM DESCRIPTION

The main program, beginning on page 1 of Appendix A, controls the flow of telescope operation and includes the main operating loop. A simplified flow chart of the main program is shown in Figure 14.

The program begins by initializing system hardware. It then proceeds to allow user entry of daily atmospheric and observation parameters, if required. The telescope is aligned by pointing to a known bright star. The star's celestial coordinates are entered, thereby giving the system a reference point. Data entry of desired stars for view is accomplished either by direct entry or by reading a previously written and formatted file. Editing of a stored file is allowed. A star is then selected for view. Once a star is selected, the program ensures that the star is visible above the current local horizon. If it is, the main loop is entered, allowing the system to find and track the star. If it is not, another star must be selected.

Since safety of the telescope structure was the primary design criterion, several layers of protection were utilized in the hardware and software. At any time that a zenith angle

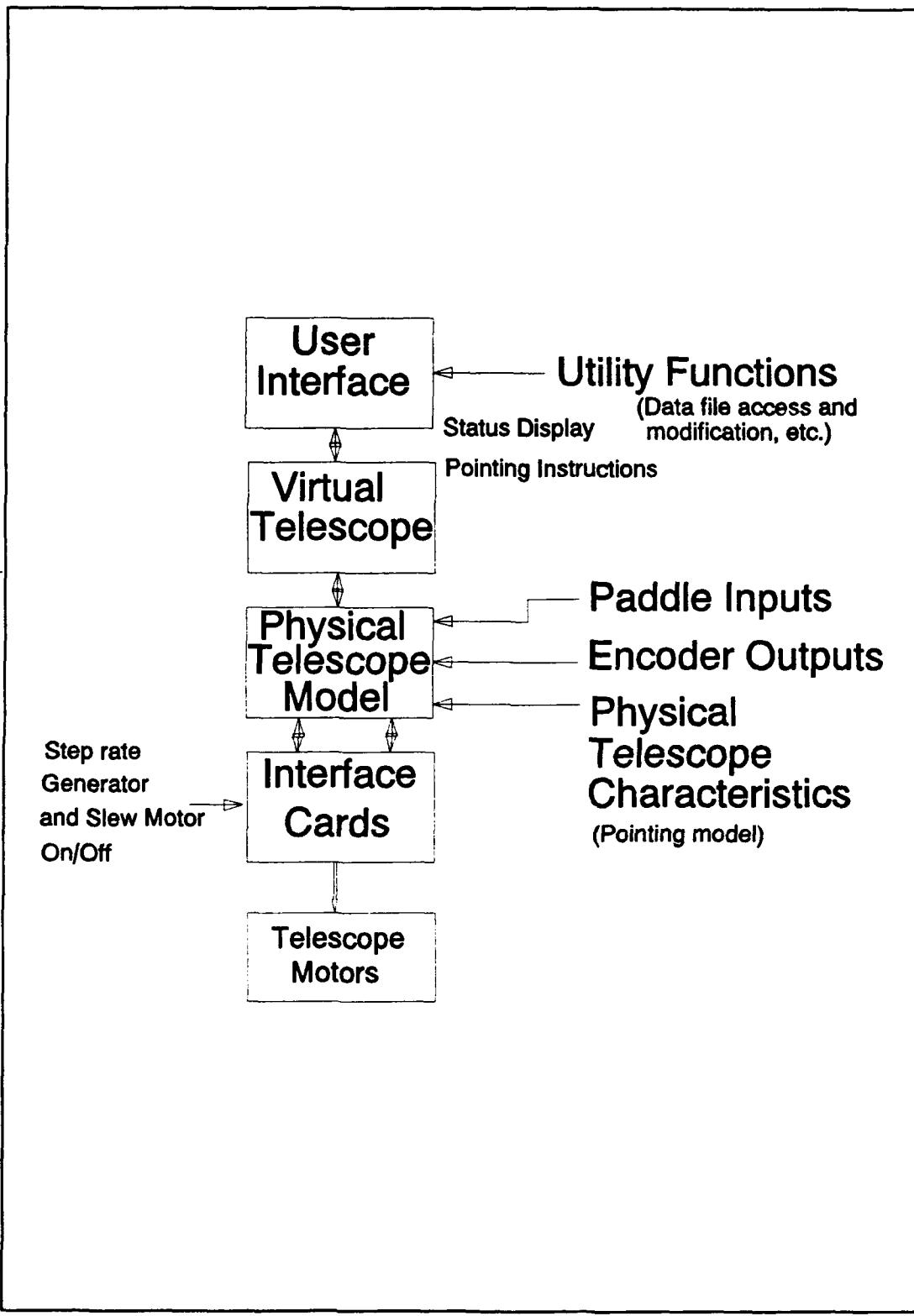


Figure 13 - Program modular concept

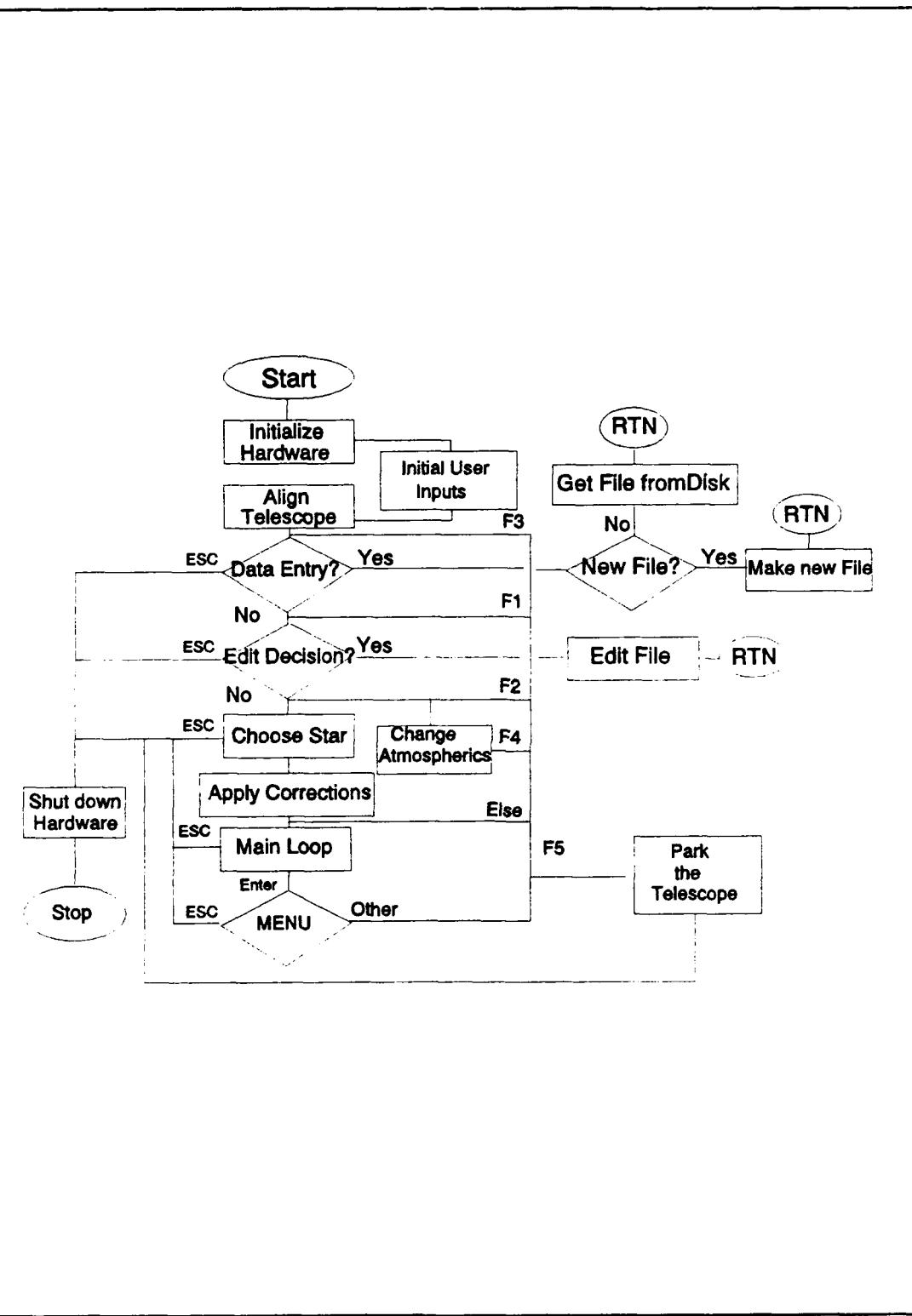


Figure 14 - Simplified flow chart of the main program

of 65 degrees or greater was calculated, a warning message would appear to the operator on the telescope status display screen. When a 75 degree zenith angle was reached, the software would automatically drop track on the target star and ask the operator to choose a new star for view. In the event of catastrophic software errors that allowed the telescope to move beyond a 75 degree zenith angle, physical hardware limit switches were installed at a 76 degree zenith angle. Two sets of limit switches exist on the telescope. The first set cuts power to the slew motor controller. This set of switches must be engaged first due to the speed involved in slewing and the time that the system requires to decelerate. The second set of switches cuts power to the stepping motor drivers.

A simplified flow chart of the main program loop is provided in Figure 15. In this loop the main tracking actions take place. Every second, time and position information is updated to the telescope status screen. The target star's right ascension, declination, hour angle, zenith angle and azimuth are displayed. Movement commands are issued to keep the differences between the stars' coordinates and the telescope's within acceptable limits. The loop also checks once per loop for user input. An ESC key hit stops telescope movement immediately and awaits user confirmation to quit the program. If confirmation is received, the program is exited via a procedure to safely shut down all hardware. If the ENTER key is hit, the change-telescope-status menu appears,

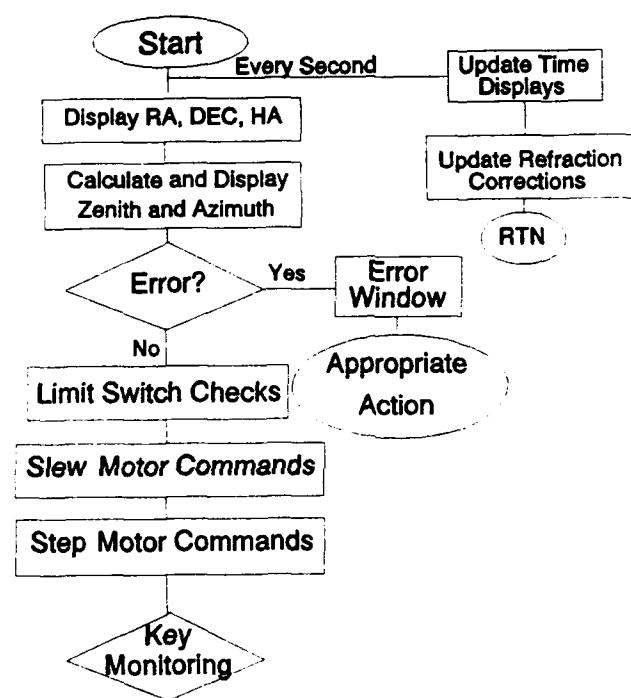


Figure 15 - Simplified flow chart of the main loop

allowing the user several options including the ability to proceed to another star.

Once the virtual telescope's right ascension and declination, the actual telescope's azimuth and zenith angles, the star's hour angle and the times are ready for display in the main loop, the TeleStatusDisplay subroutine is called to display them. It updates the Pacific Standard Time, sidereal time, and universal coordinated time clocks once per second, along with the updated refraction correction.

B. SUBROUTINE DESCRIPTIONS

1. Display Subroutines

Display subroutines are responsible for the display of data and graphics to the computer screen.

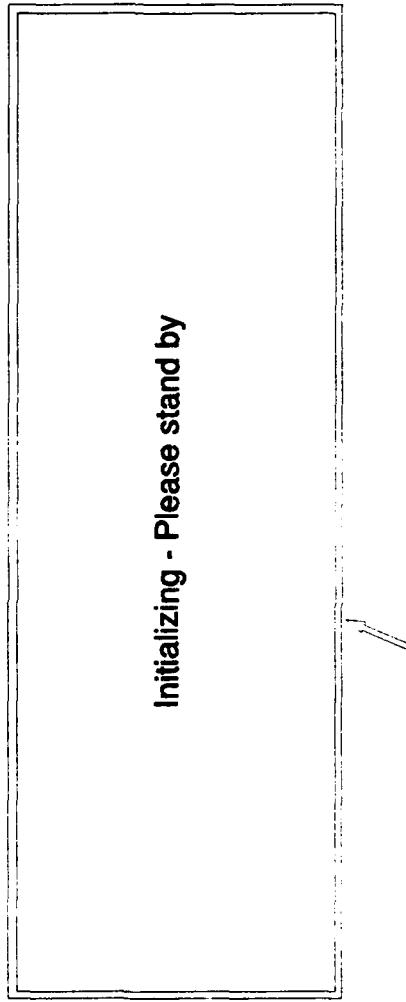
a. *InitialDisplay*

This subroutine creates the display screen graphics which will be utilized for the remainder of the program run. A user box is displayed at the bottom of the screen, in which all user interaction and display occurs. The user box, as shown in Figure 16, is simply a screen area for data exchange. Additionally, menus and error messages occur in windows which occupy other video screens.

This subroutine also colors the background and identifies the program and its intended location to the user.

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User box

Figure 16 - The initial display screen with user box

b. *StarDisplay*

This subroutine modifies the initial display screen for star data entry. It draws the individual fields in which the parser (see subroutine *StarDataEntry*) will operate and also prints instructions for leaving the data entry routine.

c. *Windows*

All error messages and most menus are displayed on video pages which are separate from the primary display page. These messages and menus are pre-defined data structures and are displayed by this subroutine. A rectangular window is placed at a position determined in the data structure. The size of this window is dependent on the number of lines of text to be displayed and the length of the longest line.

To determine the video parameters and if the next video page is available, the subroutine *VideoState* is called.

A window is removed by calling the *WindowsPop* subroutine.

d. *WindowsPop*

This subroutine removes the last window placed by the subroutine *Windows* from the active video page and makes the next lowest video page active.

e. *TeleStatusDisplay*

This subroutine modifies the user box for the display of telescope and star positions and time during the

main loop. The name of the active star is placed in the appropriate box.

f. ShowTime

There are three different times displayed on the status screen during telescope movement, universal coordinated time (UTC), mean sidereal time and Pacific Standard time (PST). These three times are updated and displayed through this subroutine.

g. TeleAngles

Using the celestial coordinates to which the telescope is pointed, this subroutine determines the position of the telescope in terrestrial terms. The terrestrial rectangular coordinates are decoded into azimuth and zenith angles. Azimuth is defined as the number of degrees measured eastward in a circle from North. North is therefore 0 degrees, East is 90 degrees, etc. Zenith angle is defined as the number of degrees from the point directly overhead the telescope to the telescope pointing vector. Zenith may also be thought of as the co-elevation, where zenith equals 90 degrees minus the elevation. Elevation is the number of degrees from the local horizon to the telescope pointing vector.

The above angles are calculated by means of direction cosine matrices. These matrices are described in detail in Chapter II.

2. Data Entry Subroutines

Data entry subroutines all accept user input prior to proceeding. User input is necessary to align the telescope, to provide the program with atmospheric parameters and a list of stars on which the telescope shall sight.

a. TeleAlign

The telescope must have an alignment point to which all future movements will be referenced. This subroutine accepts user input to determine the right ascension and declination of the telescope pointing vector at a given time. Coordinates from encoder output signals are then treated as a difference from this baseline.

Typically, the right ascension and declination of the present pointing vector will be entered from the bright star list of the Astronomical Almanac. Therefore, the epoch of the coordinates must be given and accounted for. This subroutine precesses, nutates, aberrates and refracts the given coordinates to account for the difference between tabulated and apparent positions of the bright star chosen for alignment.

b. Atmospheric

Refraction through the atmosphere depends primarily on four variables: the zenith angle, the atmospheric temperature, the atmospheric pressure and the center wavelength of light from the star being observed. The latter

three must be provided to the program in order to be accounted for. This subroutine allows user entry of these variables.

c. *StarDataEntry*

This subroutine is a parser for accepting data for up to one hundred stars. Four fields are required for each listing: star name, right ascension, declination and the epoch to which the right ascension and declination refers. The subroutine displays five lines of the listing at one time and allows scrolling, page jumps and home or end-of-file jumps. Full word-processing ability has been included for appropriate keys. Insert, delete and shift keys are active, with their usual functions.

No attempt was made in this procedure to ensure the correctness of the formatted data. If the data formats are not correct, the program will later display an error message and allow the listing to be edited.

As the listing is updated, the matrix variables Star\$(), RA\$(), DEC\$() and EPOCH\$() are updated. These four variables are each one hundred elements in length and hold the star names, right ascensions, declinations and epochs, respectively, in string form.

Figure 17 shows the *StarDataEntry* screen.

d. *ChooseStar*

Once a listing of stars has been read into the matrix variables Star\$(), RA\$(), DEC\$() and EPOCH\$, one star

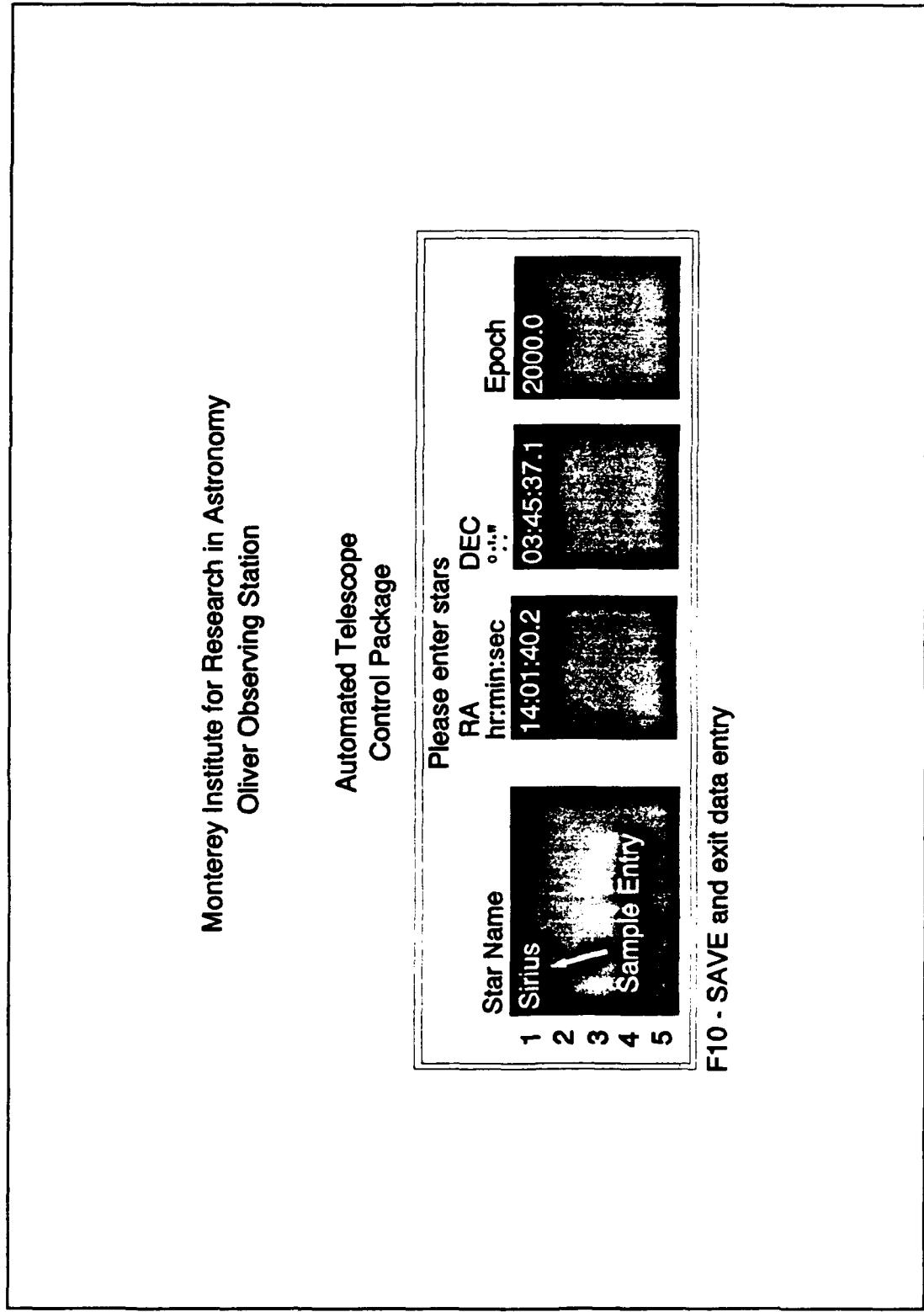


Figure 17 - StarDataEntry screen

must be chosen for view. This is referred to as the "active star". An active star will therefore have associated with it a name, a right ascension, a declination and an epoch.

This subroutine modifies the user box to show a listing of the one hundred star entries in the current file. This file will be on either a floppy or hard disk and may have any file name. The default name is "STARFILE.DAT". The subroutine allows the user to scroll, page up or down, jump home or end-of-file. Once the selected star is highlighted, it is chosen for view by pressing the F10 key. The active star's four elements are then passed to the main loop.

3. Data Correction Subroutines

As described in Chapter IV, the movement of Earth around the Sun, the wobble of Earth's axis, the rotation of the Earth, the relativistic effects of the Sun's presence and the refraction of light through the atmosphere must be accounted for. The data correction subroutines make adjustments to the right ascension and declination of the active star to account for these effects.

a. Precession

The effects of precession on astronomical sightings were discussed in Chapter IV. The Precession subroutine uses the approximate formulae for precession found in the Astronomical Almanac, page B19. The epoch of the active star is first precessed to the standard epoch of J2000.0, then to

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**Automated Telescope
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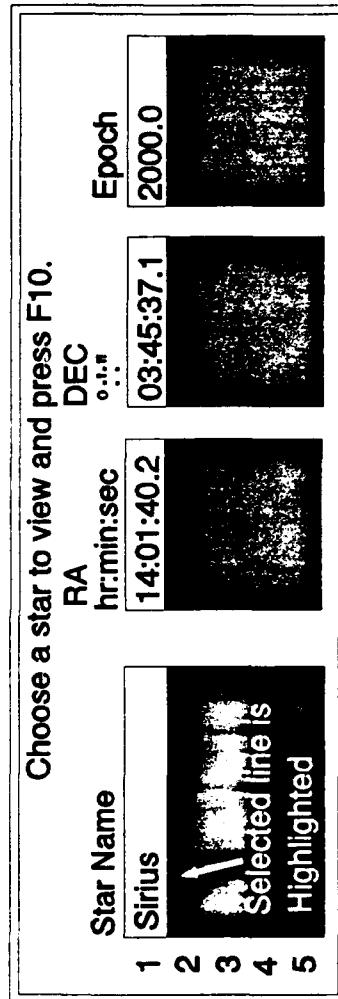


Figure 18 - ChooseStar screen

the current epoch. The current epoch is calculated to the current day. Right ascension and declination for the current star are then modified to account for the difference between the epoch entered by the user and the current epoch. Precession is accounted for at the beginning of each star sight.

b. Nutation

Nutation is described in Chapter IV. The Nutation subroutine uses the approximate formulae for nutation found in the Astronomical Almanac, page B20. A rotation matrix is used to relate the unmodified and nutated celestial rectangular coordinates. The rectangular coordinates are then decoded into right ascension and declination. Nutation is accounted for at the beginning of each star sight.

c. Aberration

The causes of aberration are briefly discussed in Chapter IV. This subroutine accounts for aberration by determining the relation between the apparent positions of the Sun and the active star and modifying the right ascension and declination accordingly. The method used was garnered from Reference 11, pages 499-502. The mean longitude of the Sun was calculated from formulae provided in the Astronomical Almanac, page C2.

d. Refraction

The effects of refraction through the Earth's atmosphere were considered in Chapter IV. Refraction is affected by four variables: zenith angle, atmospheric temperature, atmospheric pressure and the center wavelength of light for the observation. The latter three were entered by the user in the Atmospheric subroutine. The first was calculated in the TeleAngles subroutine.

Refraction was calculated by the method employed by Reference 5, pages 86-87. Refraction corrections to right ascension and declination were calculated in the subroutine and applied in the main loop. This is necessary because refraction is updated once per second and old corrections must be removed prior to new ones being applied.

4. Time Correction Subroutine

a. SiderealTimeCalc

Sidereal time is a measure of time in relation to the stars (specifically to the vernal equinox), instead of in relation to the Sun. Local mean sidereal time is the actual number of hours, minutes and seconds from the alignment of the local meridian to the mean equinox. This subroutine calculates the local mean sidereal time and adds a minor correction to account for the difference between the true and mean equinoxes. This yields the local apparent sidereal time. This time is then displayed from the main loop.

5. Key and Video Handling Subroutines

a. KeyCode

For any key pressed, KeyCode returns a unique integer value. This value may then be used to determine the key or key combination that has been pressed. Once the subroutine is entered, the program will wait until the key buffer contains a character before returning control to the calling routine. This is especially useful in any application requiring the decoding of user response. The following call KeyCode or KeyCodeNoWait: The main program, ChooseStar, StarDataEntry and Windows.

b. KeyCodeNoWait

KeyCodeNoWait is identical to the KeyCode subroutine above except that the key buffer is checked and control is immediately returned to the calling routine instead of waiting for key input. The main loop uses KeyCodeNoWait to monitor the keyboard while the telescope is tracking a star.

c. VideoState

The video mode parameters of the computer on which the program is being run must be determined in order to place the error message windows and menus. VideoState determines the current video mode, the number of text columns and the number of the active display page. The Windows subroutine uses this data to determine if there is room to place a message window prior to placement.

6. Telescope Movement

Four subroutines control motor commands to the telescope. Each subroutine issues motor commands for a certain type of movement. Movement is categorized into slew, set, track and guide. Slew uses the one-quarter horsepower slew motors to transit the telescope through large angles. This places the telescope in the neighborhood of a target star's position. Set, track and guide all use the smaller stepping motors. Set is used to adjust the telescope's position after slewing and align the telescope's pointing vector to the target star. Once slew and set have occurred, track changes the right ascension of the telescope at a constant rate to account for the Earth's rotation. As refraction, movement and other errors cause misalignments, guide adjusts the telescope's position.

To ensure structure safety, the position of the gravity limit switch is checked after every step of the stepping motors and periodically during use of the slew motors.

Arrow keys are monitored from the main program before each track update to allow manual adjustment of the telescope's position.

a. *SlewCommands*

The *SlewCommands* subroutine, immediately after a target star is chosen, slews the telescope to within one

degree of the calculated position. On successive loops, movement is left to the stepping motors in the set (fast) or guide (slow) modes.

The maximum slew motor rate is approximately 12 degrees per second. To ensure the safety of the telescope, the gravity limit switch position is checked every degree of movement or every eight-hundredths of a second during slewing.

b. SetCommands

This subroutine commands the stepping motors following a slew to a new target star. Once the SlewCommands routine has positioned the telescope to within one degree of a calculated position, this subroutine uses the stepping motors, in the fast mode, to bring the position to within one step of the star's calculated position. The astronomer operating the system will then guide the telescope to the star's actual apparent position by operating the stepping motors manually, either by the hand paddles or the arrow keys.

c. TrackCommands

Once a star has been found, the TrackCommands subroutine will update the right ascension motor stepping rate to adjust for changes in the atmospheric refraction. The telescope will, at this time, be slowly sweeping in right ascension to compensate for the rotation of the Earth.

C. ERROR AND WARNING MESSAGES

Error handling is an important aspect of any computer program. To prevent a program from halting catastrophically upon encountering an error, errors must be anticipated and dealt with appropriately. Additionally, some situations may be anticipated that require that the user be warned of impending problems. The main program contains ten error and warning windows to advise the user of anticipated situations.

1. The Quit Confirmation Window

Following standard practice, the ESC key is used to abort a telescope tracking run or to exit the program entirely. Upon striking the ESC key telescope movement is halted, if the telescope is moving. The quit-confirmation window then appears. If the user types either an upper or lower case "Y", the program is terminated. Otherwise, control returns to that segment of the program previously running.

2. The File Save Error Window

When a data file is to be saved, many errors may occur. The appropriate disk drive may not be functioning, the drive may not contain a disk, the disk may be write-protected, the disk may be full, etc. In any of these cases, the file-save-error window appears to inform the user that the data file could not be saved. Upon the press of any key, control returns to the file editor. From this point the user may correct the error or quit the program, as appropriate.

3. The File-Not-Found Error Window

When the user requests that an old data file be used, the program requests the path and file name to retrieve. If the path does not exist or if the file name cannot be located, the file-not-found-error window appears. This window informs the user of the error and waits for any key to be pressed. Control is then returned to the data entry screen menu, allowing the user to choose an old file, re-enter the file name and path, or begin a new file.

4. The Equinox Correction Error Window

Corrections to the Greenwich mean sidereal time are in the form of a number, in seconds. The default values provided are zero seconds for both the current and next days. Should the user fail to except the default values and enter a correction that could not be interpreted as a number in seconds, the equinox-correction-error window appears. Once a key is pressed, the user is allowed to except the defaults or re-enter new numbers.

5. The Slew Confirmation Window

Before the telescope is slewed to a new star, the new target star's coordinates are calculated. The zenith angle anticipated is displayed, along with a warning that the F1 key will cause the telescope to slew. Although not displayed on the screen, the F10 key will cause the program to run in

SIMULATION mode. Any other key will return the user to the ChooseStar subroutine.

6. The Change Telescope Status Window

The change-telescope-status window is a user menu of seven possible actions. It appears when the ENTER key is pressed during a telescope tracking run and allows the user to take one of the following actions:

- Edit the current list
- Choose a star from the current list
- Change to a new data file
- Change atmospheric parameters
- Park the telescope
- Exit the menu and return to whatever tracking was previously taking place.

7. The RA/DEC Format Error Window

Right ascension and declination of a target star are entered from a formatted data file. If the program attempts to decode these coordinates and fails, the RA/DEC-format-error window appears. The user is then given the opportunity to edit the data file or choose another star.

8. The Slew Error Window

If the user ignores the slew-confirmation window with a zenith angle over 75 degrees and attempts to slew the telescope to an unsafe position the slew-error window appears. No telescope movement is allowed. Upon the press of any key

control is returned to the change-telescope-status menu so that the user may choose another star or take other appropriate action.

9. The Zenith Angle Warning Window

The maximum zenith angle allowed is 74 degrees. When the zenith angle reaches 75 degrees telescope movement is halted. To warn the user of impending cessation of movement, the zenith-angle-warning window appears any time the zenith angle is 65 degrees or over. It displays the simple warning "The zenith angle is X degrees".

10. The Zenith Angle Halt Movement Window

Should the user ignore the zenith-angle-warning window and allow the telescope to track to its limits, the zenith-angle-halt-movement window appears at 75 degrees. All telescope movement is halted. The change-telescope-status menu appears upon the press of any key. The user may then take action as appropriate such as choosing another star. No action that the user takes can cause the telescope to automatically track to a zenith angle greater than 74 degrees.

11. The Telescope Status Help Window

When the telescope status screen is shown, help for the allowed key strokes is available through this subroutine. The F1 key will display a window listing the keys which control manual slew or guide of the telescope. The arrow keys and numeric keypad is used for this purpose.

12. The CTS-10 Failure Window

It is possible that the CTS-10 WWV time synchronization board will fail and reset. This usually indicates an onboard battery failure or a poor physical connection to the data bus of the computer. In this event, the CTS-10 will reset the computer's system clock to the factory default setting. This condition is tested in the main program upon initialization. If an error is detected, an error window will appear, advising the user to either repair the CTS-10 problem or manually update the system clock to correspond to UTC. After any key is pressed, the program will halt execution and must be run again after the error is rectified.

VII. PROCEDURES

Once final hard- and software modifications have been completed, an astronomer shall have automatic control of telescope movements to a known set of celestial coordinates. He or she will remove the telescope from its stowed configuration by removing obstructing objects, sliding the enclosing roof away from the telescope and opening the doors covering the primary mirror. After this point, procedures differ drastically from the current procedures.

The astronomer will provide power to motor controllers and the computer. Following systems checks, the MIRACTRL program will be run. The user will provide the celestial coordinates to which the telescope is aligned by looking at a known bright star. The astronomer will enter atmospheric data for the evening and provide the computer with a list of star positions either by a data file or by manual entry. He or she will choose a star to view and the program will check whether it is in view. If so, a single key stroke will slew and set the telescope to its desired position. The astronomer will guide the telescope's position onto the star and allow the computer to begin tracking. The computer will continue to track until the astronomer guides the telescope manually, the star is out of view or the gravity limit switch is engaged. The astronomer may choose to halt the telescope, exit the program,

thereby halting the telescope, or change the target star at any time.

The telescope status display screen is shown in Figure 19.

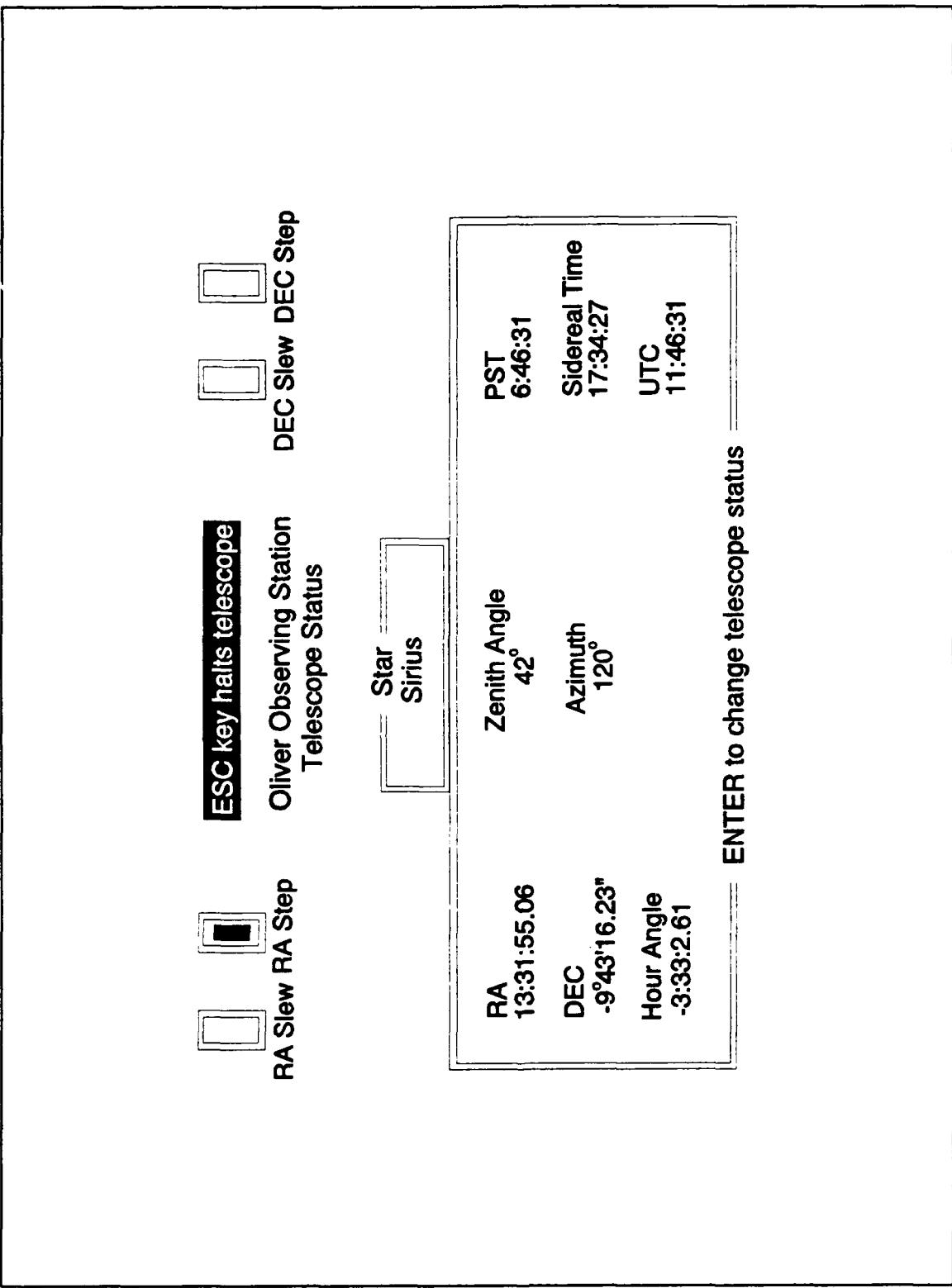


Figure 19 - The telescope status display screen with example data

VIII. PROGRAM VALIDATION

To the date of this writing, the computer program written had not been tested on the telescope due to lack of funds and the time required to connect the associated hardware. Extensive testing of each subroutine that affected telescope movement was accomplished in simulation.

A. TRACKING VALIDATION

To ensure accuracy of the program's tracking commands and related coordinate systems, several simplified simulations were performed. Initially, the latitude of the observing site was set to zero degrees and a star entered with a zero declination. Corrections for precession, nutation, refraction and aberration were not included. This ensured that, with properly defined coordinate systems, the star would have a zenith angle pointing directly East or West at any time. The simulation was run and meridian passage of a simulated star was observed. It was verified that at meridian passage, the local hour angle became zero and switched from a negative to a positive sign. It was also observed that the zenith angle shifted from East to West.

In the next validation simulation, the observing site's latitude was changed to the correct latitude at the Oliver Observing Station. The above simulation was repeated and

meridian passage of the simulated body observed. A target star from the Bright Star listing of the Astronomical Almanac, Reference 1, was then entered and its calculated azimuth and elevation verified by use of a U.S. Navy Rude Starfinder. This process was repeated in each quadrant of right ascension and for both positive and negative declination while varying the system clock to allow for each target star to be in view.

When the system clock was restored to its correct time and the target star was changed to different quadrants, the appropriate error messages were observed when the star was not in view.

B. CELESTIAL COORDINATE CORRECTION VALIDATIONS

The correction of celestial coordinates was verified by allowing only a single correction to be applied at a time. The correction under scrutiny was observed to be within the expected limits set in Table 1. Precession, nutation, aberration and refraction corrections for a given target star were calculated by hand and the results compared to the computer solution. The direction of each correction was verified at concurrently. For example, refraction of light by the Earth's atmosphere always causes the apparent position of a star to be shifted toward the zenith. This was verified in all simulations run.

Corrections to celestial coordinates must take place in the proper order, i.e. precession, nutation, aberration and

refraction. Simulations were also run to verify that the corrections were taking place in the proper sequence and that the only changes to the celestial coordinates came from those sources.

C. MOTOR COMMAND VALIDATION

Motor commands could not be fully validated since the computer was not connected to the motor drivers or controllers. Simulations to verify motor command logic were completed. The telescope status display screen also displays the motor states for all four motors. These displays were used to determine when motors were turned on or off during representative actions. Slew, guide and track motor commands were tested individually and together to ensure that the sequence of commanded motor actions was correct. Following hardware installation, a check should be made of signal output to each motor controller or driver prior to allowing telescope movement. Limit switch actions were also validated by simulation.

IX. CONCLUSIONS AND RECOMMENDATIONS

A. CONCLUSIONS

It has been shown that an automatic telescope control system is both practical and relatively inexpensive. Additional hardware requirements for the system discussed would cost approximately eight thousand dollars as of the time of this writing. While a personal computer based on an Intel 80486 microprocessor was used for this design, it was determined experimentally that the program will run in real time on an Intel 80286-based system, as long as an Intel 80287 mathematics co-processor is installed. This variation would require a dedicated computer, because multitasking on an 80286-based computer is not practical with available operating systems or data bus speeds.

B. RECOMMENDATIONS FOR ADDITIONAL WORK

There remains additional work on the MIRA telescope system before fully automatic operation is possible. Hardware updates yet required include modification of installed optical encoders, replacement of the slew motor drivers and computer control of the focusing step motor.

Completion of computer program subroutines to reflect the types of motor drivers eventually purchased will be necessary. Additionally, a warm-room computer monitor and keyboard as

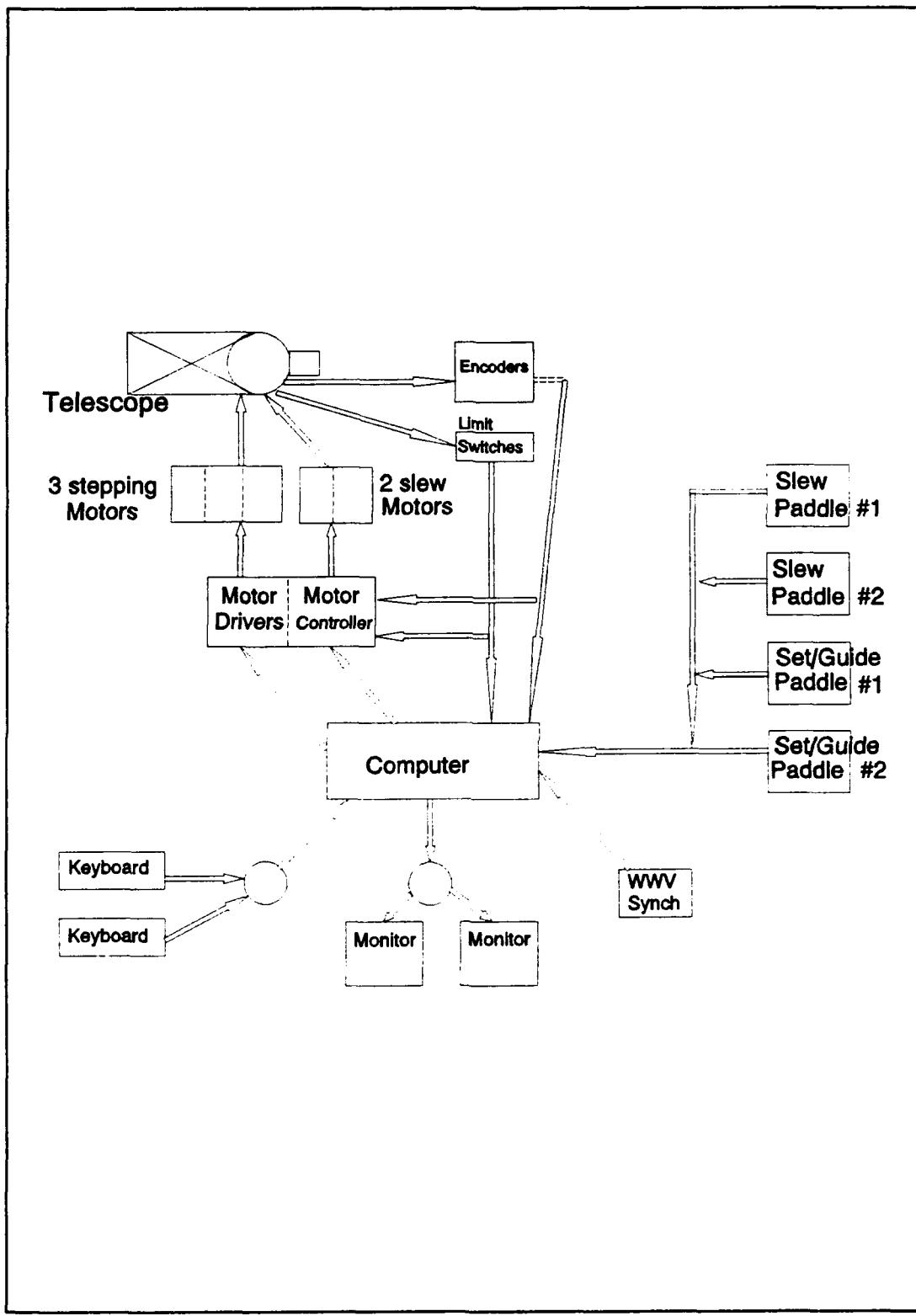


Figure 20 - Eventual hardware configuration

well as slew and guide/track paddles will need to be installed. This work was not completed because of lack of funds. Figure 20 shows the desired hardware configuration upon completion. The pointing model, which represents the changing telescope structure in software, must also be included. Testing of the completed system will require extensive simulation in both hard- and software prior to actual telescope movement. Before the telescope is first moved by computer control, a rigorous testing plan must be developed.

Automatic environment sensing of temperature and pressure is required for long duration tracking of a single body because of refractive effects. Less than one-tenth of an arc second resolution may be obtained under the current model due to manual input of time-varying parameters.

Optimization of slewing movements may be accomplished by a "traveling salesman" approach to minimize the required slew times or accelerations.

APPENDIX A MIRACTRL.EXE

A. MIRACTRL.BAS LISTING

```
'*****  
'**      Name:      MIRACTRL      **  
'**      Type:      Program      **  
'**      Module:    MIRACTRL.BAS  **  
'**      Language:   Quickbasic 4.50 **  
'**      Author:    David P. Wood **  
'*****  
'  
'Control Core for MIRA Telescope.  
'  
'Usage:          No Command Line Parameters  
' .MAK File:     (none)  
' Parameters:    (none)  
' Variables:     w*Text$()    Text for the error windows  
'                      Star$()    A 100-element matrix of desired  
stars for  
'                      viewing  
'                      RA$()     A 100-element matrix of right  
ascensions  
'                      for the stars listed in Star$()  
'                      DEC$()     A 100-element matrix of  
declinations  
'                      for the stars listed in Star$()  
'                      Epoch$()   A 100-element matrix of epochs for  
the  
'                      RA's and DEC's of the stars listed  
in Star$()  
'                      EQofEQ!()   The Equation of Equinox  
corrections to  
'                      sidereal time for successive dates  
(due  
'                      to corrections for the Earth's  
nutation  
'                      EQofEQUIN! The Equation of Equinox correction  
for the  
'                      current time  
'                      w.*       Windows Type variables, defines  
window  
'                      min%     appearance  
'                      The current minute of time
```

```

'          hour%      The current hour of time
'          day%       The current day
'          month%     The current month
'          year%      The current year
'          LocalTime$  A string representing the local
time
'          SiderealTime$ A string representing the local
'                           apparent sidereal time
'          TIME$       The system clock time string, set
to correspond
'                           to Universal Coordinated time
(UTC)
'          kee%       A unique code for any key struck
on the keyboard
'                           returned by either KeyCode% or
KeyCodeNoWait%
'                           Functions
'          Character$  A string representing a character
returned
'                           from the keyboard
'          Abort%      A variable assigned to a current
keyboard buffer
'                           input used to determine if the
user wishes to
'                           quit the program
'          Quit%       A variable assigned to a current
keyboard buffer
'                           input used to determine if the
Abort% is TRUE
'

'Define constants
CONST ENTER = 13
CONST TABHIT = 9
CONST BACKSPACE = 8
CONST ESCAPE = 27
CONST SHIFTTAB = 3840
CONST RIGHTARROW = 19712
CONST LEFTARROW = 19200
CONST UPARROW = 18432
CONST DOWNARROW = 20480
CONST NUMRIGHTARROW = 54
CONST NUMLEFTARROW = 52
CONST NUMUPARROW = 56
CONST NUMDOWNARROW = 50
CONST one = 49
CONST three = 51
CONST seven = 55
CONST nine = 57
CONST F1 = 15104
CONST F2 = 15360

```

```

CONST F3 = 15616
CONST F4 = 15872
CONST F9 = 17152
CONST F10 = 17408
CONST F11 = -31488
CONST F12 = -31232
CONST DELETE = 21248
CONST FALSE = 0
CONST TRUE = NOT FALSE
CONST PGDOWN = 20736
CONST PGUP = 18688
CONST INSERT = 20992
CONST HOME = 18176
CONST ENDFILE = 20224
CONST PLUS = 1
CONST MINUS = -1
CONST pi = 3.14159265359#

```

```

'Define color constants
CONST BLACK = 0
CONST BLUE = 1
CONST GREEN = 2
CONST CYAN = 3
CONST RED = 4
CONST MAGENTA = 5
CONST BROWN = 6
CONST WHITE = 7
CONST BRIGHT = 8
CONST BLINK = 16
CONST YELLOW = BROWN + BRIGHT

```

```

TYPE RegType
    ax      AS INTEGER
    bx      AS INTEGER
    cx      AS INTEGER
    dx      AS INTEGER
    Bp     AS INTEGER
    si      AS INTEGER
    di      AS INTEGER
    flags   AS INTEGER
END TYPE
TYPE RegTypeX
    ax      AS INTEGER
    bx      AS INTEGER
    cx      AS INTEGER
    dx      AS INTEGER
    Bp     AS INTEGER
    si      AS INTEGER
    di      AS INTEGER
    flags   AS INTEGER
    ds      AS INTEGER

```

```

        es          AS INTEGER
END TYPE

TYPE WindowsType
action      AS INTEGER
edgeline    AS INTEGER
row         AS INTEGER
col         AS INTEGER
fgdEdge     AS INTEGER
bgdEdge     AS INTEGER
fgdBody     AS INTEGER
bgdBody     AS INTEGER
fgdHighlight AS INTEGER
bgdHighlight AS INTEGER
fgdTitle    AS INTEGER
bgdTitle    AS INTEGER
fgdPrompt   AS INTEGER
bgdPrompt   AS INTEGER
returnCode  AS INTEGER
END TYPE

'Functions
DECLARE FUNCTION KeyCodeNoWait% (Character$)
DECLARE FUNCTION KeyCode% (Character$)

'Subprograms
DECLARE SUB windows (w AS WindowsType, wText$(), wTitle$,
wPrompt$)
DECLARE SUB WindowsPop ()
DECLARE SUB VideoState (mode%, columns%, page%)
DECLARE SUB InitialDisplay ()
DECLARE SUB StarDisplay ()
DECLARE SUB StarDataEntry (Star$(), RA$(), DEC$(), EPOCH$(),
NewFile%, w1 AS WindowsType, w1Text$(), w1Title$, w1Prompt$)
DECLARE SUB ChooseStar (Star$(), RA$(), DEC$(), EPOCH$(),
NewFile%, w1 AS WindowsType, w1Text$(), w1Title$, w1Prompt$,
currentStar%)
DECLARE SUB TeleStatusDisplay (ActiveStar$, Simulation%)
DECLARE SUB DayLightSaving (DayLightSavings%, hour%, day%,
month%, year%)
DECLARE SUB SiderealTimeCalc (SiderealTime$, min%, hour%,
day%, month%, year%, EQofEQUIN!, GMST0hrSEC#, JulianDate#,
dayFraction#, DaysPassed%)
DECLARE SUB EquationEquinoxes (EQofEQ(), EQofEQUIN!)
DECLARE SUB Interrupt (intnum%, inreg AS RegType, outreg AS
RegType)
DECLARE SUB ShowTime (min%, hour%, day%, month%, year%,
DayLightSavings%, SiderealTime$)
DECLARE SUB StarData (Star$(), RA$(), DEC$(), EPOCH$(),
EditDecision%, NewFile%)

```

```

DECLARE SUB Refraction (Pressure!, Temperature!, Wavelength!,
GMST0hrSEC#, HourAngle#, GHAaries#, RAofStar#, DECoFStar#,
RARefractCor#, DecRefractCor#, RAofMIRA#)
DECLARE SUB Atmospheric (Pressure!, Temperature!,
Wavelength!)
DECLARE SUB TeleAngles (Zenith%, Azimuth%, GHAaries#,
RAofStar#, DECoFStar#, CelCoord#(), DCM#(), TerraRec#())
DECLARE SUB Precession (RA$(()), DEC$(()), EPOCH$(()),
currentStar%, JulianDate#, EpochError%, RAofStar#, DECoFStar#,
RADecError%)
DECLARE SUB Aberration (RAofStar#, DECoFStar#,
RAAberrationCor#, DecAberrationCor#, dayFraction|,
DaysPassed%)
DECLARE SUB SlewCommands (Simulation%, TeleRA#, TeleDEC|,
RAofStar#, DECoFStar#)
DECLARE SUB TrackCommands ()
DECLARE SUB SetCommands (Simulation%, TeleRA#, TeleDEC|,
RAofStar#, DECoFStar#)
DECLARE SUB TeleAlign (GMST0hrSEC#, GHAaries#, TeleRA|,
TeleDEC#)
DECLARE SUB Nutation (RAofStar#, DECoFStar#, JulianDate|,
N#(), S1#(), S2#())
'
'Data Structures
DIM w1 AS WindowsType
DIM w2 AS WindowsType
DIM w3 AS WindowsType
DIM w4 AS WindowsType
DIM w5 AS WindowsType
DIM w6 AS WindowsType
DIM w7 AS WindowsType
DIM w8 AS WindowsType
DIM w9 AS WindowsType
DIM w10 AS WindowsType
DIM w11 AS WindowsType
DIM w12 AS WindowsType
'
'Arrays
DIM w1Text$(1 TO 1)
DIM w2Text$(1 TO 1)
DIM w3Text$(1 TO 1)
DIM w4Text$(1 TO 1)
DIM w5Text$(1 TO 7)
DIM w6Text$(1 TO 9)
DIM w7Text$(1 TO 1)
DIM w8Text$(1 TO 2)
DIM w9Text$(1 TO 3)
DIM w10Text$(1 TO 2)
DIM w11Text$(1 TO 13)
DIM w12Text$(1 TO 5)
DIM Star$(1 TO 100)

```

```

DIM RA$(1 TO 100)
DIM DEC$(1 TO 100)
DIM EPOCH$(1 TO 100)
DIM EQofEQ!(1 TO 2)
DIM CelCoord#(1 TO 3)
DIM DCM#(1 TO 3, 1 TO 3)
DIM TerraCoord#(1 TO 3)
DIM TerraRec#(1 TO 3)
DIM S1#(1 TO 3)
DIM S2#(1 TO 3)
DIM N#(1 TO 3, 1 TO 3)

'Define error windows
'

'The quit program confirmation window
w1.action = 1
w1.edgeline = 1
w1.row = 5
w1.col = 5
w1.fgdEdge = CYAN + BRIGHT
w1.bgdEdge = RED
w1.fgdBody = BRIGHT + WHITE
w1.bgdBody = RED
w1.fgdHighlight = 0
w1.bgdHighlight = 0
w1.fgdTitle = CYAN
w1.bgdTitle = RED
w1.fgdPrompt = YELLOW
w1.bgdPrompt = RED
w1Title$ = ""
w1Text$(1) = " Quit Program? (Y or N) "
w1Prompt$ = ""

'The file save error window
w2.action = 1
w2.edgeline = 1
w2.row = 5
w2.col = 5
w2.fgdEdge = CYAN + BRIGHT
w2.bgdEdge = RED
w2.fgdBody = BRIGHT + WHITE
w2.bgdBody = RED
w2.fgdHighlight = 0
w2.bgdHighlight = 0
w2.fgdTitle = CYAN
w2.bgdTitle = RED
w2.fgdPrompt = YELLOW
w2.bgdPrompt = RED
w2Title$ = ""
w2Text$(1) = " Save error "
w2Prompt$ = "Press any key to continue"

```

```

'
'The file-not-found error window
w3.action = 1
w3.edgeline = 1
w3.row = 5
w3.col = 5
w3.fgdEdge = CYAN + BRIGHT
w3.bgdEdge = RED
w3.fgdBody = BRIGHT + WHITE
w3.bgdBody = RED
w3.fgdHighlight = 0
w3.bgdHighlight = 0
w3.fgdTitle = CYAN
w3.bgdTitle = RED
w3.fgdPrompt = YELLOW
w3.bgdPrompt = RED
w3Title$ = ""
w3Text$(1) = " Error - File not found or path incorrect "
w3Prompt$ = "Press any key to continue"
'

'The Equinox correction error window
w4.action = 1
w4.edgeline = 1
w4.row = 5
w4.col = 5
w4.fgdEdge = CYAN + BRIGHT
w4.bgdEdge = RED
w4.fgdBody = BRIGHT + WHITE
w4.bgdBody = RED
w4.fgdHighlight = 0
w4.bgdHighlight = 0
w4.fgdTitle = CYAN
w4.bgdTitle = RED
w4.fgdPrompt = YELLOW
w4.bgdPrompt = RED
w4Title$ = ""
w4Text$(1) = " Error - Equinox corrections must be a NUMBER
in seconds "
w4Prompt$ = "Press any key to continue"
'

'The slew confirmation window
w5.action = 1
w5.edgeline = 1
w5.row = 5
w5.col = 5
w5.fgdEdge = CYAN + BRIGHT
w5.bgdEdge = RED
w5.fgdBody = BRIGHT + WHITE
w5.bgdBody = RED
w5.fgdHighlight = 0
w5.bgdHighlight = 0

```

```

w5.fgdTitle = WHITE + BLINK + BRIGHT
w5.bgdTitle = RED
w5.fgdPrompt = YELLOW
w5.bgdPrompt = RED
w5Title$ = " W A R N I N G "
w5Text$(1) = ""
w5Text$(2) = " This action will slew the telescope. "
w5Text$(3) = " The initial zenith angle will be " + Zenith$ + " deg. "
w5Text$(4) = " Any key other than F1 exits this procedure. "
w5Text$(5) = " The ESC key will halt telescope movement at "
w5Text$(6) = " any time."
w5Text$(7) = ""
w5Prompt$ = "Press the F1 key to slew"
'

'The Change telescope status window
w6.action = 1
w6.edgeline = 1
w6.row = 5
w6.col = 3
w6.fgdEdge = CYAN + BRIGHT
w6.bgdEdge = BLUE
w6.fgdBody = BRIGHT + WHITE
w6.bgdBody = BLUE
w6.fgdHighlight = 0
w6.bgdHighlight = 0
w6.fgdTitle = WHITE + BRIGHT
w6.bgdTitle = BLUE
w6.fgdPrompt = YELLOW
w6.bgdPrompt = BLUE
w6Title$ = " Change of Telescope Status "
w6Text$(1) = ""
w6Text$(2) = " F1: Edit the current list "
w6Text$(3) = " F2: Choose a star from the current list "
w6Text$(4) = " F3: Change to a new data file "
w6Text$(5) = " F4: Change atmospheric parameters "
w6Text$(6) = " F5: Park the telescope "
w6Text$(7) = " ESC: Exit the program and halt the telescope
"
w6Text$(8) = " Any other key aborts the change "
w6Text$(9) = ""
w6Prompt$ = "Choose an option"
'

'The RA/DEC format error window
w7.action = 1
w7.edgeline = 1
w7.row = 5
w7.col = 5
w7.fgdEdge = CYAN + BRIGHT
w7.bgdEdge = RED
w7.fgdBody = BRIGHT + WHITE

```

```

w7.bgdBody = RED
w7.fgdHighlight = 0
w7.bgdHighlight = 0
w7.fgdTitle = CYAN
w7.bgdTitle = RED
w7.fgdPrompt = YELLOW
w7.bgdPrompt = RED
w7Title$ = ""
w7Text$(1) = " Error - RA or DEC is formatted incorrectly for
this star "
w7Prompt$ = "Press any key to continue"
'

'The slew error window
w8.action = 1
w8.edgeline = 1
w8.row = 5
w8.col = 10
w8.fgdEdge = CYAN + BRIGHT
w8.bgdEdge = RED
w8.fgdBody = BRIGHT + WHITE
w8.bgdBody = RED
w8.fgdHighlight = 0
w8.bgdHighlight = 0
w8.fgdTitle = CYAN + BLINK
w8.bgdTitle = RED
w8.fgdPrompt = YELLOW
w8.bgdPrompt = RED
w8Title$ = " NOTICE "
w8Text$(1) = "      This star is not in view or has a zenith
angle "
w8Text$(2) = " greater than 75 degrees. Please choose
another star."
w8Prompt$ = "Press any key to continue"
'

'The zenith angle warning window
w9.action = 0
w9.edgeline = 1
w9.row = 7
w9.col = 57
w9.fgdEdge = CYAN + BRIGHT
w9.bgdEdge = YELLOW
w9.fgdBody = BRIGHT + WHITE
w9.bgdBody = YELLOW
w9.fgdHighlight = 0
w9.bgdHighlight = 0
w9.fgdTitle = CYAN + BLINK
w9.bgdTitle = YELLOW
w9.fgdPrompt = YELLOW
w9.bgdPrompt = YELLOW
w9Title$ = " NOTICE "
w9Text$(1) = " The zenith "

```

```

w9Text$(2) = " angle is"
w9Text$(3) = STR$(Zenith%) + " degrees "
w9Prompt$ = ""
'
'The zenith angle halt movement window
w10.action = 1
w10.edgeline = 1
w10.row = 5
w10.col = 5
w10.fgdEdge = CYAN + BRIGHT
w10.bgdEdge = RED
w10.fgdBody = BRIGHT + WHITE
w10.bgdBody = RED
w10.fgdHighlight = 0
w10.bgdHighlight = 0
w10.fgdTitle = CYAN + BLINK
w10.bgdTitle = RED
w10.fgdPrompt = YELLOW
w10.bgdPrompt = RED
w10Title$ = " NOTICE "
w10Text$(1) = " Telescope movement has been halted. "
w10Text$(2) = " The zenith angle has exceeded 75 degrees. "
w10Prompt$ = "Press any key to continue"
'
'The Status help window
w11.action = 1
w11.edgeline = 1
w11.row = 5
w11.col = 3
w11.fgdEdge = CYAN + BRIGHT
w11.bgdEdge = BLUE
w11.fgdBody = BRIGHT + WHITE
w11.bgdBody = BLUE
w11.fgdHighlight = 0
w11.bgdHighlight = 0
w11.fgdTitle = WHITE + BRIGHT
w11.bgdTitle = BLUE
w11.fgdPrompt = YELLOW
w11.bgdPrompt = BLUE
w11Title$ = " Telescope Status Help"
w11Text$(1) = " Slow step:                               Fast step:
"
w11Text$(2) = "                                         "
w11Text$(3) = " up arrow      = Increase DEC          8 = Increase
DEC "
w11Text$(4) = " down arrow     = Decrease DEC         2 = Decrease
DEC "
w11Text$(5) = " left arrow     = Increase RA          4 = Increase
RA "
w11Text$(6) = " right arrow    = Decrease RA         6 = Decrease
RA "

```

```

w11Text$(7) = " "
w11Text$(8) = " Slew: "
w11Text$(9) = " "
w11Text$(10) = " 1 = Increase DEC "
w11Text$(11) = " 3 = Decrease DEC "
w11Text$(12) = " 7 = Increase RA "
w11Text$(13) = " 9 = Decrease RA "
w11Prompt$ = " Press any key to exit help "
'
'The CTS-10 error window
w12.action = 1
w12.edgeline = 1
w12.row = 5
w12.col = 5
w12.fgdEdge = CYAN + BRIGHT
w12.bgdEdge = RED
w12.fgdBody = BRIGHT + WHITE
w12.bgdBody = RED
w12.fgdHighlight = 0
w12.bgdHighlight = 0
w12.fgdTitle = CYAN + BLINK
w12.bgdTitle = RED
w12.fgdPrompt = YELLOW
w12.bgdPrompt = RED
w12Title$ = "Error"
w12Text$(1) = " The CTS-10 WWV Time Syncronization board is
not "
w12Text$(2) = " responding. Take one of the following
actions: "
w12Text$(3) = " "
w12Text$(4) = " a. Check the CTS-10 onboard battery "
w12Text$(5) = " b. Manually update the system time to
UTC "
w12Prompt$ = "Press any key to exit"
'
'Initialize the display
SCREEN 0, , 0, 0
WIDTH 80
InitialDisplay
'
*****  

'** Temporary Debugging Statement **  

DO  

LOOP UNTIL INKEY$ <> ""  

*****  

'  

'Initialize the rest of the system (describe)
*****  

*****  

'**  


```

```

*** More Work Needed
**
'POKE (address of external boards), 0
**
'POKE (address of Key paddle buffer), 0
**
***** *****
*****
'
'Check the CTS-10 WWV Time Synch board for response
year% = VAL(RIGHT$(DATE$, 4))
IF year% < 1992 THEN windows w12, w12Text$, w12Title$,
w12Prompt$
SYSTEM
'

'Perform a one-time calculation of the Equation of Equinox
correction to
'sidereal time
ON ERROR GOTO EQofEQError:
    EquationEquinoxes EQofEQ(), EQofEQUIN!
ON ERROR GOTO 0
'

'Determine current time and date information for use in
DayLightSaving,
'SiderealTimeCalc SUBRoutines and Status DO LOOP, below
min% = VAL(MID$(TIME$, 4, 2))
hour% = VAL(LEFT$(TIME$, 2))
day% = VAL(MID$(DATE$, 4, 2))
month% = VAL(LEFT$(DATE$, 2))
year% = VAL(RIGHT$(DATE$, 4))
'

'Check whether Daylight Savings time applies to local time
DayLightSaving DayLightSavings%, hour%, day%, month%, year%
'

ChooseDecision% = TRUE
AtmosChange% = FALSE
DataEntry% = TRUE
RADecError% = FALSE
firstloop% = TRUE
WindowOn% = FALSE
Simulation% = FALSE
DatafileOld$ = "STARFILE.DAT"
Pressure! = 633! 'mmHg @ 5000 ft altitude default
Temperature! = 10! 'C @ 5000 ft altitude default
Wavelength! = .5 'microns @ 5000 ft altitude default
'

'Allow the user to change the default atmospheric conditions
used to calculate refraction
Atmospheric Pressure!, Temperature!, Wavelength!
'

'Perform sidereal time calculations

```

```

SiderealTimeCalc SiderealTime$, min%, hour%, day%, month%,
year%, EQofEQUIN!, GMST0hrSEC#, JulianDate#, dayFraction#,
DaysPassed%
'
'Align telescope in software

TeleAlign GMST0hrSEC#, GHAaries#, TeleRA#, TeleDEC#
'

DO
    'Perform sidereal time calculations
    SiderealTimeCalc SiderealTime$, min%, hour%, day%,
month%, year%, EQofEQUIN!, GMST0hrSEC#, JulianDate#,
dayFraction#, DaysPassed%
'
    IF DataEntry% = TRUE THEN
        DataEntry% = FALSE
        'Accept star data entry from disk or by hand
input
        COLOR WHITE, BLUE, BLUE
        FOR r% = 14 TO 21
            LOCATE r%, 6, 0
            PRINT SPACE$(68)
        NEXT r%
        EditDecision% = FALSE
        NewFile% = TRUE
        DO
            COLOR WHITE, BLUE, BLUE
            LOCATE 17, 10, 0
            PRINT SPACE$(45)
            LOCATE 17, 10, 0
            PRINT "Enter star data by: (F1) Disk"
            LOCATE 18, 31, 0
            PRINT "(F2) Keyboard"
            kee% = KeyCode%(Character$)
            SELECT CASE kee%
                CASE 15104 'F1
                    BadPathType% = FALSE
                    NewFile% = FALSE
                    FOR i% = 1 TO 2
                        LOCATE 16 +
                            PRINT
i%, 25, 0
                    SPACE$(34)

the path and file name: ", Datafile$
THEN
                INPUT "Please enter
IF LEN(Datafile$) = 0
                LOCATE 17, 47, 1
                PRINT DatafileOld$
                LOCATE 17, 10, 1

```

```

DatafileOld$                               Datafile$ =
                                         ELSE
                                         DatafileOld$ =
= Datafile$                               END IF
                                         IF LEN(Datafile$) > 0
                                         ON ERROR GOTO
THEN                                         ON ERROR GOTO
                                         OPEN
BadPath:                                         ON ERROR GOTO
                                         OPEN
Datafile$ FOR INPUT AS #1
                                         ON ERROR GOTO
0                                         IF
                                         i% =
                                         DO
BadPathType% = FALSE THEN
                                         i% =
1                                         DO
WHILE NOT EOF(1)
                                         LOOP
                                         IF i%
                                         INPUT #1, Star$(i%), RA$(i%), DEC$(i%), EPOCH$(i%)
                                         i% = i% + 1
                                         LOOP
                                         IF i%
< 100 THEN
                                         FOR h% = i% TO 100
                                         Star$(h%) = ""
                                         RA$(h%) = ""
                                         DEC$(h%) = ""
                                         EPOCH$(h%) = ""
                                         NEXT h%
                                         END IF
                                         RESET
                                         LOCATE
17, 11, 0                               PRINT
                                         LOCATE
SPACE$(63)                               PRINT
                                         LOCATE
17, 10, 0                               PRINT
                                         LOCATE
"Do you wish to edit the file"          LOCATE
18, 10, 0

```

```

"prior to proceeding? (Y or N) " PRINT
WHITE, BLUE, BLUE COLOR
18, 40, 0 LOCATE
"N" PRINT
LOCATE
18, 40, 1 kee%
= KeyCode%(Character$) IF
UCASE$(Character$) = "Y" THEN EditDecision% = TRUE Stops%
= TRUE ELSE
action, return to beginning of loop END IF
ELSE
'return to beginning of loop
END IF
CASE 15360 ' F2
EditDecision% = TRUE
NewFile% = TRUE
Stops% = TRUE
CASE ELSE
Stops% = FALSE
END SELECT
LOOP UNTIL Stops% = TRUE
END IF
'
IF EditDecision% = TRUE THEN
EditDecision% = FALSE
'Change the screen to display readiness for
star data entry
StarDisplay
'Accept star data entry
StarDataEntry Star$(), RA$(), DEC$(),
EPOCH$(), NewFile%, w1, w1Text$(), w1Title$, w1Prompt$
'
'Save data to file
IF NewFile% = TRUE THEN
NewFile% = FALSE
COLOR WHITE, BLUE, BLUE
FOR r% = 14 TO 21
LOCATE r%, 6, 0
PRINT SPACE$(68)
NEXT r%

```

```

        COLOR BLUE, BLACK, BLUE
        LOCATE 23, 5, 0
        PRINT STRING$(30, 177)
        Datafile$ = "StarFile.DAT"
        COLOR WHITE, BLUE, BLUE
        LOCATE 15, 37, 0
        PRINT Datafile$
        LOCATE 15, 10, 0
        INPUT "Please enter the filename: ",  

DataFile2$  

= DataFile2$  

#1  

DEC$(i%), EPOCH$(i%)  

        ON ERROR GOTO NewFileError:  

            OPEN Datafile$ FOR OUTPUT AS  

        ON ERROR GOTO 0
        FOR i% = 1 TO UBOUND(Star$)
            WRITE #1, Star$(i%), RA$(i%),  

        NEXT i%
        RESET  

        ELSE  

            'If an old file, store to disk
            COLOR WHITE, BLUE, BLUE
            FOR r% = 14 TO 21
                LOCATE r%, 6, 0
                PRINT SPACE$(68)
            NEXT r%
            COLOR BLUE, BLACK, BLUE
            LOCATE 23, 5, 0
            PRINT STRING$(30, 177)
            COLOR WHITE, BLUE, BLUE
            LOCATE 15, 37, 0
            PRINT Datafile$
            LOCATE 15, 10, 0
            INPUT "Please enter the filename: ",  

DataFile2$  

= DataFile2$  

#1  

        ON ERROR GOTO OldFileError:  

            OPEN Datafile$ FOR OUTPUT AS  

        ON ERROR GOTO 0
        IF FileError% = FALSE THEN
            FOR i% = 1 TO 100

```

```

                WRITE #1, Star$(i%),
RA$(i%), DEC$(i%), EPOCH$(i%)
                                NEXT i%
                END IF
                RESET
                FOR r% = 14 TO 21
                        LOCATE r%, 6, 0
                        PRINT SPACE$(68)
                NEXT r%
        END IF
END IF
'
'Allow the user to update the atmospheric parameters
on which refraction is calculated
IF AtmosChange% = TRUE THEN
    AtmosChange% = FALSE
    Atmospheric Pressure!, Temperature!,
Wavelength!
END IF
'
'Allow the user to choose the active star on which the
telescope will track
IF ChooseDecision% = TRUE THEN
    ChooseDecision% = FALSE
    ChooseStar Star$(), RA$(), DEC$(), EPOCH$(),
NewFile%, w1, w1Text$, w1Title$, w1Prompt$, currentStar%
    ActiveStar$ = Star$(currentStar%)
    'Calculate Precession and Nutation effects for
the day
    Precession RA$(), DEC$(), EPOCH$(),
currentStar%, JulianDate#, EpochError%, RAofStar#, DECoFStar#,
RADecError%
'
    'RA and DEC displayed are corrected for
precession only
    RAofStarDisplay# = RAofStar#
    DECoFStarDisplay# = DECoFStar#
'
    Nutation RAofStar#, DECoFStar#, JulianDate#,
N#(), S1#(), S2#()
'
    Aberration RAofStar#, DECoFStar#,
RAAberrationCor#, DecAberrationCor#, dayFraction#, DaysPassed%
    RAofStar# = RAofStar# + RAAberrationCor#
    DECoFStar# = DECoFStar# + DecAberrationCor#
'
    Refraction Pressure!, Temperature!,
Wavelength!, GMST0hrSEC#, HourAngle#, GHAaries#, RAofStar#,
DECoFStar#, RARefractCor#, DecRefractCor#, RAofMIRA#
    RAofStar# = RAofStar# + RARefractCor#
    DECoFStar# = DECoFStar# + DecRefractCor#

```

```

      RARefractCorOld# = RARefractCor#
      DecRefractCorOld# = DecRefractCor#
      DO WHILE RAofStar# >= (2# * pi)
          RAofStar# = RAofStar# - (2# * pi)
      LOOP
      DO WHILE RAofStar# < 0
          RAofStar# = RAofStar# + (2# * pi)
      LOOP
      DO WHILE DECoFStar# > (pi / 2#)
          DECoFStar# = pi - DECoFStar#
      LOOP
      DO WHILE DECoFStar# < -(pi / 2#)
          DECoFStar# = -(pi + DECoFStar#)
      LOOP
      TeleAngles Zenith%, Azimuth%, GHAaries#,
RAofStar#, DECoFStar#, CelCoord#(), DCM#(), TerraRec#()
Zenith$ = STR$(Zenith%)
w5Text$(3) = " The initial zenith angle will
be " + Zenith$ + " deg. "
windows w5, w5Text$, w5Title$, w5Prompt$
ELSE
    w5.returnCode% = 15104' F1
END IF
'
'If the user enters F10 vice F1, simulation of the
control process occurs
IF w5.returnCode = 17408 THEN ' F10
    w5.returnCode = 15104' F1
    Simulation% = TRUE
END IF
SELECT CASE w5.returnCode
    CASE 15104' F1
        WindowsPop
        TimeSec% = VAL(RIGHT$(TIME$, 2))
        '
        'Display telescope status screen
        TeleStatusDisplay ActiveStar$,
Simulation%
        '
        Quit% = FALSE
        FirstPass% = TRUE
        DO
            'Calculate and display local,
sidereal and UTC times
            'and calculate Refraction
corrections
            IF TimeSec% = 59 AND
VAL(RIGHT$(TIME$, 2)) = 0 THEN
                ^
                ShowTime min%, hour%,
day%, month%, year%, DayLightSavings%, SiderealTime$

```

```

TimeSec%      =
VAL(RIGHT$(TIME$, 2))          RefractionPressure!,  

Temperature!, Wavelength!, GMST0hrSEC#, HourAngle#, GHAaries#,  

RAofStar#,   DECoStar#,   RARefractCor#,   DecRefractCor#,  

RAofMIRA#  

RAofStar# = RAofStar#  

- RARefractCorOld#  

DECoStar# = DecRefractCorOld#  

RAofStar# = RAofStar#  

+ RARefractCor#  

DECoStar# = DecRefractCor#  

RAofStar# = RAofStar#  

RARefractCor#  

DecRefractCor#  

DecRefractCorOld# =  

DecRefractCorOld# =  

ELSEIF TimeSec% <  

VAL(RIGHT$(TIME$, 2)) THEN  

ShowTime min%, hour%,  

day%, month%, year%, DayLightSavings%, SiderealTime$  

TimeSec% =  

VAL(RIGHT$(TIME$, 2))          RefractionPressure!,  

Temperature!, Wavelength!, GMST0hrSEC#, HourAngle#, GHAaries#,  

RAofStar#,   DECoStar#,   RARefractCor#,   DecRefractCor#,  

RAofMIRA#  

RAofStar# = RAofStar#  

- RARefractCorOld#  

DECoStar# = DecRefractCorOld#  

RAofStar# = RAofStar#  

+ RARefractCor#  

DECoStar# = DecRefractCor#  

RAofStar# = RAofStar#  

RARefractCor#  

DecRefractCor#  

DecRefractCorOld# =  

DecRefractCorOld# =  

END IF  

DO WHILE RAofStar# >= (2# *  

pi)  

RAofStar# = RAofStar#  

- (2# * pi)  

LOOP  

DO WHILE RAofStar# < 0  

RAofStar# = RAofStar#  

+ (2# * pi)  

LOOP

```

```

DO WHILE DECofStar# > (pi /
2#)
DECofStar#
LOOP
DO WHILE DECofStar# < -(pi /
2#)
DECofStar# = -(pi +
DECofStar#)
LOOP
'
'Display current RA, DEC and
Hour Angle of the star
COLOR WHITE, BLUE, BLUE
RAofStarInHours# =
RAofStarDisplay# * 24# / (2# * pi)           RAHours% =
INT(RAofStarInHours#)                         RAMinutes% =
INT((RAofStarInHours# - CDBL(RAHours%)) * 60#)   RASeconds! =
((RAofStarInHours# - CDBL(RAHours%)) * 60# - CDBL(RAMinutes%)) * 60#
LOCATE 15, 10, 0
PRINT
RTRIM$(LTRIM$(STR$(RAHours%))) ;      ":" ;
RTRIM$(LTRIM$(STR$(RAMinutes%))) ;      ":" ;
RTRIM$(LTRIM$(STR$(INT(RASeconds! * 100) / 100))); SPACE$(5)
DECinDeg# = DECofStarDisplay#
* 360# / (2# * pi)
DECinDeg! = CSNG(DECinDeg#)
DECDegrees% = INT(DECinDeg!)
DECMinutes% = INT((DECinDeg#
- CDBL(DECDegrees%)) * 60#)
DECSeconds! =
CSNG(((DECinDeg# - CDBL(DECDegrees%)) * 60#) -
CDBL(DECMinutes%)) * 60#
LOCATE 18, 10, 0
PRINT
RTRIM$(LTRIM$(STR$(DECDegrees%))) ;      CHR$(248) ;
RTRIM$(LTRIM$(STR$(DECMinutes%))) ;      ":" ;
RTRIM$(LTRIM$(STR$(INT(DECSeconds! * 100) / 100)));     "'";
SPACE$(5)
'
'Update the star's hour angle
and display
RAofStar#    'in radians
HourAngle# = RAofMIRA# -
DO WHILE HourAngle# >= (2# *
pi)

```

```

HourAngle# = (2# * pi)
HourAngle# + (2# * pi)
* 360# / (2# * pi)
HourAngleInDeg# = 360#
HourAngleInDeg# / 15#'Conversion of degrees to hours of time
STR$(FIX(HourAngleInHr#))
AND HourAngleInDeg# < 0 THEN HourAngleHours$ = "-0"
ABS((HourAngleInHr# - CDBL(FIX(HourAngleInHr#))) * 60#
STR$(FIX(HourAngleMin#))
THEN HourAngleMinutes$ = "0" + HourAngleMinutes$
HourAngleSec# = (HourAngleMin#
- CDBL(FIX(HourAngleMin#))) * 60#
FIX(HourAngleSec# * 100#)
STR$(ABS(CSNG(HourAngleSecMod%) / 100!))
LOCATE 21, 10, 0
P R I N T
LTRIM$(RTRIM$(HourAngleHours$)); ":" ;
LTRIM$(RTRIM$(HourAngleMinutes$)); ":" ;
LTRIM$(HourAngleSeconds$); SPACE$(10)
'
'Calculate and display Zenith
Angle and Azimuth Angle on status screen
TeleAngles Zenith%, Azimuth%,
GHAaries#, RAofStar#, DECoStar#, CelCoord#(), DCM#(),
TerraRec#()
LOCATE 15, 37, 0
PRINT Zenith%;
LOCATE 15, POS(0) - 1, 0
PRINT CHR$(248); SPACE$(5)
LOCATE 18, 36, 0
PRINT Azimuth%;
LOCATE 18, POS(0) - 1, 0
PRINT CHR$(248); SPACE$(5)
IF RADecError% = TRUE THEN

```

```

                    RADecError% = FALSE
                    kee% = ENTER
                    windows w7, w7Text$(),
w7Title$, w7Prompt$                               WindowsPop
firstloop% = TRUE THEN                         ELSEIF Zenith% >= 75 AND
                                                firstloop% = FALSE
                                                windows w8, w8Text$(),
w8Title$, w8Prompt$                               WindowsPop
                                                Quit% = TRUE
firstloop% = FALSE THEN                         ELSEIF Zenith% >= 75 AND
                                                windows w10,
w10Text$(), w10Title$, w10Prompt$               WindowsPop
                                                Quit% = TRUE
                                                ELSE
                                                IF Zenith% > 65 AND
WindowOn% = FALSE THEN                         WindowOn% =
                                                TRUE
                                                Zenith%
STR$(Zenith%) + " degrees "
w9Text$(), w9Title$, w9Prompt$                  windows w9,
                                                END IF
                                                IF Zenith% > 65 AND
WindowOn% = TRUE THEN                         IF Zenith% <>
                                                ZenithOld% =
                                                ZenithOld% = Zenith%
                                                WindowsPop
                                                w9Text$(3) =
                                                windows w9,
                                                END IF
                                                IF Zenith% <>
ZenithOld% AND WindowOn% = TRUE THEN          Zenith% <>
                                                ZenithOld% =
                                                Zenith%
                                                WindowsPop
                                                WindowOn% =
FALSE

```

```

        END IF
        '
        'Clear the motor state
        COLOR  WHITE,  BLUE,
        BLUE
        FOR i% = 1 TO 2
            LOCATE 3, 10
        * i%, 0
                PRINT
        SPACE$(1)
        NEXT i%
        FOR i% = 1 TO 2
            LOCATE 3, (10
        * i%) + 50, 0
                PRINT
        SPACE$(1)
        NEXT i%
        '
        'Read in Telescope
position data from encoders and operate on them
        'to define TeleRA# and
TeleDEC# in RADIANS

'*****          *** Read in encoder
data  **

'*****          RAdiff# = TeleRA# -
RAofStar#
DECofStar#
        IF FirstPass% = TRUE
THEN
        IF
ABS(RAendiff#) > .0175 OR ABS(DECendiff#) > .0175 THEN
        SlewCommands  Simulation%,  TeleRA#,  TeleDEC#,  RAofStar#,
DECofStar#
                END IF
                RAendiff#      =
                DECendiff#     =
TeleRA# - RAofStar#
TeleRA# - DECofStar#
        IF
ABS(RAendiff#) < .0175 OR ABS(DECendiff#) < .0175 THEN
        SetCommands  Simulation%,  TeleRA#,  TeleDEC#,  RAofStar#,
DECofStar#
                END IF

```

```

FirstPass% =
FALSE
END IF
TrackCommands
kee% =
KeyCodeNoWait%(Character$)
END IF
SELECT CASE kee%
CASE ENTER
    Quit% = TRUE
CASE ESCAPE
    windows w1,
    WindowsPop
    IF
w1.returnCode = 121 OR w1.returnCode = 89 THEN Abort% = TRUE
    IF Abort% =
TRUE THEN

'***** ****
System Shutdown Procedures **

go here!!      **

'***** ****
CLS
SYSTEM
END IF
CASE F1
    windows w1,
    WindowsPop
CASE LEFTARROW
    'Change RA
step multiplier to SLOW speed, INCREASE right ascension
    '
    'POKE (address
of RA step controller), X
    'Updatemotor
display box
    COLOR RED,
BLUE, BLUE
    LOCATE 3, 20,
    0
    PRINT
CHR$(222)
CASE RIGHTARROW
    'Change RA
step multiplier to SLOW speed, DECREASE right ascension
    '

```

```

        'POKE (address
of RA step controller), X
display box
BLUE, BLUE
0
PRINT
CHR$(222)

CASE NUMLEFTARROW
'Change RA
step multiplier to FAST speed, INCREASE right ascension
'
        'POKE (address
of RA step controller), X
display box
BLUE, BLUE
0
PRINT
CHR$(222)

CASE NUMRIGHTARROW
'Change RA
step multiplier to FAST speed, DECREASE right ascension
'
        'POKE (address
of RA step controller), X
display box
BLUE, BLUE
0
PRINT
CHR$(222)

CASE UPARROW
'Change DEC
step multiplier to SLOW speed, INCREASE declination
'
        'POKE (address
of DEC step controller), X
display box
BLUE, BLUE
0

```

```

PRINT
CHR$(222)
CASE DOWNARROW
'Change DEC
step multiplier to SLOW speed, DECREASE declination
'
'POKE (address
of DEC step controller), X
'Updatemotor
display box
COLOR RED,
BLUE, BLUE
LOCATE 3, 70,
0
PRINT
CHR$(222)
CASE NUMUPARROW
'Change DEC
step multiplier to FAST speed, INCREASE declination
'
'POKE (address
of DEC step controller), X
'Updatemotor
display box
COLOR RED,
BLUE, BLUE
LOCATE 3, 70,
0
PRINT
CHR$(222)
CASE NUMDOWNARROW
'Change DEC
step multiplier to FAST speed, DECREASE declination
'
'POKE (address
of DEC step controller), X
'Updatemotor
display box
COLOR RED,
BLUE, BLUE
LOCATE 3, 70,
0
PRINT
CHR$(222)
CASE seven
'Manually slew
to INCREASE right ascension
'
'Updatemotor
display box

```

```

        COLOR    RED,
BLUE, BLUE
        LOCATE 3, 10,
0
        PRINT
CHR$(222)

        'POKE (address
of RA slew controller), X
        CASE nine
        'Manually slew
        'Updatemotor
        COLOR    RED,
        LOCATE 3, 10,
0
        PRINT
CHR$(222)

        'POKE (address
of RA slew controller), X
        CASE one
        'Manually slew
        'Updatemotor
        COLOR    RED,
        LOCATE 3, 60,
0
        PRINT
CHR$(222)

        'POKE (address
of DEC slew controller), X
        CASE three
        'Manually slew
        'Updatemotor
        COLOR    RED,
        LOCATE 3, 60,
0
        PRINT
CHR$(222)

        'POKE (address
of DEC slew controller), X

```

```

        CASE ELSE
            'Turn slew
        motors OFF once no key is pressed
        '
        'POKE (address
        of RA slew controller), X
        'POKE (address
        of DEC slew controller), X
        END SELECT
        LOOP UNTIL Quit% = TRUE
        IF WindowOn% = TRUE THEN
            WindowOn% = FALSE
            WindowsPop
        END IF
        Quit% = FALSE
        windows w6, w6Text$, w6Title$,
w6Prompt$
        WindowsPop
        SELECT CASE w6.returnCode%
        CASE 15104 ' F1
            EditDecision% = TRUE
            ChooseDecision% = TRUE
            firstloop% = TRUE
        CASE 15360 ' F2
            ChooseDecision% = TRUE
            firstloop% = TRUE
        CASE 15616 ' F3
            DataEntry% = TRUE
            ChooseDecision% = TRUE
            firstloop% = TRUE
        CASE 15872 ' F4
            AtmosChange% = TRUE
            ChooseDecision% = TRUE
            firstloop% = TRUE
        CASE 16128 ' F5
        *****
        ***      Park     the
telescope    **
        *****
        CASE ESCAPE
            windows w1, w1Text$(),
w1Title$, w1Prompt$
            WindowsPop
            IF w1.returnCode = 121
OR w1.returnCode = 89 THEN Abort% = TRUE
            IF Abort% = TRUE THEN
        *****

```

```

        ' **      System
Shutdown Procedures  **
        ' **
go here!!          **

'***** *****
        CLS
SYSTEM
        END IF
        CASE ELSE
END SELECT
'
CASE ELSE
    WindowsPop
    windows w6,   w6Text$(),   w6Title$,
w6Prompt$
    WindowsPop
    SELECT CASE w6.returnValue%
        CASE 15104 ' F1
            EditDecision% = TRUE
            ChooseDecision% = TRUE
        CASE 15360 ' F2
            ChooseDecision% = TRUE
        CASE 15616 ' F3
            DataEntry% = TRUE
            ChooseDecision% = TRUE
        CASE 15872 ' F4
            AtmosChange% = TRUE
            ChooseDecision% = TRUE
        CASE 16128 ' F5

'***** *****
        ' **      Park      the
telescope      **

'***** *****
        CASE ESCAPE
            windows w1, w1Text$(),
w1Title$, w1Prompt$
            WindowsPop
            IF w1.returnValue = 121
OR w1.returnValue = 89 THEN Abort% = TRUE
            IF Abort% = TRUE THEN

'***** *****
        ' **      System
Shutdown Procedures  **
        ' **
go here!!          **

'***** *****

```

```

        CLS
        SYSTEM
        END IF
        CASE ELSE
        END SELECT
    END SELECT
    '
    LOOP
    '
    'Error Handling and time updates
    GOTO EndLabel:
    '
    NewFileError:
        RESET
        BEEP
        windows w2, w2Text$, w2Title$, w2Prompt$
        WindowsPop
        FileError% = TRUE
        RESUME NEXT
    '
    OldFileError:
        RESET
        BEEP
        windows w2, w2Text$, w2Title$, w2Prompt$
        WindowsPop
        FileError% = TRUE
        RESUME NEXT
    '
    BadPath:
        RESET
        BadPathType% = TRUE
        BEEP
        windows w3, w3Text$, w3Title$, w3Prompt$
        WindowsPop
        RESUME NEXT
    '
    EQofEQError:
        BEEP
        windows w4, w4Text$, w4Title$, w4Prompt$
        WindowsPop
        RESUME
    '
EndLabel:
END

*****
***     Name:      Aberration      **
***     Type:      Subroutine      **
***     Module:    MIRACTRL.BAS    **
***     Language:   Quickbasic 4.50  **
*****

```

```

'
' Calculates the aberration correction to celestial
coordinates.
'

' Example of use: Aberration RAofStar#, DECoStar#, 
RAAberrationCor#, DecAberrationCor#, dayFraction#, DaysPassed%
'

SUB Aberration (RAofStar#, DECoStar#, RAAberrationCor#,
DecAberrationCor#, dayFraction#, DaysPassed%)
'

' Calculate the mean longitude of the Sun, L#
d# = dayFraction# + CDBL(DaysPassed%)
L# = 298.926 + .985647 * d#      ' in degrees
L# = L# * (2# * pi) / 360#       ' Convert to radians
'

' Define the constant of aberration, k#
k# = 20.496                      ' in arc seconds
k# = k# * (2# * pi) / (360# * 3600#)' convert to radians
'

' Define the obliquity of the ecliptic
epsilon# = 23.440332# - .00000036# * d#  'degrees
epsilon# = epsilon# * (2# * pi) / 360#    'radians
'

' Calculate the corrections to RA and Declination due to
aberration
RAAberrationCor# = (-k# * SIN(L#) * SIN(RAofStar#) - k# *
COS(L#) * COS(epsilon#) * COS(RAofStar#)) / COS(DECoStar#)
DecAberrationCor# = -k# * SIN(L#) * COS(RAofStar#) *
SIN(DECoStar#) + k# * COS(L#) * (COS(epsilon#) *
SIN(RAofStar#) * SIN(DECoStar#) - SIN(epsilon#) *
COS(DECoStar#))
'

END SUB

*****
**      Name: Atmospheric      **
**      Type: Subroutine        **
**      Module: MIRACTRL.BAS   **
**      Language: Quickbasic 4.50**
*****


' Allows user entry of atmospheric temperature, pressure and
the center
' wavelength of light for an observation. These data are used
in refraction
' calculations.
'

' Example of use: Atmospheric Pressure!, Temperature!,
Wavelength!
'

SUB Atmospheric (Pressure!, Temperature!, Wavelength!)

```

```

'Clear the body of the status box
COLOR WHITE, BLUE, BLUE
FOR r% = 14 TO 21
    LOCATE r%, 6, 0
    PRINT SPACE$(68);
NEXT r%

LOCATE 15, 55, 0
PRINT Temperature!
LOCATE 15, 10, 0
INPUT "Please enter the atmospheric temperature (C): ", NewTemp$
IF LEN(NewTemp$) > 0 THEN Temperature! = VAL(NewTemp$)
LOCATE 17, 55, 0
PRINT Pressure!
LOCATE 17, 10, 0
INPUT "Please enter the atmospheric pressure (mmHg): ", NewPres$
IF LEN(NewPres$) > 0 THEN Pressure! = VAL(NewPres$)
LOCATE 20, 40, 0
PRINT Wavelength!
LOCATE 19, 10, 0
PRINT "Please enter the center wavelength of light"
LOCATE 20, 10, 0
INPUT "for the chosen star (microns): ", NewWave$
IF LEN(NewWave$) > 0 THEN Wavelength! = VAL(NewWave$)

END SUB

*****
**      Name:      ChooseStar      **
**      Type:      Subroutine      **
**      Module:     MIRACTRL.BAS   **
**      Language:   Quickbasic 4.50  **
*****
'

'Allows choice of stars from list. Sets active star for slew
and track
'of telescope.

'

'Example of use: StarDataEntry Star$(), RA$(), DEC$(),
Epoch$(), NewFile%, w1, w1Text$(), w1Title$, w1Prompt$,
CurrentStar%
'Parameters:      (none)
'Variables:       Star$()      A 5-element matrix of desired
stars for
'
'
RA$()          viewing
'
'
ascensions      A 5-element matrix of right
'
'
for the stars listed in Star$()

```

```

'
DEC$()      A 5-element matrix of declinations
'
Epoch$()    for the stars listed in Star$()
the          A 5-element matrix of epochs for
             RA's and DEC's of the stars listed
in Star$()  in Star$()
'
g%, h%, i%  Looping indices
j%          A flag activated when the
Backspace key is hit
l%          A flag activated when the Shift
Tab, Up or Down
'
row%, col%  Arrows are hit
for the      The current row and column markers
'
Stop$        cursor
not the      A string representing whether or
             character entry loop is to be
exited       exited
kee%         A unique code returned by the
KeyCode% or  KeyCodeNoWait% Functions for each
key hit      key hit
buffer       Character$ A string read from the keyboard
'
the current  Current$   A string representing the entry of
one of the   field. This variable will become
'
following:  Star$(), RA$(),
DEC$(), EPOCH$()
'

'Module Level
'  DECLARATIONS:  DECLARE SUB StarDataEntry (Star$(), RA$(),
DEC$(),
'
                           Epoch$(), NewFile%, w1,
w1Text$, w1Title$, w1Prompt$, CurrentStar%)
'
SUB ChooseStar (Star$(), RA$(), DEC$(), EPOCH$(), NewFile%, w1
AS WindowsType, w1Text$, w1Title$, w1Prompt$, currentStar%)
'

'Initialize variables
current$ = Star$(1)
row% = 17
col% = 9
FirstRow% = 17
LastRow% = 21
LastCol% = 62
FileLine% = 1
TopFileLineShown% = 1
currentStar% = 1

```

```

ChangeFileLine% = FALSE
HomeOrEnd% = FALSE
Stops% = FALSE
QuitEntry% = FALSE
'
'Erase Box
COLOR WHITE, BLUE, BLUE
FOR r% = 14 TO 21
    LOCATE r%, 6, 0
    PRINT SPACE$(68);
NEXT r%
'
'Draw bottom edge of the status box
LOCATE 22, 5, 0
PRINT CHR$(200); STRING$(68, 205); CHR$(188);
'
'Place lettering
COLOR WHITE, BLUE, BLUE
LOCATE 14, 22, 0
PRINT "Choose a star to view and press F10."
LOCATE 15, 33, 0
PRINT "RA           DEC";
LOCATE 16, 10, 0
PRINT "Star Name          hr:min:sec";
LOCATE 16, 48, 0
PRINT CHR$(248); ":"'':"";
LOCATE 16, 63, 0
PRINT "Epoch";
'
'Complete data fields by reading in one key at a time, ESC
exits
DO
    IF TopFileLineShown% < 1 THEN TopFileLineShown% = 1
    IF TopFileLineShown% > 96 THEN TopFileLineShown% = 96
    IF FileLine% < 1 THEN FileLine% = 1
    IF FileLine% > 100 THEN FileLine% = 100
    IF row% < FirstRow% THEN row% = FirstRow%
    IF row% > LastRow% THEN row% = LastRow%
    'Update the row numbers and fields
    FOR g% = 1 TO 5
        COLOR WHITE, BLUE, BLUE
        LOCATE 16 + g%, 6, 0
        PRINT Star$(TopFileLineShown% + g% - 1) +
        SPACE$(19 - LEN(Star$(TopFileLineShown% + g% - 1)))
        LOCATE 16 + g%, 33, 0
        PRINT RA$(TopFileLineShown% + g% - 1) +
        SPACE$(10 - LEN(RA$(TopFileLineShown% + g% - 1)))
        LOCATE 16 + g%, 48, 0

```

```

        PRINT DEC$(TopFileLineShown% + g% - 1) +
SPACE$(10 - LEN(DEC$(TopFileLineShown% + g% - 1)))
        LOCATE 16 + g%, 63, 0
        PRINT EPOCH$(TopFileLineShown% + g% - 1) +
SPACE$(10 - LEN(EPOCH$(TopFileLineShown% + g% - 1)))
NEXT g%
COLOR BLACK, WHITE, BLUE
LOCATE row%, 10, 0
PRINT Star$(FileLine%) + SPACE$(19) -
LEN(Star$(FileLine%)))
LOCATE row%, 33, 0
PRINT RA$(FileLine%) + SPACE$(10) -
LEN(RA$(FileLine%)))
LOCATE row%, 48, 0
PRINT DEC$(FileLine%) + SPACE$(10) -
LEN(DEC$(FileLine%)))
LOCATE row%, 63, 0
PRINT EPOCH$(FileLine%) + SPACE$(10) -
LEN(EPOCH$(FileLine%)))
'
COLOR WHITE, BLACK, BLUE
'Read the next character and operate on it
DO
    LOCATE row%, col% + 1, 1
    kee% = KeyCode%(Character$)
    SELECT CASE kee%
        CASE ENTER
            IF r~w% = LastRow% AND col% =
LastCol% THEN
                row% = row% + 1
                IF row% = LastRow% +
1 THEN
                    TopFileLineShown% = TopFileLineShown% + 1
                    FileLine% =
FileLine% + 1
                    END IF
                ELSE
                    'no action
                END IF
                Stops% = TRUE
            CASE ESCAPE
                windows w1, w1Text$(),
w1Title$, w1Prompt$
                IF w1.returnCode = 121 OR
w1.returnCode = 89 THEN Abort% = TRUE
                IF Abort% = TRUE THEN
'***** ****
                    *** System Shutdown
Procedures **

```

```

        ***          go here!!
**
*****
WindowsPop
CLS
SYSTEM
ELSE
WindowsPop
END IF
CASE UPARROW
IF row% = FirstRow% THEN
    TopFileLineShown% =
TopFileLineShown% - 1
END IF
row% = row% - 1
ChangeFileLine% = MINUS
Stops% = TRUE
CASE DOWNARROW
IF row% = LastRow% THEN
    TopFileLineShown% =
TopFileLineShown% + 1
END IF
row% = row% + 1
ChangeFileLine% = PLUS
Stops% = TRUE
CASE PGDOWN
IF TopFileLineShown% > 95 THEN
    ChangeFileLine% = PLUS
* (96 - TopFileLineShown%)
    TopFileLineShown% = 96
ELSE
    TopFileLineShown% =
TopFileLineShown% + 5
    ChangeFileLine% = PLUS
* 5
END IF
Stops% = TRUE
CASE PGUP
IF TopFileLineShown% < 6 THEN
    ChangeFileLine% =
MINUS * (TopFileLineShown% - 1)
    TopFileLineShown% = 1
ELSE
    TopFileLineShown% =
TopFileLineShown% - 5
    ChangeFileLine% =
MINUS * 5
END IF
Stops% = TRUE
CASE HOME

```

```

        TopFileLineShown% = 1
        HomeOrEnd% = 1
        Stops% = TRUE
    CASE ENDFILE
        TopFileLineShown% = 96
        HomeOrEnd% = 100
        Stops% = TRUE
    CASE 17408 'F10
        currentStar% = FileLine%
        QuitEntry% = TRUE
        Stops% = TRUE
    CASE ELSE
        CASE ELSE

' *****
' **
' **
' ** Error Handling
Needed      **
' **
' **
' *****

        END SELECT
    LOOP UNTIL Stops% = TRUE
    Stops% = FALSE
    IF FileLine% < 1 THEN FileLine% = 1
    IF FileLine% > 100 THEN FileLine% = 100
    IF ChangeFileLine% <> FALSE THEN
        FileLine% = FileLine% + ChangeFileLine%
        IF FileLine% < 1 THEN FileLine% = 1
        IF FileLine% > 100 THEN FileLine% = 100
        ChangeFileLine% = FALSE
        current$ = Star$(FileLine%)
    END IF
    '
    IF HomeOrEnd% <> FALSE THEN
        IF HomeOrEnd% = 1 THEN
            row% = FirstRow%
        ELSEIF HomeOrEnd% = 100 THEN
            row% = LastRow%
        END IF
        FileLine% = HomeOrEnd%
        HomeOrEnd% = FALSE
        COLOR WHITE, BLUE, BLUE
        LOCATE LastRow%, 9, 0
        PRINT SPACE$(1)
        current$ = Star$(FileLine%)
    END IF
    '
    LOOP UNTIL QuitEntry% = TRUE

```

```
END SUB
```

```
'*****  
'**      Name:      DayLightSaving    **  
'**      Type:      Subroutine      **  
'**      Module:     MIRACTRL.BAS   **  
'**      Language:   Quickbasic 4.50  **  
'*****  
  
'Determine if Daylight Savings Time applies to local time  
  
'Example of use:  DayLightSaving DayLightSavings%, hour%,  
day%, month%, year%  
'Parameters:      (none)  
'Variables:       DayLightSavings% A varaiable representing  
whether or not  
                                daylight savings time is  
in effect  
'                                dayOfWeek$ A string representing the day of  
the week  
'                                SundayPassed% A varialbe representing  
whether or not  
                                Sunday has passed in the  
current week  
'                                hour%      The current hour of time  
'                                day%       The current day  
'                                month%     The current month  
'                                year%      The current year  
  
'Module Level  
  
'      DECLARATIONS:  DECLARE SUB DayLightSaving  
(DayLightSavings%, hour%, day%, month%, year%)  
  
SUB DayLightSaving (DayLightSavings%, hour%, day%, month%,  
year%)  
    '*****  
    '** Determine the day of the week (=dayOfWeek$)  **  
    '** and if Sunday has passed (=SundayPassed$)    **  
    '*****  
    IF month% > 10 OR month% < 4 THEN DayLightSavings% =  
FALSE  
    IF month% > 4 AND month% < 10 THEN DayLightSavings% =  
TRUE  
    IF month% = 4 THEN  
        IF day% > 23 THEN  
            IF dayOfWeek$ = "SUNDAY" THEN  
                IF hour% > 1 THEN  
DayLightSavings% = TRUE  
ELSEIF SundayPassed% = TRUE THEN  
DayLightSavings% = TRUE
```

```

        END IF
    END IF
    IF month% = 10 THEN
        IF day% > 24 THEN
            IF dayOfWeek$ = "SUNDAY" THEN
                IF hour% > 1 THEN
DayLightSavings% = TRUE
ELSEIF SundayPassed% = TRUE THEN
DayLightSavings% = TRUE
                END IF
            END IF
        END IF
    END IF
END SUB

' *****
' **      Name:      EquationEquinoxes **
' **      Type:      Subroutine      **
' **      Module:     MIRACTRL.BAS   **
' **      Language:   Quickbasic 4.50  **
' *****

'
' Allows input of time correction argument from the Equation of
Equinoxes
'(See page B6-B15, Astronomical Almanac) (Correction needed due
to nutation
'of Earth)
'

' Example of use: EquationEquinoxes EQofEQ(), EQofEQUIN!
' Parameters:      (none)
' Variables:       EQofEQ()      An array of the the equation of
equinox
'
'
'
correction for
'
'
'
the current
'
'
'
NextDay$      the current time
A string representing either
'
'
'
the current
'
'
'
NextDate$      or next day
A string representing either
'
'
'
day, year)      or the next total date (month,
'
'

' Module Level
' DECLARATIONS: DECLARE SUB EquationEquinoxes (EQofEQ(),
EQofEQUIN!)

'
SUB EquationEquinoxes (EQofEQ(), EQofEQUIN!)
    'Input the equation of equinox corrections for the
current day

```

```

'and the next day
COLOR WHITE, BLUE, BLUE
FOR i% = 1 TO 2
    NextDay$ = LTRIM$(STR$(VAL(MID$(DATE$, 4, 2))
+ i% - 1))
    NextDate$ = LEFT$(DATE$, 3) + NextDay$ +
RIGHT$(DATE$, 5)
    LOCATE 16 + i%, 57, 0
    PRINT EQofEQ(i%)
    LOCATE 16 + i%, 10, 0
    PRINT "Enter the Equation of Equinoxes for ";
NextDate$; ";";
    LOCATE 17 + i%, 10, 0
    PRINT "(in seconds)"
    LOCATE 16 + i%, 58, 1
    INPUT "", EQofEQ(i%)
NEXT i%
'
'Interpolate to find the equation of equinox
correction for the
'current time
EQofEQUIN! = (TIMER * (EQofEQ(2) - EQofEQ(1)) /
86400!) + EQofEQ(1)
'
'Erase the above cues
FOR i% = 1 TO 4
    LOCATE 16 + i%, 21, 0
    PRINT SPACE$(45)
NEXT i%
END SUB

*****
'*      Name:      InitialDisplay      **
'*      Type:      Subroutine      **
'*      Module:     MIRACTRL.BAS      **
'*      Language:   Quickbasic 4.50      **
*****
'

'Creates the initial display screen graphics
'

'Example of use:  InitialDisplay
'Parameters:      (none)
'Variables:       i%, r%      Looping indices
'

'Module Level
'    DECLARATIONS:  DECLARE SUB InitialDisplay
'

SUB InitialDisplay
CLS
'Color the background
FOR i% = 1 TO 25

```

```

        COLOR BLUE, BLACK, BLUE
        PRINT STRING$(80, 177)
NEXT i%
LOCATE 24, 1, 0
PRINT STRING$(80, 177);
LOCATE 25, 1, 0
PRINT STRING$(80, 177);

'Place the lettering
LOCATE 4, 18, 0
COLOR WHITE, BLUE, BLUE
PRINT "Monterey Institute for Research in Astronomy"
LOCATE 6, 28, 0
PRINT "Oliver Observing Station"
LOCATE 10, 30, 0
PRINT "Automated Telescope"
LOCATE 11, 32, 0
PRINT "Control Package"

'Draw top edge of status box
LOCATE 13, 5, 0
COLOR WHITE, BLUE, BLUE
PRINT CHR$(201); STRING$(68, 205); CHR$(187);

'Draw the body of the status box
FOR r% = 14 TO 21
    LOCATE r%, 5, 0
    PRINT CHR$(186);
    PRINT SPACE$(68);
    PRINT CHR$(186);
NEXT r%

'Draw bottom edge of the status box
LOCATE 22, 5, 0
PRINT CHR$(200); STRING$(68, 205); CHR$(188);

'Display "Initializing" status in box
COLOR WHITE + BLINK, BLUE, BLUE
LOCATE 17, 25, 0
PRINT "Initializing"
LOCATE 17, 37, 0
COLOR WHITE, BLUE, BLUE
PRINT " - Please stand by"
END SUB

***** *****
**      Name:      KeyCode%      **
**      Type:      Function      **
**      Module:     MIRACTRL.BAS  **
**      Language:   Quickbasic 4.50  **
*****

```

```

'
'Returns a unique integer for any key pressed
'
'Example of Use: kee% = KeyCode% Character$
'Parameters:      (none)
'Variables:       Character$      A string representing the
current keyboard
'                           buffer value
'
'Module Level
'DECLARATIONS:  DECLARE FUNCTION KeyCode% (Character$)
'
FUNCTION KeyCode% (Character$) STATIC
    DO
        Character$ = INKEY$
        LOOP UNTIL Character$ <> ""
        KeyCode% = CVI(Character$ + CHR$(0))
    END FUNCTION

*****
'**      Name:      KeyCodeNoWait%   **
'**      Type:     Function      **
'**      Module:    MIRACTRL.BAS  **
'**      Language:  Quickbasic 4.50 **
*****
'

'Returns a unique integer for any key pressed, but does not
wait
'for a key to be pressed
'
'Example of Use: kee% = KeyCodeNoWait% Character$
'Parameters:      (none)
'Variables:      (none)
'Module Level
'DECLARATIONS:  DECLARE FUNCTION KeyCode% (Character$)
'
FUNCTION KeyCodeNoWait% (Character$) STATIC
    Character$ = INKEY$
    IF Character$ <> "" THEN
        KeyCodeNoWait% = CVI(Character$ + CHR$(0))
    ELSE
        KeyCodeNoWait% = 0
    END IF
END FUNCTION

*****
'**      Name:      Nutation      **
'**      Type:     Subroutine    **
'**      Module:    MIRACTRL.BAS  **
'**      Language:  Quickbasic 4.50 **
*****

```

```

'
' Calculates the nutation corrections to celestial coordinates
for a given
'day and time.
'

' Example of use: Nutation RAofStar#, DECoFStar#, JulianDate#,
N#(), S1#(), S2#()

'

' Note: Astronomical Almanac notation is used wherever possible
'

SUB Nutation (RAofStar#, DECoFStar#, JulianDate#, N#(), S1#(),
S2#())
'

' Unmodified celestial coordinates converted to rectangular
form
S1#(1) = COS(DECoFStar#) * COS(RAofStar#)
S1#(2) = COS(DECoFStar#) * SIN(RAofStar#)
S1#(3) = SIN(DECoFStar#)

'

' Nutation Parameters
d# = JulianDate# - 2447871.5#
epsilon# = 23.440332# * (2# * pi) / 360#

DeltaPsi# = -.0048# * SIN((279.9# - .053# * d#) * (2# * pi) /
/ 360#) - .0004# * SIN((197.9# + 1.971# * d#) * (2# * pi) /
360#)
DeltaPsi# = DeltaPsi# * (2# * pi) / 360#

DeltaEpsilon# = .0026# * COS((279.9# - .053# * d#) * (2# * pi) /
/ 360#) + .0002# * COS((197.9# + 1.971# * d#) * (2# * pi) /
/ 360#)
DeltaEpsilon# = DeltaEpsilon# * (2# * pi) / 360#

'

' Nutation Rotation Matrix
N#(1, 1) = 1
N#(1, 2) = -DeltaPsi * COS(epsilon#)
N#(1, 3) = -DeltaPsi * SIN(epsilon#)
N#(2, 1) = DeltaPsi * COS(epsilon#)
N#(2, 2) = 1
N#(2, 3) = -DeltaEpsilon#
N#(3, 1) = DeltaPsi * SIN(epsilon#)
N#(3, 2) = DeltaEpsilon#
N#(3, 3) = 1

'

' Celestial Coordinates modified for nutation in rectangular
form
S2#(1) = N#(1, 1) * S1#(1) + N#(1, 2) * S1#(2) + N#(1, 3) *
S1#(3)
S2#(2) = N#(2, 1) * S1#(1) + N#(2, 2) * S1#(2) + N#(2, 3) *
S1#(3)

```

```

S2#(3) = N#(3, 1) * S1#(1) + N#(3, 2) * S1#(2) + N#(3, 3) *
S1#(3)
'
'Decode S1#() elements from rectangular form to RA and DEC
IF S2#(1) >= 0# AND S2#(2) < 0# THEN
    alpha# = ATN(-S2#(1) / S2#(2)) + (3# * pi) / 2#
ELSEIF S2#(1) < 0# AND S2#(2) < 0# THEN
    alpha# = ATN(S2#(2) / S2#(1)) + pi
ELSEIF S2#(1) < 0# AND S2#(2) >= 0# THEN
    alpha# = ATN(-S2#(1) / S2#(2)) + pi / 2#
ELSE alpha# = ATN(S2#(2) / S2#(1))
END IF
'
delta# = ATN(S2#(3) * COS(alpha#) / S2#(1))
'
RAofStar# = alpha#
DECofStar# = delta#
'
END SUB

*****
'**      Name:      Precession      **
'**      Type:      Subroutine      **
'**      Module:    MIRACTRL.BAS   **
'**      Language:   Quickbasic 4.50  **
*****
'

'Calculates the precession corrections to celestial
coordinates for a given
'day and time.
'

'Example of use: Precession RA$(), DEC$(), EPOCH$(),
currentStar%, JulianDate#, EpochError%, RAofStar#, DECoStar#, RADecError%
'

'Note: Astronomical Almanac notation is used wherever possible
'

SUB Precession (RA$(), DEC$(), EPOCH$(), currentStar%,
JulianDate#, EpochError%, RAofStar#, DECoStar#, RADecError%)
'
'Decode RA$() and DEC$() into radians
RAFirstColon% = INSTR(RA$(currentStar%), ":")
RASecondColon% = INSTR((RAFirstColon% + 1),
RA$(currentStar%), ":")
IF RAFirstColon% = 0 OR RASecondColon% = 0 THEN
    RADecError% = TRUE
ELSE
    RAofStar# = VAL(LEFT$(RA$(currentStar%),
(RAFirstColon% - 1)) * 3600 + VAL(MID$(RA$(currentStar%),
(RAFirstColon% + 1), 2)) * 60 + VAL(RIGHT$(RA$(currentStar%),
(LEN(RA$(currentStar%)) - RASecondColon%)))

```

```

    RAofStar# = RAofStar# * 2# * pi / 86164.09053#
'Seconds of time converted to radians of sidereal movement
    DO WHILE RAofStar# >= (2# * pi)
        RAofStar# = RAofStar# - (2# * pi)
    LOOP
    DO WHILE RAofStar# < 0
        RAofStar# = RAofStar# + (2# * pi)
    LOOP
END IF
DECFirstColon% = INSTR(DEC$(currentStar%), ":")
DECSecondColon% = INSTR((DECFirstColon% + 1),
DEC$(currentStar%), ":")
IF DECFirstColon% = 0 OR DECSecondColon% = 0 THEN
    RADecError% = TRUE
ELSE
    DECoFStar# = VAL(LEFT$(DEC$(currentStar%),
(DECFirstColon% - 1)) + VAL(MID$(DEC$(currentStar%),
(DECFirstColon% + 1), 2)) / 60 +
VAL(RIGHT$(DEC$(currentStar%),
(LEN(RA$(currentStar%)) -
DECFirstColon%)) / 3600
    DECoFStar# = DECoFStar# * (2# * pi) / 360#'Degrees
converted to radians
    DO WHILE DECoFStar# > (pi / 2#)
        DECoFStar# = pi - DECoFStar#
    LOOP
    DO WHILE DECoFStar# < -(pi / 2#)
        DECoFStar# = -(pi + DECoFStar#)
    LOOP
END IF
'Account for precession between the epoch chosen by the user
and J2000.0
'
'Determine the Julian date at last midnight
year! = VAL(EPOCH$(currentStar%))
IF year! = 0 THEN
    EpochError% = TRUE
ELSE
    JDofEpoch! = (4713! + year!) * 365.25
    '
    'Using the approximate formulae for precession in the
Astronomical Almanac:
    'Account for precession FROM the star's known epoch TO
J2000.0
    T1# = (CDBL(JDofEpoch!) - 2451545#) / 36525#
    M1# = 1.2812323# * T1# + .0003879# * T1# ^ 2 +
.0000101# * T1# ^ 3' degrees
    N1# = .5567530000000001# * T1# - .0001185# * T1# ^ 2
- .0000116# * T1# ^ 3' degrees
    '
    alpha# = RAofStar# * 360# / (2# * pi)' RA in degrees

```

```

        delta# = DECoFStar# * 360# / (2# * pi)'           Dec in
degrees
        alphaInRad# = RAoFStar#
        deltaInRad# = DECoFStar#
        '
        'M suffix refers to mean epoch
        alphaM# = alpha# - .5# * (M1# - N1# * SIN(alphaInRad#)
* TAN(deltaInRad#))
        alphaMinRad# = alphaM# * (2# * pi) / 360#
        deltaM# = delta# - .5# * N1# * COS(alphaMinRad#)
        deltaMinRad# = deltaM# * (2# * pi) / 360#
        '
        '0 suffix refers to J2000.0 epoch
        alpha0# = alpha# - M1# - N1# * SIN(alphaMinRad#) *
TAN(deltaMinRad#)
        alpha0InRad# = alpha0# * (2# * pi) / 360#
        delta0# = delta# - N1# * COS(alphaMinRad#)
        delta0InRad# = delta0# * (2# * pi) / 360#
        '
        'Account for precession FROM J2000.0 TO the current
date
        T2# = (CDBL(JulianDate#) - 2451545#) / 36525#
        M2# = 1.2812323# * T2# + .0003879# * T2# ^ 2 +
.0000101# * T2# ^ 3' degrees
        N2# = .5567530000000001# * T2# - .0001185# * T2# ^ 2
- .0000116# * T2# ^ 3' degrees
        '
        'M suffix refers to mean epoch
        alphaM# = alpha0# - .5# * (M2# - N2# * *
SIN(alpha0InRad#) * TAN(delta0InRad#))
        alphaMinRad# = alphaM# * (2# * pi) / 360#
        deltaM# = delta0# - .5# * N2# * COS(alphaMinRad#)
        deltaMinRad# = deltaM# * (2# * pi) / 360#
        '
        'The new alpha and delta are the CORRECTED right
ascension and
        'declination IN DEGREES
        alpha# = alpha0# - M2# - N2# * SIN(alphaMinRad#) *
TAN(deltaMinRad#)
        delta# = delta0# - N2# * COS(alphaMinRad#)
        RAoFStar# = alpha# * (2# * pi) / 360#
        DO WHILE RAoFStar# >= (2# * pi)
            RAoFStar# = RAoFStar# - (2# * pi)
        LOOP
        DO WHILE RAoFStar# < 0
            RAoFStar# = RAoFStar# + (2# * pi)
        LOOP
        '
        DECoFStar# = delta# * (2# * pi) / 360#
        DO WHILE DECoFStar# > (pi / 2#)
            DECoFStar# = pi - DECoFStar#

```

```

        LOOP
        DO WHILE DECoFStar# < -(pi / 2#)
            DECoFStar# = -(pi + DECoFStar#)
        LOOP
    '
END IF
END SUB

'*****
'*      Name:      Refraction      **
'*      Type:      Subroutine      **
'*      Module:     MIRACTRL.BAS   **
'*      Language:   Quickbasic 4.50  **
'*****

'
'Calculates the refraction corrections to celestial
coordinates for a given
'day and time. Uses data from the Atmospheric Subroutine.
'
'Example of use: Refraction Pressure!, Temperature!,
Wavelength!, GMST0hrSEC#, HourAngle#, GHAaries#, RAofStar#,
DECoFStar#, RARefractCor#, DecRefractCor#, RAofMIRA#
'

SUB Refraction (Pressure!, Temperature!, Wavelength!,
GMST0hrSEC#, HourAngle#, GHAaries#, RAofStar#, DECoFStar#,
RARefractCor#, DecRefractCor#, RAofMIRA#)
'
'Initialize variables
MIRALatitude# = .63908139#           'radians, (36 deg 37 min
North)
MIRAlongitude# = 2.12639281#         'radians, (121 deg 50 min
West)
sec% = VAL(RIGHT$(TIME$, 2))
min% = VAL(MID$(TIME$, 4, 2))
hour% = VAL(LEFT$(TIME$, 2))
'
'Calculate Greenwich Mean Sidereal Time
'Determine the day fraction that has passed since midnight
'in SECONDS
sec& = sec%
hours& = hour%
mins& = min%
dayFraction# = CDBL((hours& * 3600) + (mins& * 60) + sec&)
'
'Determine the Greenwich Mean Sidereal Time at the current
'time in SECONDS
GMSTinSEC# = GMST0hrSEC# + dayFraction#
'
'Calculate Greenwich Hour Angle of Aries
GHAaries# = GMSTinSEC# * 2# * pi / 86164.09053099999#'Seconds
of solar time converted to radians

```

```

DO WHILE GHAaries# > (2# * pi)
    GHAaries# = GHAaries# - (2# * pi)
LOOP
DO WHILE GHAaries# < 0#
    GHAaries# = GHAaries# + (2# * pi)
LOOP
'
'Calculate the current star's hour angle for this moment
RAofMIRA# = GHAaries# - MIRALongitude#
'
HourAngle# = RAofMIRA# - RAofStar#    'in radians
DO WHILE HourAngle# > (2# * pi)
    HourAngle# = HourAngle# - (2# * pi)
LOOP
DO WHILE HourAngle# < 0#
    HourAngle# = HourAngle# + (2# * pi)
LOOP
'
'Calculate the refraction constant in seconds
k# = 21.3 * Pressure! * (1 + (.0057 / (Wavelength! ^ 2))) /
(273 + Temperature!)
'Note: Temperature in degrees C, Pressure in mmHg, Wavelength
in microns
'      K# in seconds of arc
'
'Calculate the refraction corrections in seconds
RARefractCor# = k# * SIN(HourAngle#) / ((COS(DECofStar#) ^ 2)
* (TAN(DECofStar#) * TAN(MIRALatitude#) + COS(HourAngle#)))
DecRefractCor# = k# * (TAN(DECofStar#) * COS(HourAngle#) -
TAN(MIRALatitude#)) / (TAN(DECofStar#) * TAN(MIRALatitude#) +
COS(HourAngle#))
RARefractCor# = RARefractCor# * (2# * pi) / 3600#' Convert
seconds of arc to radians
DecRefractCor# = DecRefractCor# * (2# * pi) / 3600#' Convert
seconds of arc to radians
'
END SUB

*****
**      Name:      SetCommands      **
**      Type:      Subroutine      **
**      Module:     MIRACTRL.BAS   **
**      Language:   Quickbasic 4.50  **
*****
'
'Controls motor commands to set the Telescope following a slew
to a new
'target star.  Uses the stepping motors in high speed mode.
'

```

```

'Example of use: SetCommands Simulation%, TeleRA#, TeleDec#,
RAofStar#, DECofStar#
'
SUB SetCommands (Simulation%, TeleRA#, TeleDEC#, RAofStar|,
DECofStar#)
'
'Set as appropriate in Right Ascension
DO
    RAdiff# = TeleRA# - RAofStar#
    'Update motor display box
    COLOR RED, BLUE, BLUE
    LOCATE 3, 20, 0
    PRINT CHR$(222)
    'POKE (address of RA Set controller), X
    '*****
    IF Simulation% = TRUE THEN
        IF RAdiff# > 0 THEN TeleRA# = TeleRA# - .00001
        IF RAdiff# < 0 THEN TeleRA# = TeleRA# + .00001
    ELSE
        'Read new TeleRA# from encoders
    END IF
    '
    'Check limit switch
    'PEEK (address of limit switch)
    '*****
    '** Change RAdiff# < .0003 to   **
    '** acceptable limit           **
LOOP UNTIL RAdiff# < .0003
    '*****'

'Set as appropriate in Declination
DO
    DECdiff# = TeleDEC# - DECofStar#
    'Update motor display box
    COLOR RED, BLUE, BLUE
    LOCATE 3, 70, 0
    PRINT CHR$(222)
    'POKE (address of DEC Set controller), X
    '*****
    IF Simulation% = TRUE THEN
        IF DECdiff# > 0 THEN TeleDEC# = TeleDEC# -
.00001
        IF DECdiff# < 0 THEN TeleDEC# = TeleDEC# +
.00001
    ELSE
        'Read new TeleDEC# from encoders
    END IF
    '
    'Check limit switch
    'PEEK (address of limit switch)
    '*****

```



```

        '
        FirstColonLoc% = INSTR(SiderealTime$, ":")
        LastColonLoc%   =   INSTR((FirstColonLoc% + 1),
SiderealTime$, ":")
        LengthSec% = 2
        BeginMin% = FirstColonLoc% + 1
        '
        SiderealSec$ = LTRIM$(STR$(VAL(RIGHT$(SiderealTime$,
LengthSec%)) + 1)))
        SiderealMin$ = MID$(SiderealTime$, BeginMin%,
LastColonLoc% - FirstColonLoc% - 1)
        SiderealHr$ = LEFT$(SiderealTime$, (FirstColonLoc% -
1))
        '
        IF LEN(SiderealSec$) = 1 THEN SiderealSec$ = "0" +
SiderealSec$
        IF VAL(SiderealSec$) > 59 THEN
            SiderealMin$ = STR$(VAL(SiderealMin$) + 1)
            SiderealSec$ = "00"
        END IF
        IF VAL(SiderealMin$) > 59 THEN
            SiderealHr$ = STR$(VAL(SiderealHr$) + 1)
            SiderealMin$ = "00"
        END IF
        IF VAL(SiderealHr$) > 23 THEN SiderealHr$ = "00"
        SiderealTime$ = LTRIM$(SiderealHr$) + ":" + +
LTRIM$(SiderealMin$) + ":" + LTRIM$(SiderealSec$)
        COLOR WHITE, BLUE, BLUE
        LOCATE 15, 58, 0
        PRINT LTRIM$(LocalTime$)
        LOCATE 18, 58, 0
        PRINT LTRIM$(SiderealTime$); SPACE$(1)
        LOCATE 21, 58, 0
        PRINT TIME$
END SUB

*****
'**      Name:      SiderealTimeCalc  **
'**      Type:      Subroutine       **
'**      Module:    MIRACTRL.BAS    **
'**      Language:  Quickbasic 4.50  **
*****
'
'Calculate Local Apparent Sidereal Time
'
'Example of use: SiderealTimeCalc SiderealTime$, min%, hour%,
day%, month%, year%, EQofEQUIN!, GMST0hrSEC#, JulianDate#,
dayFraction#, DaysPassed%
'Parameters:      (none)
'Variables:       SiderealTime$     A string representing the
local

```

```

corrections to
dates (due
nutation
since a known
above in string
the value of
not the
or not noon
the current
seconds
calculation of
at 0 hour,
Astronomical
at 0 hour on
current day that
at the current
length when a
reset

EQofEQ!()      apparent sidereal time
                           The Equation of Equinox
                           sidereal time for successive
                           to corrections for the Earth's

min%           The current minute of time
hour%          The current hour of time
day%           The current day
month%         The current month
year%          The current year
yearDiff!      The number of years passed
               leap year (1992) divided by 4
yearDiffStrg$  The decimal portion of the
               form. yearDiff! then becomes
               this string
               A string indicating whether or
               current year is a leap year
noonPassed$    A string representing whether
               has passed on the current day
daysPassed%    The number of days passed in
               year
JulianDate#   The current Julian date in
               Tu#           A variable used in the
                           Greenwich Mean Sidereal Time
                           following the convention in the
                           Almanac
                           Greenwich Mean Sidereal Time
                           the current day
dayFraction!   The decimal fraction of the
                           has passed by the current time
                           Greenwich Mean Sidereal Time
                           time in seconds
                           A string that becomes of some
                           loop is to be exited, then is
                           OKtoQUIT$
```

```

'Module Level
' DECLARATIONS:      DECLARE SUB SiderealTimeCalc
(SiderealTime$, min%, hour%, day%, month%, year%, EQofEQUIN!,
GMST0hrSEC#,   GMST0hrSEC#, JulianDate#,   dayFraction#,
DaysPassed%)

SUB SiderealTimeCalc (SiderealTime$, min%, hour%, day%,
month%, year%, EQofEQUIN!, GMST0hrSEC#, JulianDate#,
dayFraction#, DaysPassed%)
    'Determine if the current year is a leap year
    yearDiff! = CSNG(year% - 1992) / 4!
    yearDiffStrg$ = RIGHT$(STR$(yearDiff!), ,
(LEN(yearDiff!) - INSTR(STR$(yearDiff%), ".") ))
    yearDiff! = VAL(yearDiffStrg$)
    IF yearDiff! > 0 THEN
        LeapYear$ = "FALSE"
    ELSE
        LeapYear$ = "TRUE"
    END IF
    '
    'Determine if noon has passed on the current day
    IF hour% < 12 THEN
        noonPassed$ = "FALSE"
    ELSE
        noonPassed$ = "TRUE"
    END IF
    'Determine days that have passed in the current year
    SELECT CASE month%
        CASE 1
            DaysPassed% = day% - 1
        CASE 2
            DaysPassed% = day% + 30
        CASE 3
            IF LeapYear$ = "TRUE" THEN
                DaysPassed% = day% + 59
            ELSE
                DaysPassed% = day% + 58
            END IF
        CASE 4
            IF LeapYear$ = "TRUE" THEN
                DaysPassed% = day% + 90
            ELSE
                DaysPassed% = day% + 89
            END IF
        CASE 5
            IF LeapYear$ = "TRUE" THEN
                DaysPassed% = day% + 120
            ELSE
                DaysPassed% = day% + 119
            END IF

```

```

CASE 6
    IF LeapYear$ = "TRUE" THEN
        DaysPassed% = day% + 151
    ELSE
        DaysPassed% = day% + 150
    END IF
CASE 7
    IF LeapYear$ = "TRUE" THEN
        DaysPassed% = day% + 181
    ELSE
        DaysPassed% = day% + 180
    END IF
CASE 8
    IF LeapYear$ = "TRUE" THEN
        DaysPassed% = day% + 212
    ELSE
        DaysPassed% = day% + 211
    END IF
CASE 9
    IF LeapYear$ = "TRUE" THEN
        DaysPassed% = day% + 243
    ELSE
        DaysPassed% = day% + 242
    END IF
CASE 10
    IF LeapYear$ = "TRUE" THEN
        DaysPassed% = day% + 273
    ELSE
        DaysPassed% = day% + 272
    END IF
CASE 11
    IF LeapYear$ = "TRUE" THEN
        DaysPassed% = day% + 304
    ELSE
        DaysPassed% = day% + 303
    END IF
CASE 12
    IF LeapYear$ = "TRUE" THEN
        DaysPassed% = day% + 334
    ELSE
        DaysPassed% = day% + 333
    END IF
CASE ELSE
END SELECT
'
'Determine the Julian date at last midnight
'***** ****
'** Change to be good for all years, not just 1992 **
JulianDate# = CDBL(2448621.5# + DaysPassed% + 1)  '**'
'***** ****
'
```

```

        'Determine the Greenwich Mean Sidereal Time at 0 hour
UTC
        'in SECONDS
        Tu# = (JulianDate# - 2451545#) / 36525!
        GMST0hrSEC# = 24110.54841# + (8640184.812866# * Tu#)
+ (.093104 * (Tu) ^ 2) - (6.2 * (10 ^ (-6)) * (Tu) ^ 3)
        '
        'Determine the day fraction that has passed since
midnight
        'in SECONDS
        hours& = hour%
        mins& = min%
        dayFraction# = CDBL((hours& * 3600) + (mins& * 60))
        '
        'Determine the Greenwich Mean Sidereal Time at the
current
        'time in SECONDS
        GMSTinSEC# = GMST0hrSEC# + dayFraction#
        '
        'Calculate Greenwich Apparent Sidereal Time in
SECONDS
        GASTinSEC# = GMSTinSEC# + CDBL(EQofEQ!)
        'Add or subtract multiples of 24 hours as necessary
        DO
            IF GASTinSEC# < 0 THEN
                GASTinSEC# = GASTinSEC# + (24! * 3600!)
            ELSEIF GASTinSEC# > 86400 THEN
                GASTinSEC# = GASTinSEC# - (24 * 3600)
            END IF
            IF GASTinSEC# >= 0 AND GASTinSEC# <= 86400
        THEN OKtoQUIT% = TRUE
            LOOP UNTIL OKtoQUIT% = TRUE
            OKtoQUIT% = TRUE
            '
            GASThr& = INT(GASTinSEC# / 3600)
            GASTmin% = INT((GASTinSEC# - (GASThr& * 3600)) / 60)
            GASTsec% = INT(GASTinSEC# - (GASThr& * 3600) -
(GASTmin% * 60))
            SiderealTime$ = LTRIM$(STR$(GASThr&)) + ":" +
LTRIM$(STR$(GASTmin%)) + ":" + LTRIM$(STR$(GASTsec%))
        END SUB

'*****
'*      Name:      SlewCommands      **
'*      Type:      Subroutine      **
'*      Module:    MIRACTRL.BAS    **
'*      Language:  Quickbasic 4.50  **
'*****'
'
'Controls motor commands to slew the Telescope to a new
'target star.  Uses the slew motors.

```

```

'Example of use: SlewCommands (Simulation%, TeleRA#, TeleDec#, RAofStar#, DECoFStar#)
'
SUB SlewCommands (Simulation%, TeleRA#, TeleDEC#, RAofStar#, DECoFStar#)
'
'Slew as appropriate in Right Ascension
DO
    RAdiff# = TeleRA# - RAofStar#
    DO WHILE RAdiff# >= (2# * pi)
        RAdiff# = RAdiff# - (2# * pi)
    LOOP
    DO WHILE RAdiff# <= -(2# * pi)
        RAdiff# = RAdiff# + (2# * pi)
    LOOP
    'Update motor display box
    COLOR RED, BLUE, BLUE
    LOCATE 3, 10, 0
    PRINT CHR$(222)
    'Turn RA slew motor ON
    'POKE (address of RA slew controller), X
    '*****
    IF Simulation% = TRUE THEN
        IF RAdiff# > 0 THEN TeleRA# = TeleRA# - .0001
        IF RAdiff# < 0 THEN TeleRA# = TeleRA# + .0001
    ELSE
        'Read new TeleRA# from encoders
    END IF
    '
    'Check limit switch
    'PEEK (address of limit switch)
    '*****
    '** Change RAdiff# < .0003 to   **
    '** acceptable limit           **
LOOP UNTIL RAdiff# < .0003
'*****
'Update motor display box
COLOR RED, BLUE, BLUE
LOCATE 3, 10, 0
PRINT SPACE$(1)
'
'Slew as appropriate in Declination
DO
    DECdiff# = TeleDEC# - DECoFStar#
    'Update motor display box
    COLOR RED, BLUE, BLUE
    LOCATE 3, 60, 0
    PRINT CHR$(222)
    'Turn DEC slew motor ON
    'POKE (address of DEC slew controller), X

```

```

'*****
IF Simulation% = TRUE THEN
    IF DECdiff# > 0 THEN TeleDEC# = TeleDEC# -
.0001
    IF DECdiff# < 0 THEN TeleDEC# = TeleDEC# +
.0001
ELSE
    'Read new TeleDEC# from encoders
END IF
'
'Check limit switch
'PEEK (address of limit switch)
'*****
'** Change DECdiff# < .0003 to   **
'** acceptable limit      **
LOOP UNTIL DECdiff# < .0003
'*****
'Update motor display box
COLOR RED, BLUE, BLUE
LOCATE 3, 60, 0
PRINT SPACE$(1)

'
END SUB

'*****
'**      Name:      StarDataEntry      **
'**      Type:      Subroutine      **
'**      Module:     MIRACTRL.BAS      **
'**      Language:   Quickbasic 4.50      **
'*****


'Allows entry of star position data as an initial user input
'
'Example of use: StarDataEntry Star$(), RA$(), DEC$(),
Epoch$, NewFile%, w1, w1Text$, w1Title$, w1Prompt$
'Parameters:      (none)
'Variables:       Star$()      A 5-element matrix of desired
stars for
'
                           viewing
'
                           RA$()      A 5-element matrix of right
ascensions
'
                           DEC$()      for the stars listed in Star$()
'
                           Epoch$()    A 5-element matrix of declinations
                           for the stars listed in Star$()
'
                           RA's and DEC's of the stars listed
in Star$()
'
                           g%, h%, i%  Looping indices
'
                           j%          A flag activated when the
Backspace key is hit

```

```

'           1%      A flag activated when the Shift
Tab, Up or Down
'           Arrows are hit
'           row%, col% The current row and column markers
for the
'           cursor
'           Stop$     A string representing whether or
not the
'           character entry loop is to be
exited
'           kee%     A unique code returned by the
KeyCode% or
'           KeyCodeNoWait% Functions for each
key hit
'           buffer
'           Character$  A string read from the keyboard
the current
'           Current$   A string representing the entry of
one of the
'           field. This variable will become
following: Star$(), RA$(),
DEC$(), EPOCH$()
'
'Module Level
'   DECLARATIONS: DECLARE SUB StarDataEntry (Star$(), RA$(),
DEC$(),
'                           Epoch$(), NewFile%, w1,
w1Text$, w1Title$, w1Prompt$)
'
SUB StarDataEntry (Star$(), RA$(), DEC$(), EPOCH$(), NewFile%,
w1 AS WindowsType, w1Text$, w1Title$, w1Prompt$)
'
'Initialize variables
IF NewFile% = TRUE THEN
  FOR h% = 1 TO 100
    Star$(h%) = ""
    RA$(h%) = ""
    DEC$(h%) = ""
    EPOCH$(h%) = ""
  NEXT h%
END IF
current$ = Star$(1)
j% = 0
L% = 0
i% = 1
row% = 17
col% = 9
InsertOn% = FALSE
FirstRow% = 17
LastRow% = 21
LastCol% = 62

```

```

FileLine% = 1
TopFileLineShown% = 1
ChangeCol% = FALSE
ChangeFileLine% = FALSE
HomeOrEnd% = FALSE
Stops% = FALSE
QuitEntry% = FALSE

'Draw bottom edge of the status box
LOCATE 22, 5, 0
PRINT CHR$(200); STRING$(68, 205); CHR$(188);

'Complete data fields by reading in one key at a time, ESC
exits
DO
    IF TopFileLineShown% < 1 THEN TopFileLineShown% = 1
    IF TopFileLineShown% > 96 THEN TopFileLineShown% = 96
    IF FileLine% < 1 THEN FileLine% = 1
    IF FileLine% > 100 THEN FileLine% = 100
    IF row% < FirstRow% THEN row% = FirstRow%
    IF row% > LastRow% THEN row% = LastRow%
    'Update the row numbers and fields
    FOR g% = 1 TO 5
        COLOR WHITE, BLUE, BLUE
        LOCATE 16 + g%, 6, 0
        PRINT TopFileLineShown% + g% - 1
        COLOR WHITE, BLACK, BLUE
        LOCATE 16 + g%, 10, 0
        PRINT Star$(TopFileLineShown% + g% - 1) +
SPACE$(19 - LEN(Star$(TopFileLineShown% + g% - 1)))
        LOCATE 16 + g%, 33, 0
        PRINT RA$(TopFileLineShown% + g% - 1) +
SPACE$(10 - LEN(RA$(TopFileLineShown% + g% - 1)))
        LOCATE 16 + g%, 48, 0
        PRINT DEC$(TopFileLineShown% + g% - 1) +
SPACE$(10 - LEN(DEC$(TopFileLineShown% + g% - 1)))
        LOCATE 16 + g%, 63, 0
        PRINT EPOCH$(TopFileLineShown% + g% - 1) +
SPACE$(10 - LEN(EPOCH$(TopFileLineShown% + g% - 1)))
    NEXT g%
    COLOR WHITE, BLACK, BLUE
    'Read the next character and operate on it
    DO
        LOCATE row%, col% + i%, 1
        IF i% = 0 THEN
            i% = 1
            j% = 0
        END IF
        IF j% = 1 THEN
            PRINT " "
            IF LEN(current$) > 1 THEN

```

```

current$ = LEFT$(current$,
(LEN(current$) - 1))
    ELSE
        current$ = ""
    END IF
    j% = 0
END IF
LOCATE row%, col% + i%, 1
kee% = KeyCode%(Character$)
SELECT CASE kee%
    CASE ENTER
        IF row% = LastRow% AND col% =
LastCol% THEN
            row% = row% + 1
            IF row% = LastRow% +
1 THEN
                TopFileLineShown% = TopFileLineShown% + 1
                FileLine% =
FileLine% + 1
            END IF
        ELSE
            'no action
        END IF
        IF col% = 9 THEN
            i% = 20
        ELSE
            i% = 11
        END IF
        Stops% = TRUE
    CASE TABHIT
        IF col% = 9 THEN
            i% = 20
        ELSE
            i% = 11
        END IF
    CASE ESCAPE
        windows   w1,      w1Text$(),
w1Title$, w1Prompt$
        IF w1.returnCode = 121 OR
w1.returnCode = 89 THEN Abort% = TRUE
        IF Abort% = TRUE THEN
'***** Procedures  **
**                                         *** System Shutdown
                                         *** go here!!
**
'***** WindowsPop

```

```

        CLS
        SYSTEM
    ELSE
        WindowsPop
    END IF
CASE BACKSPACE
    i% = i% - 1
    j% = 1
CASE SHIFTTAB
    IF col% = 32 THEN col% = 9
    IF col% = 47 THEN col% = 32
    IF col% = 62 THEN col% = 47
    Stops% = TRUE
CASE UPARROW
    IF row% = FirstRow% THEN
        TopFileLineShown% =
TopFileLineShown% - 1
    ELSE
        row% = row% - 1
    END IF
    ChangeFileLine% = MINUS
    Stops% = TRUE
CASE DOWNARROW
    IF row% = LastRow% THEN
        TopFileLineShown% =
TopFileLineShown% + 1
    ELSE
        row% = row% + 1
    END IF
    ChangeFileLine% = PLUS
    Stops% = TRUE
CASE PGDOWN
    IF TopFileLineShown% > 95 THEN
        ChangeFileLine% = PLUS
        TopFileLineShown% = 96
    ELSE
        TopFileLineShown% =
        ChangeFileLine% = PLUS
* (96 - TopFileLineShown%)
        TopFileLineShown% + 5
        * 5
        END IF
        Stops% = TRUE
CASE PGUP
    IF TopFileLineShown% < 6 THEN
        ChangeFileLine% =
        TopFileLineShown% = 1
MINUS * (TopFileLineShown% - 1)
        ELSE
        TopFileLineShown% =
TopFileLineShown% - 5

```

```

        ChangeFileLine%      =
MINUS * 5
                END IF
                Stops% = TRUE
CASE HOME
                TopFileLineShown% = 1
                HomeOrEnd% = 1
                Stops% = TRUE
CASE ENDFILE
                TopFileLineShown% = 96
                HomeOrEnd% = 100
                Stops% = TRUE
CASE LEFTARROW
                i% = i% - 1
CASE RIGHTARROW
                IF LEN(current$) > i% THEN
                    i% = i% + 1
                END IF
CASE DELETE
                IF i% > 1 AND i% <
LEN(current$) + 1 THEN
                    current$ =
LEFT$(current$, i% - 1) + RIGHT$(current$, LEN(current$) - i%)
                ELSEIF i% = 1 AND i% <
LEN(current$) + 1 THEN
                    current$ =
RIGHT$(current$, LEN(current$) - i%)
                ELSEIF i% = 1 THEN
                    current$ = ""
                ELSE
                    current$ =
LEFT$(current$, i% - 1)
                END IF
                LOCATE row%, col% + 1, 1
                IF col% = 9 THEN
                    PRINT      current$;
SPACE$(19 - LEN(current$))
                ELSE
                    PRINT      current$;
SPACE$(10 - LEN(current$))
                END IF
CASE INSERT
                IF InsertOn% = TRUE THEN
                    InsertOn% = FALSE
                    LOCATE , , , 7, 7
                ELSE
                    InsertOn% = TRUE
                    LOCATE , , , 3, 7
                END IF
CASE 17408 'F10
                QuitEntry% = TRUE

```

```

        Stops% = TRUE
CASE 97 TO 122
        AcceptChar% = TRUE
CASE 65 TO 90
        AcceptChar% = TRUE
CASE 48 TO 57
        AcceptChar% = TRUE
CASE 58 TO 59
        AcceptChar% = TRUE
CASE 34
        AcceptChar% = TRUE
CASE 39
        AcceptChar% = TRUE
CASE 44
        AcceptChar% = TRUE
CASE 46
        AcceptChar% = TRUE
CASE 32
        AcceptChar% = TRUE
CASE ELSE
*****
**
**          Error Handling
Needed      **
*****
**
*****
END SELECT
IF col% = 9 AND i% = 20 THEN
    Stops% = TRUE
    ChangeCol% = TRUE
ELSEIF col% > 9 AND i% = 11 THEN
    Stops% = TRUE
    ChangeCol% = TRUE
END IF
IF AcceptChar% = TRUE THEN Stops% = TRUE
LOOP UNTIL Stops% = TRUE
Stops% = FALSE
IF AcceptChar% = TRUE THEN
    IF InsertOn% = TRUE THEN
        IF i% > 1 AND i% < LEN(current$) + 1
THEN
            current$ = LEFT$(current$, i%
- 1) + Character$ + RIGHT$(current$, LEN(current$) - i% + 1)
        ELSEIF i% = 1 AND i% < LEN(current$)
+ 1 THEN
            current$ = Character$ +
RIGHT$(current$, LEN(current$) - i% + 1)

```

```

        ELSEIF i% = 1 THEN
            current$ = Character$
        ELSE
            current$ = LEFT$(current$, i%
- 1) + Character$
        END IF
    ELSE
        IF i% > 1 AND i% < LEN(current$) THEN
            current$ = LEFT$(current$, i%
- 1) + Character$ + RIGHT$(current$, LEN(current$) - i%)
        ELSEIF i% = 1 AND i% < LEN(current$)
THEN
            current$ = Character$ +
RIGHT$(current$, LEN(current$) - i%)
        ELSEIF i% = 1 THEN
            current$ = Character$
        ELSE
            current$ = LEFT$(current$, i%
- 1) + Character$
        END IF
    END IF
    LOCATE row%, col% + 1, 1
    IF col% = 9 THEN
        PRINT      current$;      SPACE$(19) -
LEN(current$))
    ELSE
        PRINT      current$;      SPACE$(10) -
LEN(current$))
    END IF
    i% = i% + 1
    AcceptChar% = FALSE
END IF
IF FileLine% < 1 THEN FileLine% = 1
IF FileLine% > 100 THEN FileLine% = 100
SELECT CASE col%
    CASE 9
        Star$(FileLine%) = current$
    CASE 32
        RA$(FileLine%) = current$
    CASE 47
        DEC$(FileLine%) = current$
    CASE 62
        EPOCH$(FileLine%) = current$
    CASE ELSE
        *****
        ***
        ***
        Error Handling
Needed      **

```

```

    ***
    **

***** END SELECT

IF ChangeFileLine% <> FALSE THEN
    FileLine% = FileLine% + ChangeFileLine%
    IF FileLine% < 1 THEN FileLine% = 1
    IF FileLine% > 100 THEN FileLine% = 100
    ChangeFileLine% = FALSE
    SELECT CASE col%
        CASE 9
            current$ = Star$(FileLine%)
        CASE 32
            current$ = RA$(FileLine%)
        CASE 47
            current$ = DEC$(FileLine%)
        CASE 62
            current$ = EPOCH$(FileLine%)
    CASE ELSE

***** Error Handling
Needed      **
    ***
    **

***** END SELECT
END IF
'
IF HomeOrEnd% <> FALSE THEN
    IF HomeOrEnd% = 1 THEN
        row% = FirstRow%
    ELSEIF HomeOrEnd% = 100 THEN
        row% = LastRow%
    END IF
    FileLine% = HomeOrEnd%
    HomeOrEnd% = FALSE
    COLOR WHITE, BLUE, BLUE
    LOCATE LastRow%, 9, 0
    PRINT SPACE$(1)
    SELECT CASE col%
        CASE 9
            current$ = Star$(FileLine%)
        CASE 32
            current$ = RA$(FileLine%)
        CASE 47

```

```

                current$ = DEC$(FileLine%)
CASE 62          current$ = EPOCH$(FileLine%)
CASE ELSE

' *****
' **
**
' ***          Error Handling
Needed      **
' ***
**
' *****

END SELECT
END IF
'
IF ChangeCol% = TRUE THEN
    SELECT CASE col%
        CASE 9
            col% = 32
            current$ = RA$(FileLine%)
        CASE 32
            col% = 47
            current$ = DEC$(FileLine%)
        CASE 47
            col% = 62
            current$ = EPOCH$(FileLine%)
        CASE 62
            col% = 9
            row% = row% + 1
            FileLine% = FileLine% + 1
            IF FileLine% < 1 THEN FileLine% = 1
            IF FileLine% > 100 THEN FileLine% =
100
            current$ = Star$(FileLine%)
        CASE ELSE
' *****
' **
**
' ***          Error Handling
Needed      **
' ***
**
' *****

END SELECT
i% = 1
ChangeCol% = FALSE
END IF

```

```

LOOP UNTIL QuitEntry% = TRUE
'
END SUB

'*****
'*      Name:      StarDisplay      **
'*      Type:      Subroutine      **
'*      Module:    MIRACTRL.BAS   **
'*      Language:  Quickbasic 4.50  **
'*****

'
'Creates the display screen graphics for star data entry
'
'Example of use:  StarDisplay
'Parameters:      (none)
'Variables:       i%           Looping index
'Module Level
'    DECLARATIONS:  DECLARE SUB StarDisplay
'

SUB StarDisplay
'
'Clear field
COLOR WHITE, BLUE, BLUE
FOR r% = 14 TO 21
    LOCATE r%, 6, 0
    PRINT SPACE$(68)
NEXT r%
'
'Fill in star data entry fields
COLOR WHITE, BLUE, BLUE
LOCATE 14, 31, 0
PRINT "Please enter stars"
LOCATE 15, 33, 0
PRINT "RA          DEC";
LOCATE 16, 10, 0
PRINT "Star Name           hr:min:sec";
LOCATE 16, 48, 0
PRINT CHR$(248); ":";;
LOCATE 16, 63, 0
PRINT "Epoch";
'
'Draw fields for star data input
FOR i% = 1 TO 5
    COLOR WHITE, BLACK, BLUE
    LOCATE 16 + i%, 10, 0
    PRINT SPACE$(19)
    LOCATE 16 + i%, 33, 0
    PRINT SPACE$(10)
    LOCATE 16 + i%, 48, 0
    PRINT SPACE$(10)
    LOCATE 16 + i%, 63, 0

```

```

        PRINT SPACE$(10)
NEXT i%
LOCATE 23, 5, 0
COLOR WHITE, BLUE, BLUE
PRINT "F10 - SAVE and exit data entry"
'
END SUB

'*****
'*      Name:      TeleAlign      **
'*      Type:      Subroutine      **
'*      Module:    MIRACTRL.BAS   **
'*      Language:  Quickbasic 4.50  **
'*****

'Allows user input of the telescope's Hour Angle and
Declination to align
'the telescope in the local coordinate system.
'
'Example of use: TeleAlign GMST0hrSEC#, GHAaries#, TeleRA#,
TeleDEC#
SUB TeleAlign (GMST0hrSEC#, GHAaries#, TeleRA#, TeleDEC#)
'
'Initialize variables
MIRALatitude# = .63908139#           'radians, (36 deg 37 min
North)
MIRAlongitude# = 2.12639281#         'radians, (121 deg 50 min
West)
sec% = VAL(RIGHT$(TIME$, 2))
min% = VAL(MID$(TIME$, 4, 2))
hour% = VAL(LEFT$(TIME$, 2))

'Calculate Greenwich Mean Sidereal Time
'Determine the day fraction that has passed since midnight
'in SECONDS
sec& = sec%
hours& = hour%
mins& = min%
dayFraction# = CDBL((hours& * 3600) + (mins& * 60) + sec&)

'Determine the Greenwich Mean Sidereal Time at the current
'time in SECONDS
GMSTinSEC# = GMST0hrSEC# + dayFraction#
'

'Calculate Greenwich Hour Angle of Aries
GHAaries# = GMSTinSEC# * 2# * pi / 86164.09053099999# 'Seconds
of solar time converted to radians
DO WHILE GHAaries# > (2# * pi)
    GHAaries# = GHAaries# - (2# * pi)
LOOP
DO WHILE GHAaries# < 0#

```

```

        GHAaries# = GHAaries# + (2# * pi)
LOOP
'
'Calculate the telescope's hour angle for this moment
RAofMIRA# = GHAaries# - MIRAlongitude#
'
'***** *****
'Read telescope's Hour Angle from the encoders **
'Store to variable HourAngle# in radians      **
'***** *****
TeleRA# = RAofMIRA# - HourAngle#    'in radians
DO WHILE TeleRA# > (2# * pi)
    TeleRA# = TeleRA# - (2# * pi)
LOOP
DO WHILE TeleRA# < 0#
    TeleRA# = TeleRA# + (2# * pi)
LOOP
'
'***** *****
'Read telescope's Declination from the encoders **
'Store to variable TeleDEC# in radians      **
'***** *****
'
END SUB

'***** *****
'**      Name:      TeleAngles      **
'**      Type:      Subroutine      **
'**      Module:    MIRACTRL.BAS    **
'**      Language:   Quickbasic 4.50  **
'***** *****

'
'Calculates the telescope's zenith and azimuth angles in the
'terrestrial coordinate frame.
'
'Example of use:  TeleAngles Zenith%, Azimuth%, GHAaries#,
RAofStar#, DECoStar#, CelCoord#(), DCM#(), TerraRec#()
'
SUB TeleAngles (Zenith%, Azimuth%, GHAaries#, RAofStar#,
DECoStar#, CelCoord#(), DCM#(), TerraRec#())
'
'Initialize variables
MIRALatitude# = .63908139#           'radians, (36 deg 37 min
North)
MIRAlongitude# = 2.12639281#         'radians, (121 deg 50 min
West)
'
'Define a celestial coordinate vector in RECTANGULAR
coordinates
CelCoord#(1) = COS(RAofStar#) * COS(DECoStar#)
CelCoord#(2) = SIN(RAofStar#) * COS(DECoStar#)

```

```

CelCoord#(3) = SIN(DECofStar#)
'
'Define the direction cosine matrix which relates celestial
and terrestrial coordinates
DCM#(1, 1) = (-SIN(GHAaries#) * SIN(MIRALongitude#) *
SIN(MIRALatitude#) - COS(GHAaries#) * COS(MIRALongitude#) *
SIN(MIRALatitude#))
DCM#(1, 2) = (-SIN(GHAaries#) * COS(MIRALongitude#) *
SIN(MIRALatitude#) + COS(GHAaries#) * SIN(MIRALongitude#) *
SIN(MIRALatitude#))
DCM#(1, 3) = COS(MIRALatitude#)
DCM#(2, 1) = (-COS(GHAaries#) * SIN(MIRALongitude#) +
SIN(GHAaries#) * COS(MIRALongitude#))
DCM#(2, 2) = (-SIN(GHAaries#) * SIN(MIRALongitude#) -
COS(GHAaries#) * COS(MIRALongitude#))
DCM#(2, 3) = 0
DCM#(3, 1) = (COS(GHAaries#) * COS(MIRALongitude#) *
COS(MIRALatitude#) + SIN(GHAaries#) * SIN(MIRALongitude#) *
COS(MIRALatitude#))
DCM#(3, 2) = (SIN(GHAaries#) * COS(MIRALongitude#) *
COS(MIRALatitude#) - COS(GHAaries#) * SIN(MIRALongitude#) *
COS(MIRALatitude#))
DCM#(3, 3) = SIN(MIRALatitude#)
'
'Calculate the terrestrial rectangular coordinates
TerraRec#(1) = DCM#(1, 1) * CelCoord#(1) + DCM#(1, 2) *
CelCoord#(2) + DCM#(1, 3) * CelCoord#(3)
TerraRec#(2) = DCM#(2, 1) * CelCoord#(1) + DCM#(2, 2) *
CelCoord#(2) + DCM#(2, 3) * CelCoord#(3)
TerraRec#(3) = DCM#(3, 1) * CelCoord#(1) + DCM#(3, 2) *
CelCoord#(2) + DCM#(3, 3) * CelCoord#(3)
'
'Calculate the azimuth and zenith angles
IF TerraRec#(1) < 0# AND TerraRec#(2) < 0# THEN
    Azimuth# = ATN(TerraRec#(1) / TerraRec#(2)) + (pi /
2#)
ELSEIF TerraRec#(1) >= 0# AND TerraRec#(2) < 0# THEN
    Azimuth# = ATN(-TerraRec#(2) / TerraRec#(1))
ELSEIF TerraRec#(1) < 0# AND TerraRec#(2) >= 0# THEN
    Azimuth# = ATN(-TerraRec#(2) / TerraRec#(1)) + pi
ELSE Azimuth# = ATN(TerraRec#(1) / TerraRec#(2)) + ((3# * pi) /
2#)
END IF
Azimuth% = INT(Azimuth# * 360# / (2# * pi))
DO WHILE Azimuth% < 0
    Azimuth% = Azimuth% + 360
LOOP
DO WHILE Azimuth% > 360
    Azimuth% = Azimuth% - 360
LOOP

```



```

COLOR WHITE, BLUE, BLUE
LOCATE 10, 38, 0
PRINT "Star"
'
'Draw the body of the star name box and place the name
COLOR WHITE, BLUE, BLUE
LOCATE 11, 29, 0
PRINT CHR$(186); STRING$(19, 0); CHR$(186);
LOCATE 11, INT((80 - LEN(ActiveStar$)) / 2), 0
PRINT ActiveStar$
'
'Draw bottom edge of the star name box
COLOR WHITE, BLUE, BLUE
LOCATE 12, 29, 0
PRINT CHR$(200); STRING$(19, 205); CHR$(188);
'
'Place the lettering
COLOR WHITE, RED, BLUE
LOCATE 2, 27, 0
PRINT " ESC key halts telescope "
COLOR WHITE, BLUE, BLUE
LOCATE 6, 28, 0
PRINT "Oliver Observing Station"
LOCATE 7, 32, 0
PRINT "Telescope Status"
'
'Draw top edge of status box
LOCATE 13, 5, 0
COLOR WHITE, BLUE, BLUE
PRINT CHR$(201); STRING$(68, 205); CHR$(187);
'
'Draw the body of the status box
FOR r% = 14 TO 21
    LOCATE r%, 5, 0
    PRINT CHR$(186);
    PRINT SPACE$(68);
    PRINT CHR$(186);
NEXT r%
'
'Draw bottom edge of the status box
LOCATE 22, 5, 0
PRINT CHR$(200); STRING$(68, 205); CHR$(188);
LOCATE 22, 23, 0
PRINT " ENTER to change telescope status "
'
'Place the help instructions
LOCATE 23, 31, 0
PRINT "F1 for help screen"
'Place the lettering
COLOR WHITE, BLUE, BLUE
LOCATE 14, 10, 0

```

```

PRINT "RA"
LOCATE 17, 10, 0
PRINT "DEC"
LOCATE 20, 10, 0
PRINT "Hour Angle"
LOCATE 14, 34, 0
PRINT "Zenith Angle"
LOCATE 17, 36, 0
PRINT "Azimuth"
LOCATE 14, 58, 0
PRINT "PST"
LOCATE 17, 58, 0
PRINT "Sidereal Time"
LOCATE 20, 58, 0
PRINT "UTC"
'
'Show motor state boxes
COLOR RED, BLUE, BLUE
LOCATE 1, 35, 0
PRINT "Simulation"
COLOR WHITE, BLUE, BLUE
LOCATE 5, 7, 0
PRINT "RA Slew"
LOCATE 5, 17, 0
PRINT "RA Step"
LOCATE 5, 57, 0
PRINT "DEC Slew"
LOCATE 5, 66, 0
PRINT "DEC Step"
FOR i% = 1 TO 2
    LOCATE 2, 10 * i% - 1, 0
    PRINT CHR$(201); CHR$(205); CHR$(187)
    LOCATE 3, 10 * i% - 1, 0
    PRINT CHR$(186); SPACE$(1); CHR$(186)
    LOCATE 4, 10 * i% - 1, 0
    PRINT CHR$(200); CHR$(205); CHR$(188)
NEXT i%
FOR i% = 1 TO 2
    LOCATE 2, (10 * i% - 1) + 50, 0
    PRINT CHR$(201); CHR$(205); CHR$(187)
    LOCATE 3, (10 * i% - 1) + 50, 0
    PRINT CHR$(186); SPACE$(1); CHR$(186)
    LOCATE 4, (10 * i% - 1) + 50, 0
    PRINT CHR$(200); CHR$(205); CHR$(188)
NEXT i%
'
END SUB

*****
**      Name:      TrackCommands      **
**      Type:      Subroutine      **

```

```

'**      Module:      MIRACTRL.BAS      **
'**      Language:    Quickbasic 4.50   **
'***** ****
'
'Controls the right ascension stepping motor only to
compensate
'for local coordinate drift due to the rotation of the Earth.
'
'Example of use: TrackCommands
'
SUB TrackCommands
'
'Track as appropriate in Right Ascension only
'Update motor display box
COLOR RED, BLUE, BLUE
LOCATE 3, 20, 0
PRINT CHR$(222)
'POKE (address of RA track controller), X
END SUB

'*****
'**      Name:      VideoState      **
'**      Type:     Subprogram      **
'**      Module:    BIOSCALL.BAS   **
'**      Language:  Quickbasic 4.50 **
'***** ****
'
'Determines the current video mode parameters.
'
'Example of use: VideoState mode%, columns%, page%
'Parameters:      mode%      Current video mode
'                  columns%    Current number of text columns
'                  page%      Current active display page
'Variables:       reg        Structure of type RegType
'MODULE LEVEL
'DECLARATIONS:   TYPE RegType
'                  ax        AS INTEGER
'                  bx        AS INTEGER
'                  cx        AS INTEGER
'                  dx        AS INTEGER
'                  Bp        AS INTEGER
'                  si        AS INTEGER
'                  di        AS INTEGER
'                  flags    AS INTEGER
'END TYPE
'
DECLARE SUB Interrupt (intnum%, inreg AS RegType, outreg
AS RegType)
'    DECLARE SUB VideoState (mode%, columns%, page%)
'
SUB VideoState (mode%, columns%, page%) STATIC

```

```

DIM reg AS RegType
reg.ax = &HF00
Interrupt &H10, reg, reg
mode% = reg.ax AND &HFF
columns% = (CLNG(reg.ax) AND &HFF00) \ 256
page% = (CLNG(reg.bx) AND &HFF00) \ 256
END SUB

*****
'**      Name:          Windows      **
'**      Type:          Subprogram    **
'**      Module:        WINDOWS.BAS   **
'**      Language:      Quickbasic 4.50 **
*****


'Displays a rectangular window for information display
'or menu selection.

'Example of use: Windows w1, wText$(), wTitle$, wPrompt$

'Parameters:      w1           Structure of type WindowsType
'                  wText$()     Array of strings to be displayed
'                  wTitle$      Title string
'                  wPrompt$     Prompt string

'Variables:       mode%         Current video mode
'                  columns%     Current number of character
columns
'                  page%         Current video page
'                  cursorRow%   Saved cursor row position
'                  cursorCol%   Saved cursor column position
'                  newpage%     Next video page
'                  lbText%      Lower boundary of array of text
lines
'                  ubText%      Upper boundary of array of text
lines
'                  i%           Looping index
'                  maxlen%     Length of longest string to
display
'                  length%     Length of each array string
'                  row2%       Row number at bottom right
corner of window
'                  col2%       Column number at bottom right
corner of
'                  window      window
character code
'                  ul%         Upper left corner border
character code
'                  ur%         Upper right corner border
character code
'                  ll%         Lower left corner border

```

```

'           lr%          Lower right corner border
character code      v1%          Vertical border character code
'           hl%          Horizontal border character
code               r%           Index to each line of text
'           choice$       Set of unique characters for
each menu line     tmp$          Work string
'           kee%          Key code returned by InKeyCode%
function
' MODULE LEVEL
' DECLARATIONS:   SUB Windows (w AS WindowsType, wText$(),
wTitle$,           wPrompt$) STATIC
'
SUB windows (w AS WindowsType, wText$(), wTitle$, wPrompt$)
STATIC
' Key code numbers
CONST DOWNARROW = 20480
CONST ENTER = 13
CONST ESCAPE = 27
CONST UPARROW = 18432
'
' Determine current video page
VideoState mode%, columns%, page%
'
' Record current cursor location
cursorRow% = CSRLIN
cursorCol% = POS(0)
'
' Window will be on the next page, if available
newpage% = page% + 1
IF newpage% > 7 THEN
  SCREEN , , 0, 0
  PRINT "Error: Windows - not enough video pages"
  SYSTEM
END IF
'
' Copy current page to new page
PCOPY page%, newpage%
'
' Show the current page while building window on new page
SCREEN , , newpage%, page%
'
' Determine array bounds
lbText% = LBOUND(wText$)
ubText% = UBOUND(wText$)
'
' Check the text array bounds, lower always 1, upper > 0
IF lbText% <> 1 OR ubText% < 1 THEN

```

```

        SCREEN , , 0, 0
        PRINT "Error: Windows - text array dimensioned
incorrectly"
        SYSTEM
    END IF
'
' Determine longest string in text array
maxLen% = 0
FOR i% = lbText% TO ubText%
    length% = LEN(wText$(i%))
    IF length% > maxLen% THEN
        maxLen% = length%
    END IF
NEXT i%
'
' Determine the bottom right corner of window
row2% = w.row + ubText% + 1
col2% = w.col + maxLen% + 3
'
' Check that window fits on screen
IF w.row < 1 OR w.col < 1 OR row2% > 25 OR col2% > columns%
THEN
    SCREEN , , 0, 0
    PRINT "Error: Windows - part of window is off screen"
    PRINT "columns% = "; columns%
    SYSTEM
END IF
'
' Set the edge characters
SELECT CASE w.edgeline
CASE 0
    ul% = 32
    ur% = 32
    ll% = 32
    lr% = 32
    vl% = 32
    hl% = 32
CASE 1
    ul% = 218
    ur% = 191
    ll% = 192
    lr% = 217
    vl% = 179
    hl% = 196
CASE 2
    ul% = 201
    ur% = 187
    ll% = 200
    lr% = 188
    vl% = 186
    hl% = 205

```

```

CASE ELSE
  SCREEN , , 0, 0
  PRINT "Error: Windows - Edge line types incorrect"
  SYSTEM
END SELECT

'Draw top edge of box
LOCATE w.row, w.col, 0
COLOR w.fgdEdge, w.bgdEdge
PRINT CHR$(ul%); STRING$(maxLen% + 2, hl%); CHR$(ur%);

'Draw the body of the window
FOR r% = w.row + 1 TO row2% - 1
  LOCATE r%, w.col, 0
  COLOR w.fgdEdge, w.bgdEdge
  PRINT CHR$(vl%);
  COLOR w.fgdBody, w.bgdBody
  tmp$ = LEFT$(wText$(r% - w.row) + SPACE$(maxLen%), maxLen%)
  PRINT " "; tmp$; " ";
  COLOR w.fgdEdge, w.bgdEdge
  PRINT CHR$(vl%);
NEXT r%

'Draw bottom edge of the box
LOCATE row2%, w.col, 0
COLOR w.fgdEdge, w.bgdEdge
PRINT CHR$(ll%); STRING$(maxLen% + 2, hl%); CHR$(lr%);

'Center and print top title, if present
IF wTitle$ <> "" THEN
  LOCATE w.row, (w.col + col2% - LEN(wTitle$) + 1) \ 2, 0
  COLOR w.fgdTitle, w.bgdTitle
  PRINT wTitle$;
END IF

'Center and print prompt, if present
IF wPrompt$ <> "" THEN
  LOCATE row2%, (w.col + col2% - LEN(wPrompt$) + 1) \ 2, 0
  COLOR w.fgdPrompt, w.bgdPrompt
  PRINT wPrompt$;
END IF

'Now make the new page visible and active
SCREEN , , newpage%, newpage%

'Take next action based on action code
SELECT CASE w.action
CASE 1
  'Get a key code number and return it
  DO

```

```

        w.returnCode = KeyCode$(Character$)
    LOOP UNTIL w.returnCode

CASE ELSE
    w.returnCode = 0
END SELECT

'Reset the cursor position
LOCATE cursorRow%, cursorCol%

END SUB

'*****
'**      Name:      WindowsPop      **
'**      Type:      Subprogram      **
'**      Module:     WINDOWS.BAS    **
'**      Language:   Quickbasic 4.50 **
'*****'

'Removes last displayed window

'Example of use:WindowsPop
'Parameters:      (none)
'Variables:       mode%      Current video mode
'                 columns%    Current number of display columns
'                 page%      Current display page

'Module level
'  DECLARATIONS:  DECLARE SUB WindowsPop

SUB WindowsPop STATIC
    VideoState mode%, columns%, page%
    IF page% THEN
        SCREEN 0, , page% - 1, page% - 1
    END IF
END SUB

```

B. SAMPLE DATA FILE "STARFILE.DAT"

LIST OF REFERENCES

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